

自由曲面の製造

Manufacturing of free-form optics

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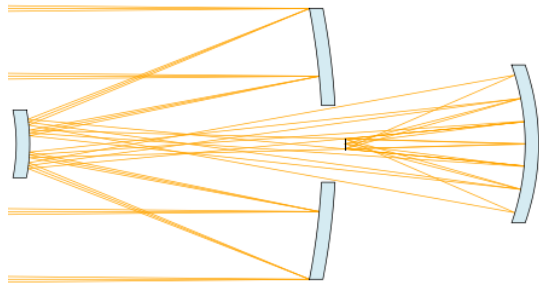
Kyoto Univ.

可視赤外線観測装置ワークショップ2022

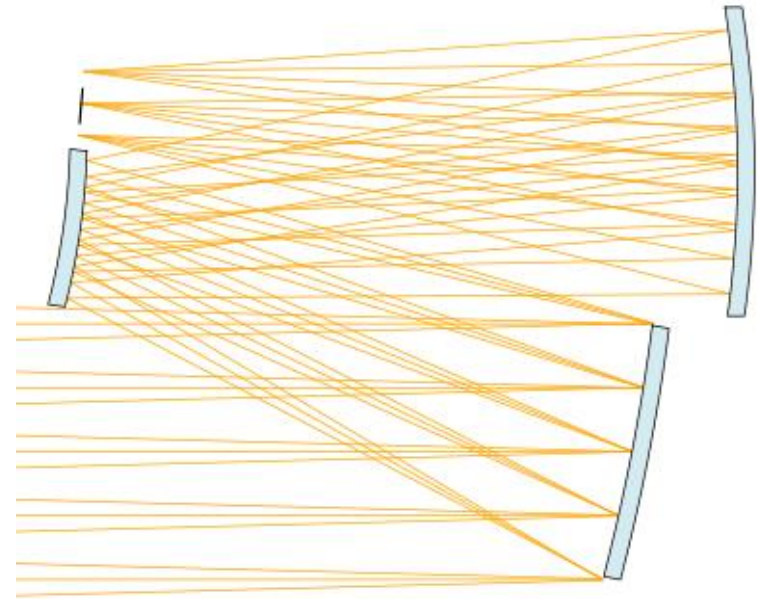
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20221223

Introduction: Potential of Free-form Optics



Conventional Optical Design
Korsch FoV $\sim \Phi 1$ deg

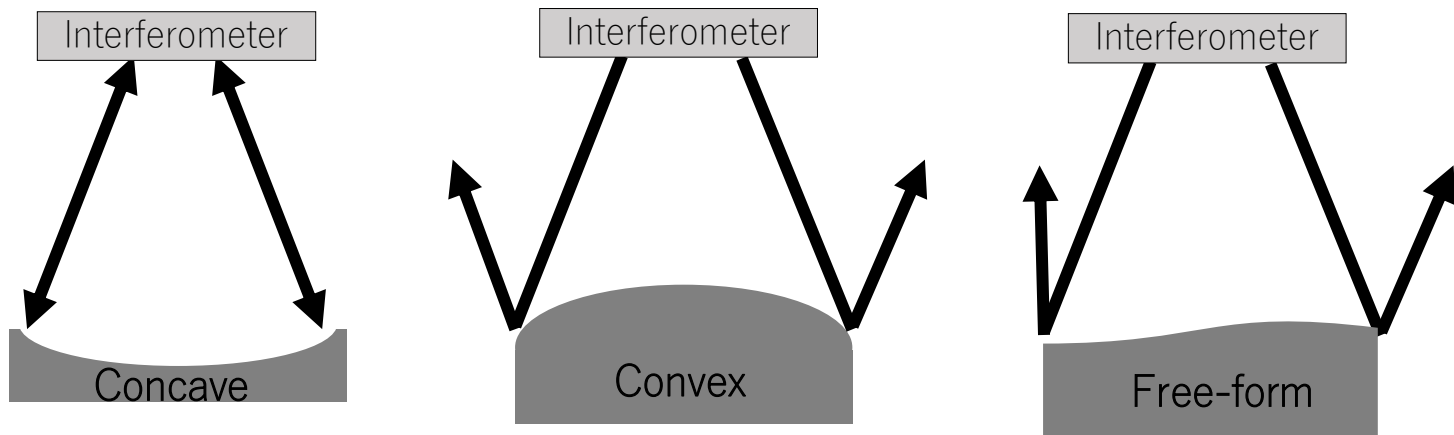


Three Mirror Anastigmat
TMA (designed by us) FoV $\sim \Phi 4$ deg

Introduction

Limitation of Interferometer

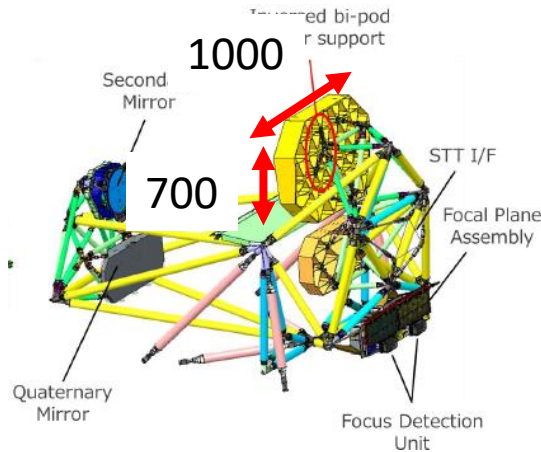
- Interferometer is a powerful tool for a test of optics.
 - requiring an accurate and custom-made reference surface (null optics)
 - requiring a stable and large space equivalent to the radius of curvature (RoC) of the test mirror
 - **challenging to test a large flat, and convex surface**
nearly impossible to test free-form optics highly aspheric



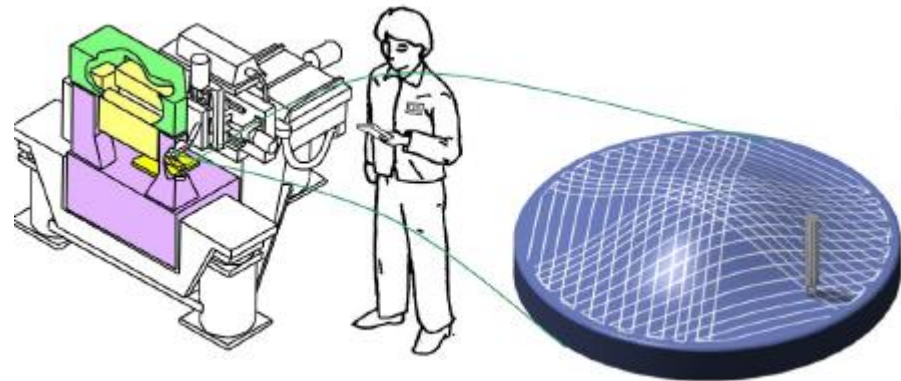
Introduction

Limitation of CMM/Mechanical method

- Coordinate measurement machine is a powerful tool for testing free-form optics.
 - limited to size up to 1 m
 - very few tools which can both test and polish

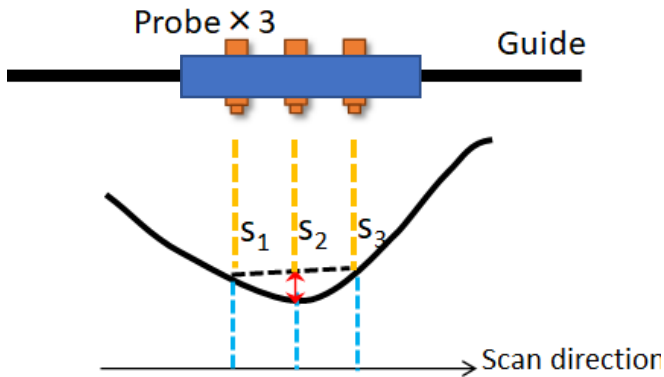


Off-axis four mirror optics
ALOS-3 (JAXA), Watarai+2018



"A-Ruler" developed by Canon Inc
Watarai+2018

Three Point Method (TPM)



- Three-sensor unit travels with linear guide
- Sag or hump (local curvature) is obtained with $2s_2 - s_1 - s_3$
- Local curvatures are doubly integrated along the scanning path
- The cross-sectional shape of the test surface is obtained

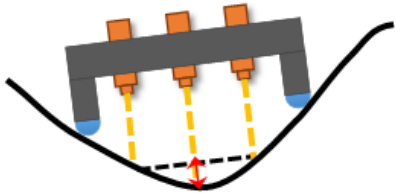


- Measuring local curvature is insensitive to motion error of the unit
- **Precise reference (null optics) is not required → Free-form**



- Low spatial resolution
- Sensor noise is also magnified with the integration process.
- **Inapplicable to surface departing from flat**

Dragging TPM



- Dragging TPM drags a probe unit on the test surface.
- Difference of the local curvature from the design value is integrated along the path for obtaining the figure error.

Dragging TPM



A linear guide is not required.

A large measurement range of sensors is not required.

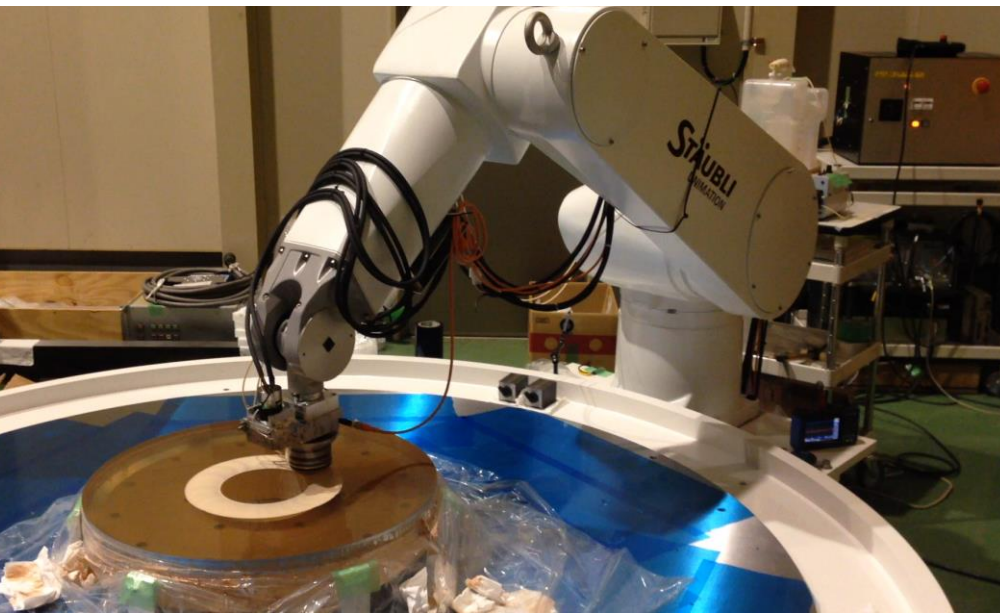
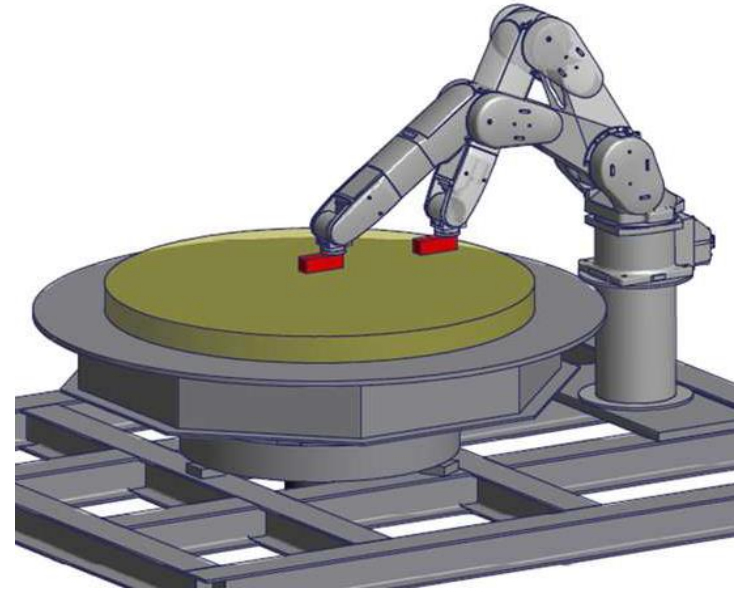
Resistant to temperature changes and vibration

It can be mounted on a polishing machine.

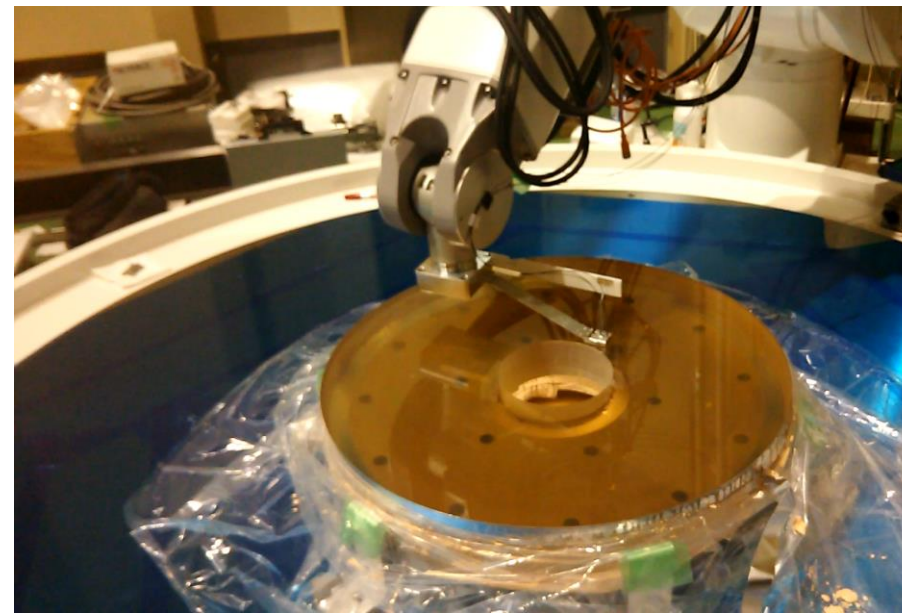
→ Measurable range is not limited

Robot Arm System

- Polish & Profilometer
- Work Area: $\phi 1.8\text{m}$
- Small space
- Robust
- Cost saving



Polish



Measurement

Data Stitching Algorithm

- Sensor error is emphasized with the integral procedure.
- The cross-sectional data is inconsistent with each other at the intersection after least squares.
- New algorithm deals with the data as a deformable *linear elastic body* and stitches the data without inconsistency.
- The solver is based on FEM (minimum energy).

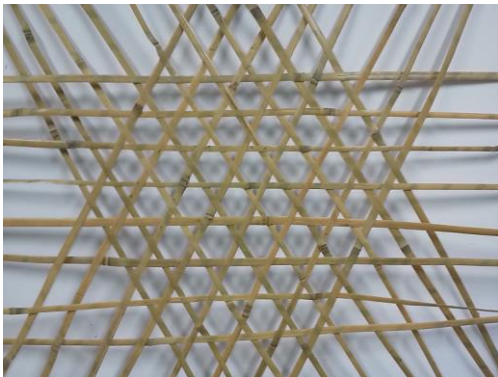
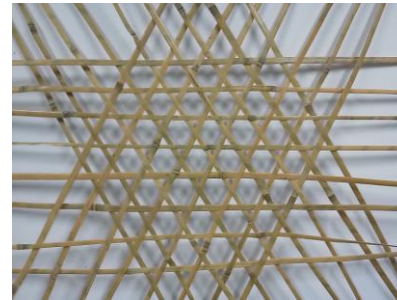
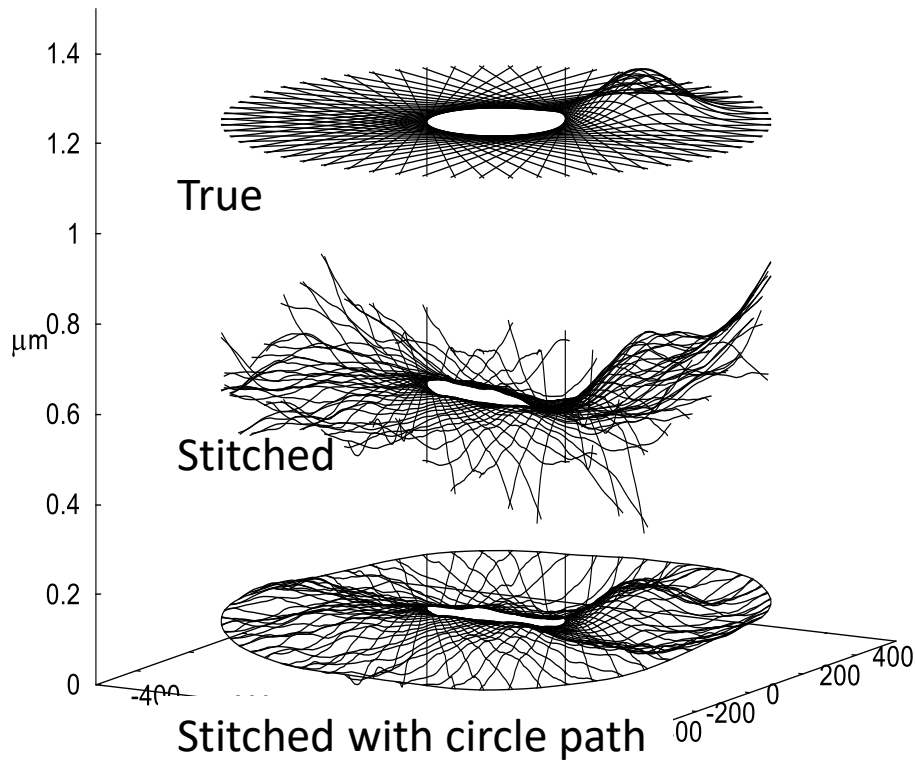


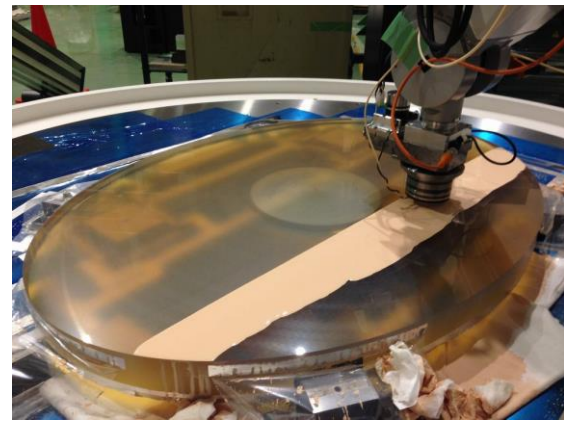
Image of stitching elastic bodies

Simulation

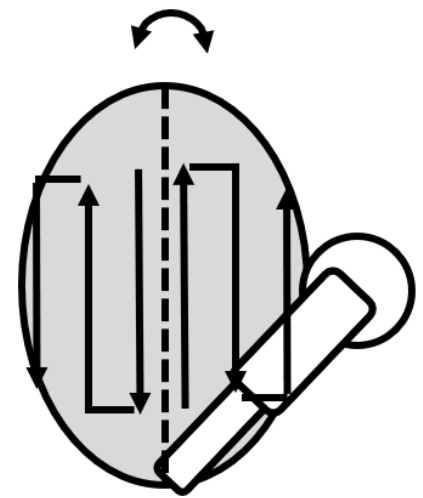
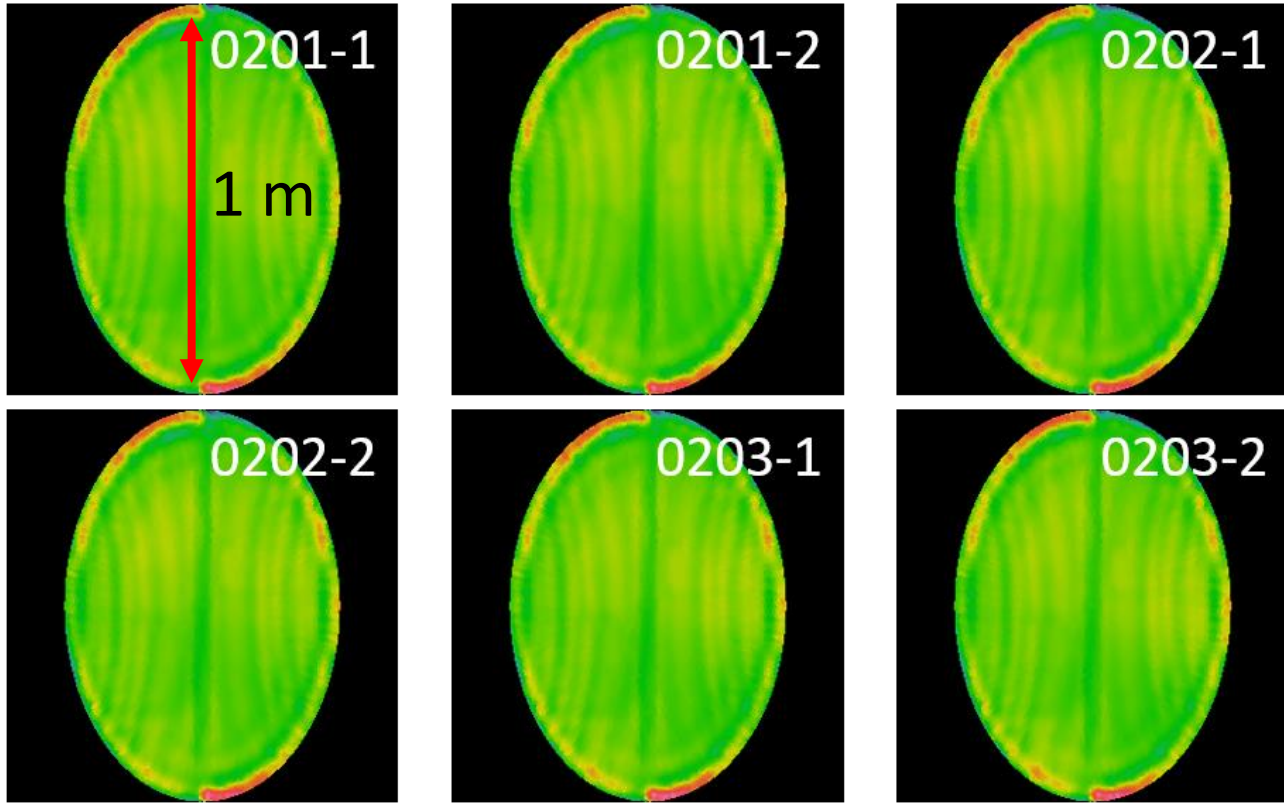


Raw data is made using a three-point method with a sensor noise of 1 nm and 100 steps. Raw data has inconsistencies at the intersections and has a discontinuous surface. Stitched surface reduces figure error and has a **continuous** surface. The circle path at the edge dramatically improves the figure error.

Repeatability on flat mirror



Polishing flat mirror



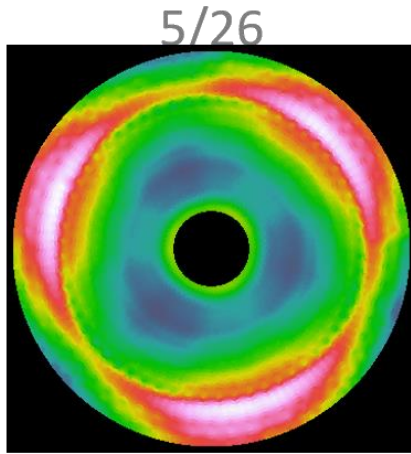
Motion of robot-arm and table

-0.3 μm

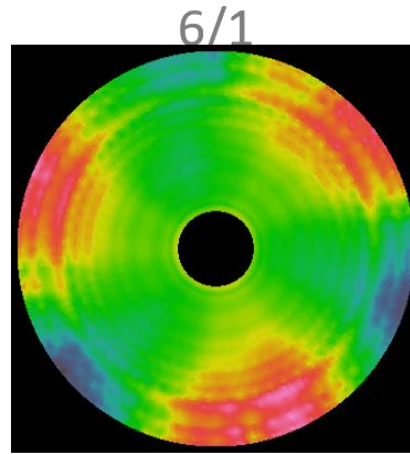
0.3 μm

6 times measurements
RMS < 10 nm

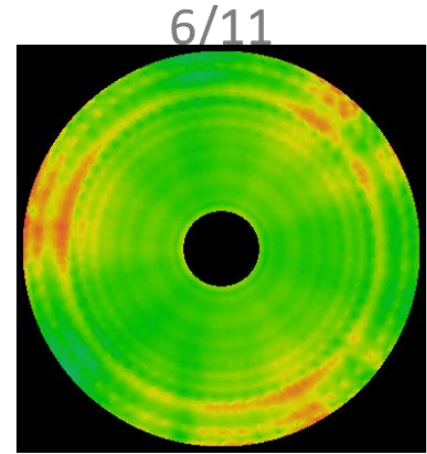
Convex Aspheric



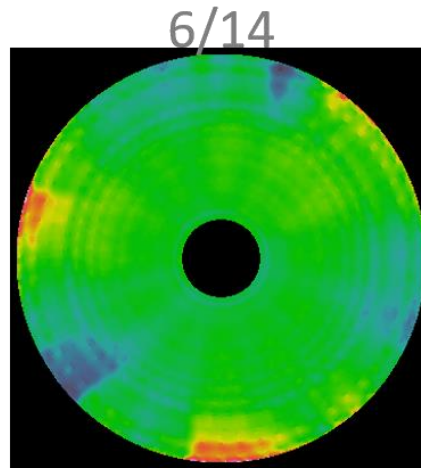
P-V = 2,400 nm
RMS = 550 nm
After grinding



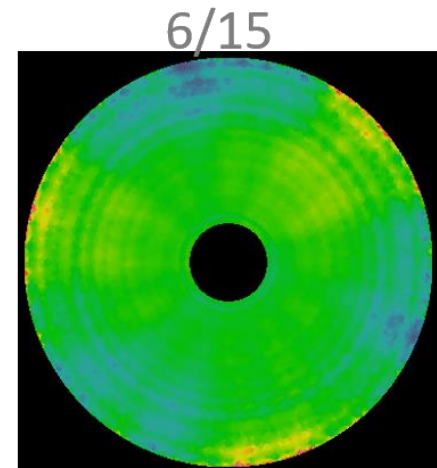
RMS = 350 nm



RMS = 150 nm



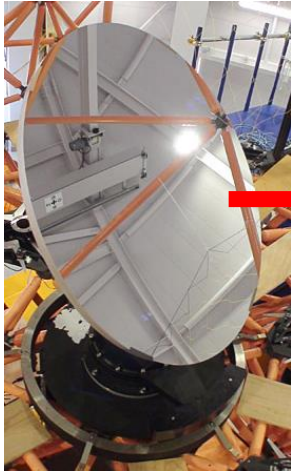
RMS = 130 nm



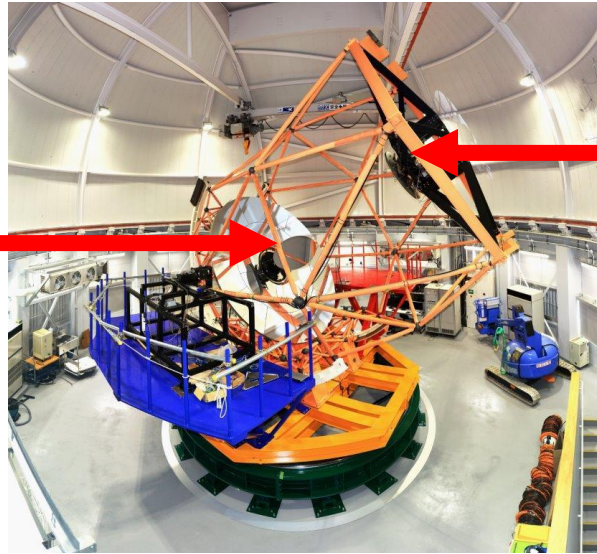
RMS = 35 nm

1 m

Seimei Telescope

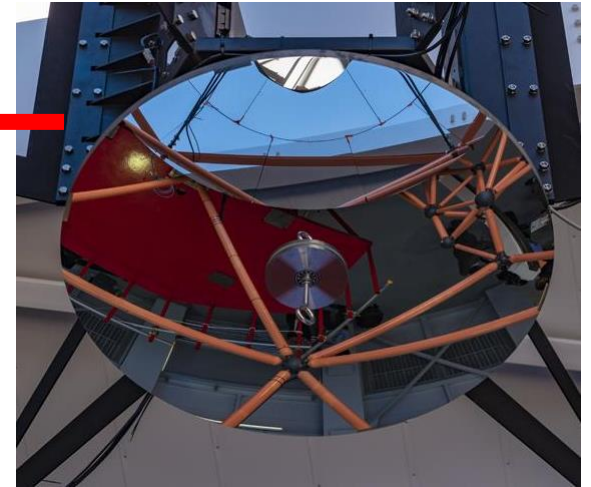


Flat mirror

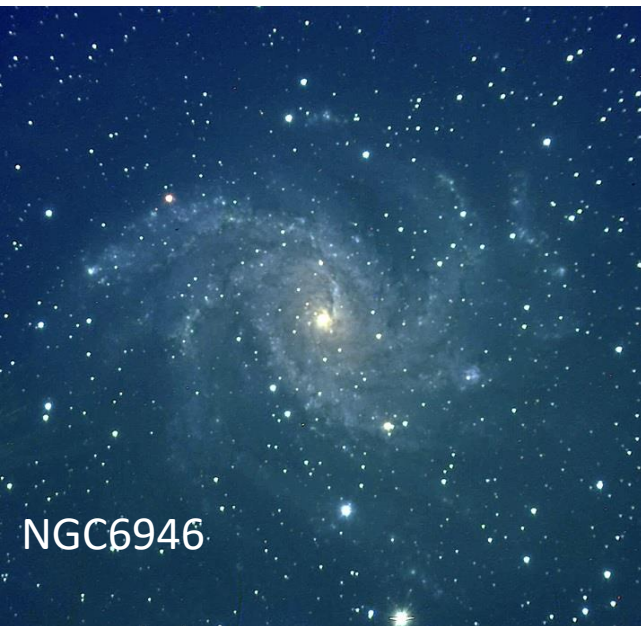


Seimei 3.8 m telescope

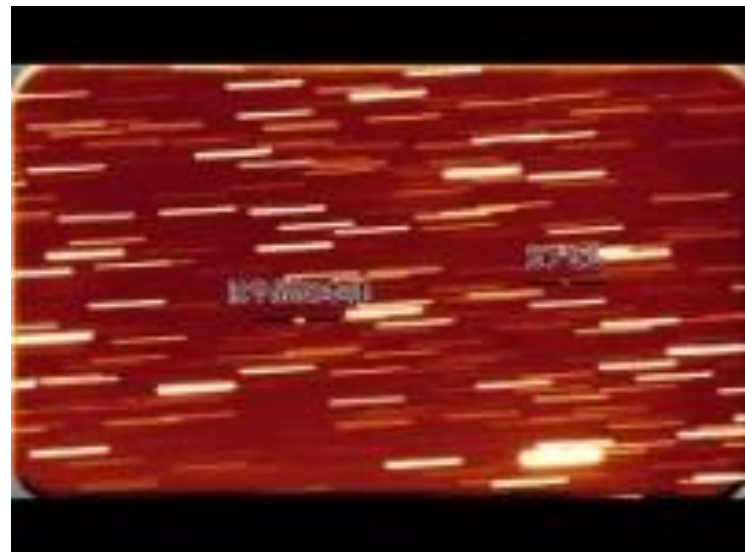
The largest telescope in East-Asia



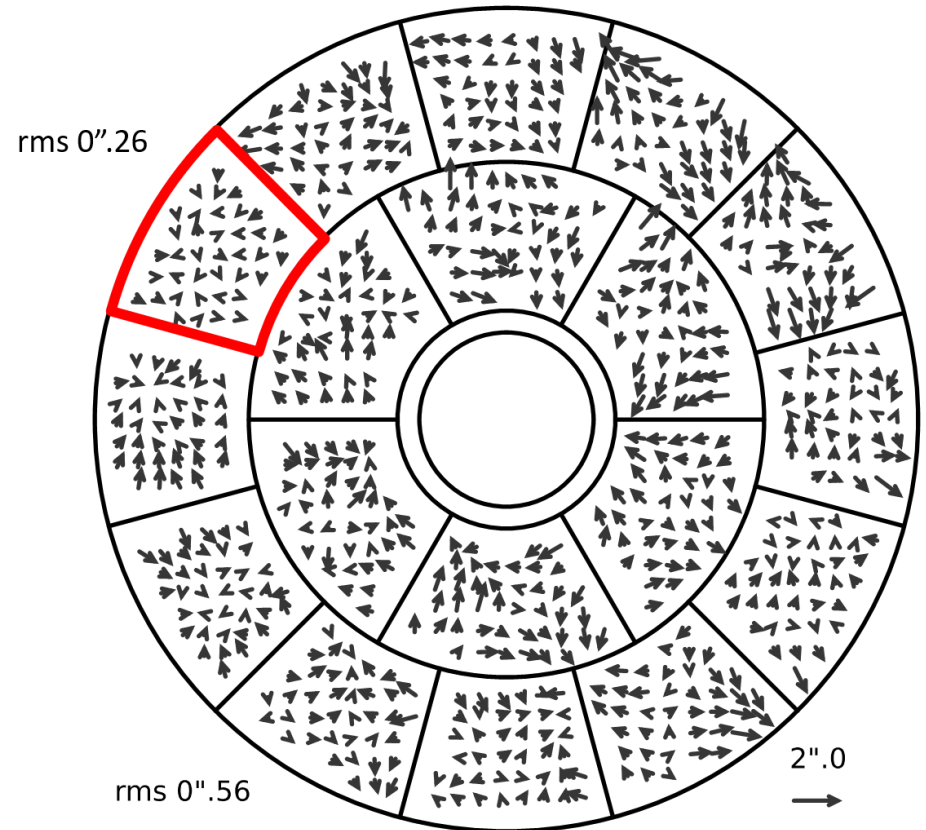
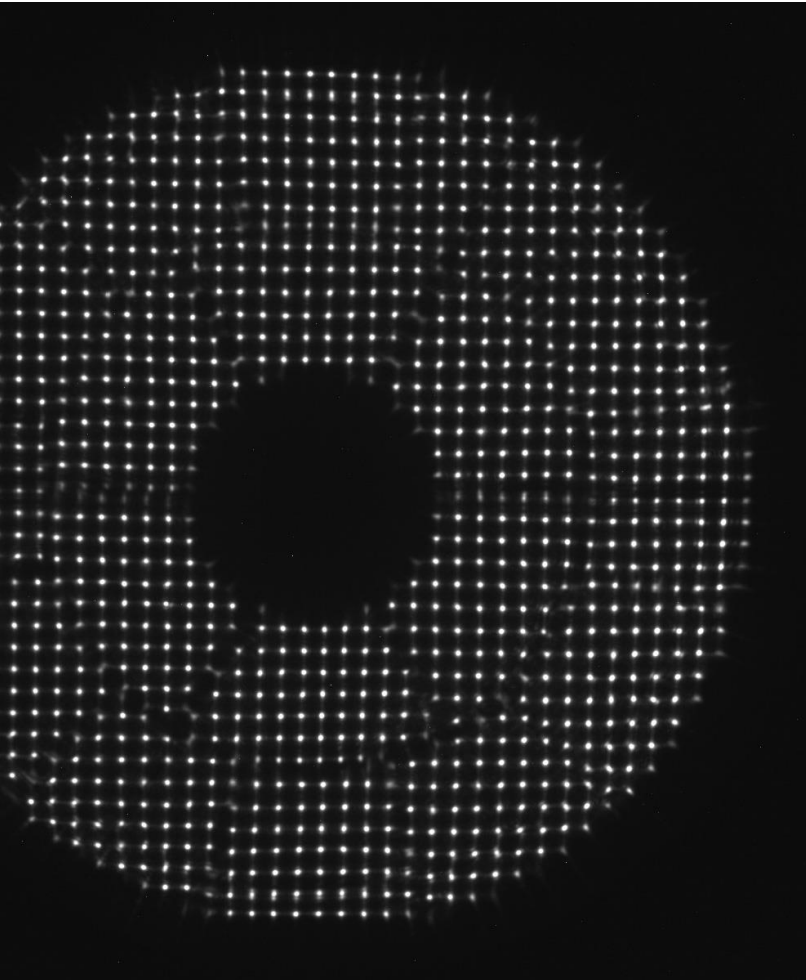
Convex Aspheric



NGC6946

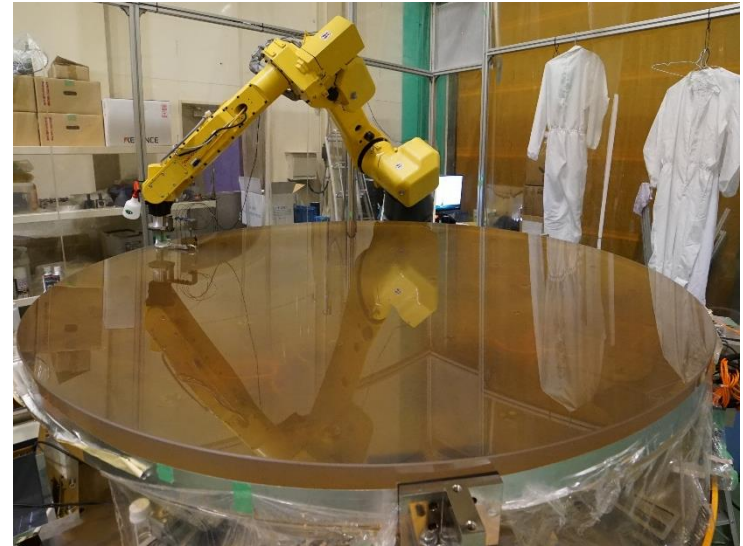


Hartmann Test

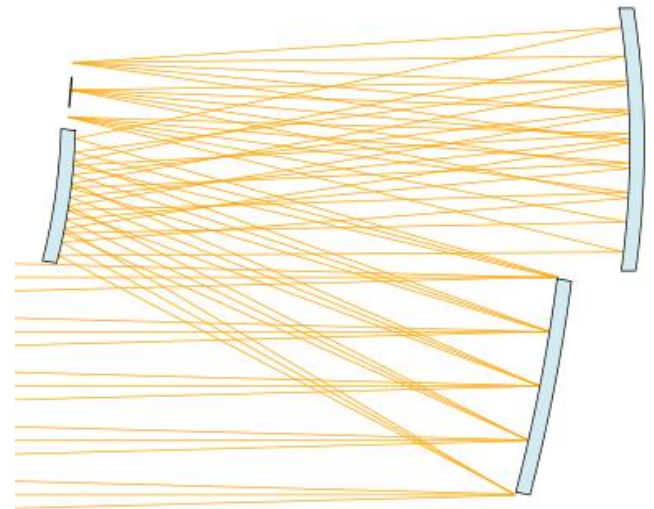


Current and Future Work

- 2-m off axis parabola
 - Telescope for observing planets in our solar system
- Three mirror anastigmat (TMA) telescope
 - 0.5 m aperture
 - FoV of $\Phi 4$ deg
 - Off-axis and good MTF



2-m off axis parabola for PLANETS



Conclusion

- We developed new manufacturing system for free form and one-meter scale optics.
- The system figured out the M2 and M3 mirrors of Seimei telescope.
- The system finished 1m convex M2 within one month with the figure error of 35 nm RMS.
- We confirmed the optical performance of M2 and M3 with the Shack Hartmann test and obtained the Hartmann constant of 0".26.