

Basic Group

2025. 09. 25. – 26.

Kagoshima University

Aims & ...

- You can make your own image from the proposed/archive data.
 - You already found adequate ALMA data
 - Band (Frequencies)/ Beam sizes (configurations)
 - When No image product in archive
 - When you want to (slightly) modify the image
- Generally, there are NO Calibration issues because the issues are already checked by ARC (QA0-2 pass)
- Download ALMA data
 - 1) Run the scriptForPI.py file & Restore the calibrated data
 - 2) Request the calibrated data to EA ARC via helpdesk
 - Ready to imaging (tclean)

NRAO Synthesis Image Summer school/workshops.

<https://web.cvent.com/event/90ae72df-7675-41b5-b056-48af11e1aa7f/summary>



09:15 - 10:00	Basics of Radio Astronomy: Slides , Video (Dominic Eddowes)
10:00 - 10:20	Coffee Break
10:20 - 11:10	Antennas and receiver systems: Slides , Video (Jay Blanchard)
11:15 - 12:00	Fundamentals of radio interferometry I: Slides , Video (Rick Perley)
12:00 - 14:00	Lunch break (self-catered)
14:00 - 14:45	Fundamentals of radio interferometry II: Slides , Video (Rick Perley)
14:45 - 15:30	Basic radio interferometry - Geometry: Slides , Video (Rick Perley)
15:30 - 16:00	Coffee Break
16:00 - 17:10	Cross Correlators: Slides , Video (Adam Deller)
17:15 - 18:15	Discussion groups
18:30 - 20:00	Welcome Reception
Thursday, 16 May	
08:00 - 09:00	Registration
09:00 - 10:00	Calibration: Slides , Video (Ian Heywood)
10:00 - 10:20	Coffee Break
10:20 - 11:20	Introduction to imaging and deconvolution: Slides , Video (Joshua Marvil)
11:20 - 12:30	Advanced calibration: Slides , Video (Ian Heywood)

There are excellent lecture materials on interferometry available online, and if you want to gain a deeper understanding, I recommend watching the lectures



Previous Workshops

[19th Synthesis Imaging Workshop](#) (2023)

[Virtual 18th Synthesis Imaging Workshop](#) (2022)

[Virtual 17th Synthesis Imaging Workshop](#) (2020)

[16th Synthesis Imaging Workshop](#) (2018)

[15th Synthesis Imaging Workshop](#) (2016)

[14th Synthesis Imaging Workshop](#) (2014)

[13th Synthesis Imaging Workshop](#) (2012)

[12th Synthesis Imaging Workshop](#) (2010)

[11th Synthesis Imaging Workshop](#) (2008)

[10th Synthesis Imaging Workshop](#) (2006)

[9th Synthesis Imaging Summer School](#) (2004)

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<https://www.eso.org/sci/facilities/alma/arc/alma-archive-school2022.html>

References

eso.org/sci/facilities/alma/arc/alma-archive-school2022.html



ALMA Advanced Data Products
ALMA Regional Centre

Final Program

Wednesday October 5	
Time	Topic
09:00	Welcome and logistics
09:30	Introduction to interferometry and ALMA
10:30	Coffee break
11:00	ALMA Science Archive basics
12:00	ALMA Science Archive content
13:00	Lunch
14:00 - 18:00	Hands-on session on archival queries (including coffee break)

Thursday October 6	
Time	Topic
09:30	ALMA Calibration basics
10:45	Coffee break
11:15	ALMA Imaging basics
13:00	Lunch
14:00 - 17:30	Hands-on session on CARTA (including coffee break)

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Quick Basics for ALMA imaging process

- Scales
- Visibility data/ Dirty beam / Dirty Image ..
- Deconvolution (clean)
- Image parameter selection
 - Pixel size/image size
 - Weightings (Natural- Briggs – Uniform)

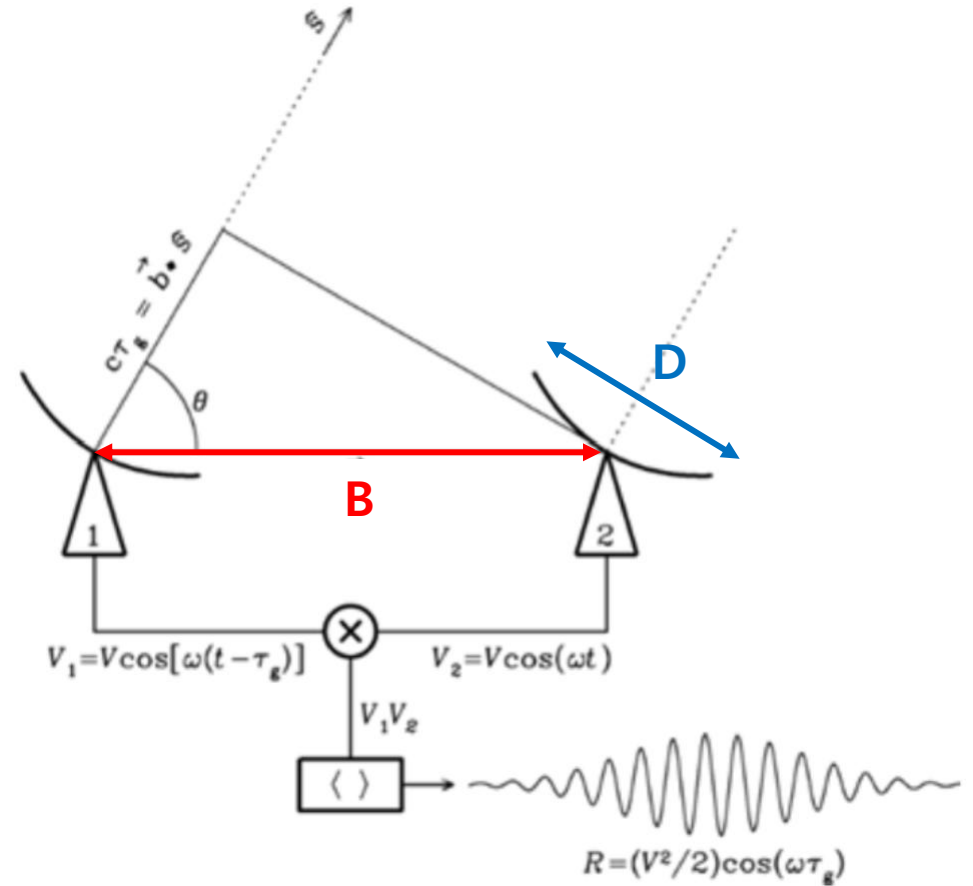
Scales

$$\theta \sim \frac{\lambda}{D(B)}$$

D : Diameter of telescope

B : Baseline; distance between antenna pair

- **Angular Resolution (beam size)**
 - depend on the longest baseline
 - $0.2'' \times (300 \text{ GHz/Freq}) \times (1 \text{ km/longest baseline})$
- **Maximum Recoverable Scale (MRS)**
 - depends on the shortest baseline ($\sim 10 \times$ beam size)
 - $1.4'' \times (300 \text{ GHz/Freq}) \times (150\text{m/shortest baseline})$
 - When the emission is more extended than MRS, the emission is resolve out $\rightarrow$ multiple configuration or ACA and TP are added.
- **Field of View (FOV)**
 - FWHM of the 12m telescope primary beam
 - $\sim 19 \text{ arcsec}$ (33 arcsec) @ 300 GHz for 12m (7m)
 - Area of target is larger than $1/3 \text{ FOV}$, mosaic observation.



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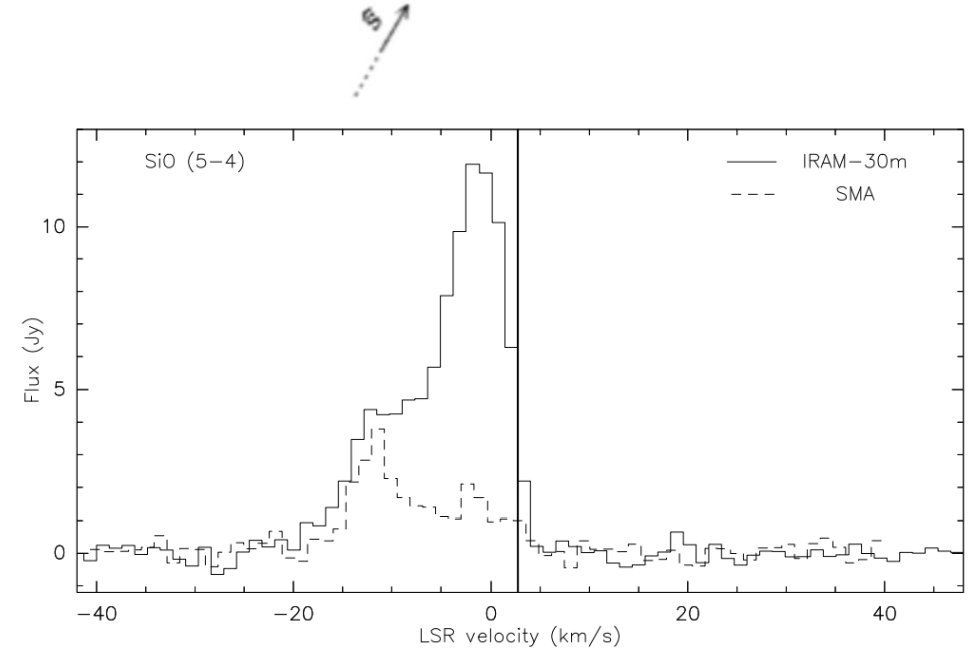


Fig. 7. SiO (5–4) spectra at the B1 position observed by IRAM 30m telescope (thin solid line) and the SMA (dashed line). Cloud velocity is indicated by a thick vertical line ($V_{\text{LSR}} = 2.7 \text{ km s}^{-1}$). The SMA observations were convolved to the 11 arcsec resolution so as to match the beam of the IRAM 30m telescope. Spectral resolution is 1.3 km s^{-1} in both cases.

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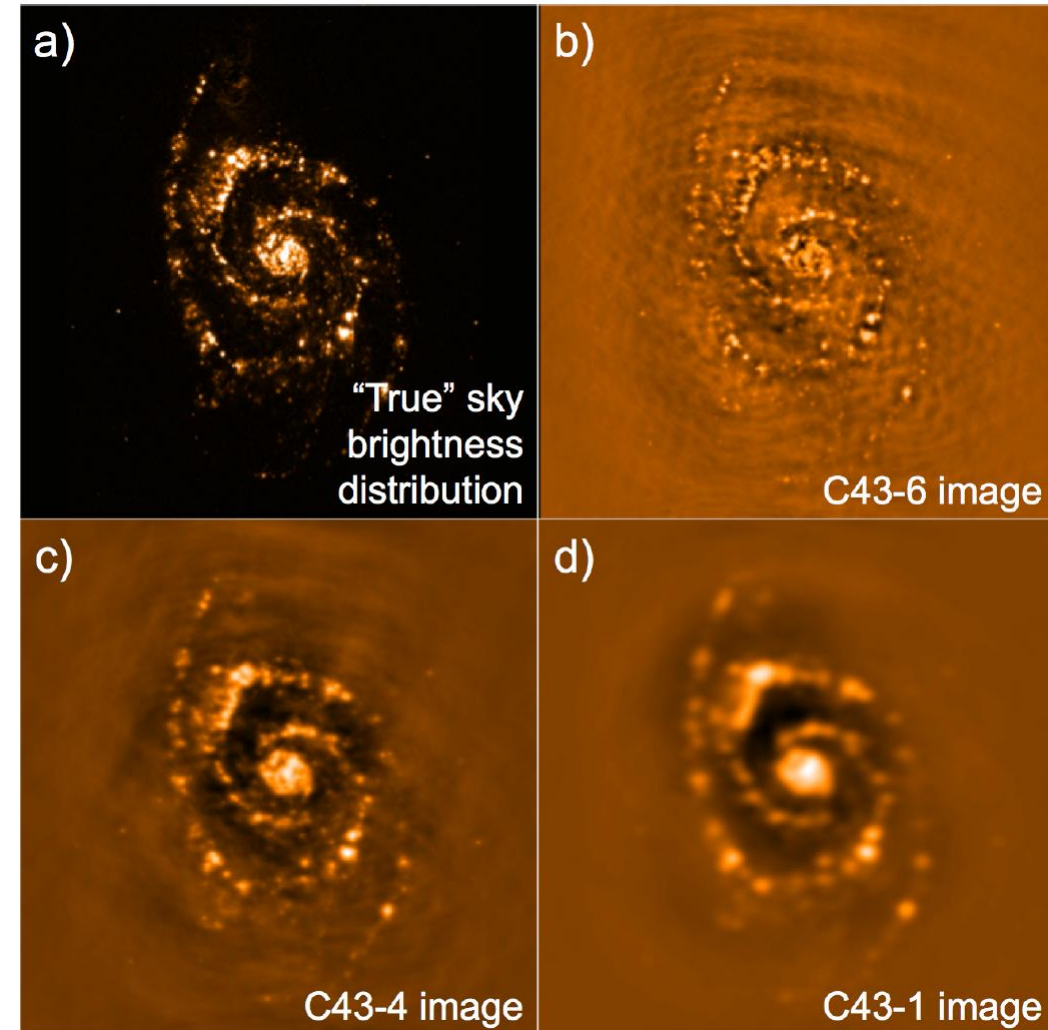
ALMA provide multi-configurations

		Band	1	3	4	5	6	7	8	9	10
Config.	L _{max}	Freq. (GHz)	40	100	150	185	230	345	460	650	870
	L _{min}										
7-m	45 m	θ_{res} (arcsec)	31.5	12.5	8.35	6.77	5.45	3.63	2.72	1.93	1.44
	9 m	θ_{MRS} (arcsec)	167	66.7	44.5	36.1	29.0	19.3	14.5	10.3	7.67
C-1	161 m	θ_{res} (arcsec)	8.45	3.38	2.25	1.83	1.47	0.98	0.74	0.52	0.39
	15 m	θ_{MRS} (arcsec)	71.2	28.5	19.0	15.4	12.4	8.25	6.19	4.38	3.27
C-2	314 m	θ_{res} (arcsec)	5.75	2.30	1.53	1.24	1.00	0.67	0.50	0.35	0.26
	15 m	θ_{MRS} (arcsec)	56.5	22.6	15.0	12.2	9.81	6.54	4.90	3.47	2.59
C-3	500 m	θ_{res} (arcsec)	3.55	1.42	0.94	0.77	0.62	0.41	0.31	0.22	0.16
	15 m	θ_{MRS} (arcsec)	40.5	16.2	10.8	8.73	7.02	4.68	3.51	2.48	1.86
C-4	784 m	θ_{res} (arcsec)	2.30	0.92	0.61	0.50	0.40	0.27	0.20	0.14	0.11
	15 m	θ_{MRS} (arcsec)	28.0	11.2	7.50	6.08	4.89	3.26	2.44	1.73	1.29
C-5	1.4 km	θ_{res} (arcsec)	1.38	0.55	0.36	0.30	0.24	0.16	0.12	0.084	0.063
	15 m	θ_{MRS} (arcsec)	16.8	6.70	4.47	3.62	2.91	1.94	1.46	1.03	0.77
C-6	2.5 km	θ_{res} (arcsec)	0.78	0.31	0.20	0.17	0.13	0.089	0.067	0.047	0.035
	15 m	θ_{MRS} (arcsec)	10.3	4.11	2.74	2.22	1.78	1.19	0.89	0.63	0.47
C-7	3.6 km	θ_{res} (arcsec)	0.53	0.21	0.14	0.11	0.092	0.061	0.046	0.033	0.024
	64 m	θ_{MRS} (arcsec)	6.45	2.58	1.72	1.40	1.12	0.75	0.56	0.40	0.30
C-8	8.5 km	θ_{res} (arcsec)	0.24	0.096	0.064	0.052	0.042	0.028	0.021	0.015	0.011
	110 m	θ_{MRS} (arcsec)	3.55	1.42	0.95	0.77	0.62	0.41	0.31	0.22	0.16
C-9	13.9 km	θ_{res} (arcsec)	0.14	0.057	0.038	0.031	0.025	0.017	0.012	0.0088	0.0066
	368 m	θ_{MRS} (arcsec)	2.03	0.81	0.54	0.44	0.35	0.24	0.18	0.13	0.093
C-10	16.2 km	θ_{res} (arcsec)	0.11	0.042	0.028	0.023	0.018	0.012	0.0091	0.0065	0.0048
	244 m	θ_{MRS} (arcsec)	1.25	0.50	0.33	0.27	0.22	0.14	0.11	0.077	0.057

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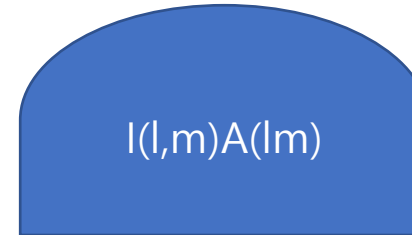


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Antenna Primary Beam Response

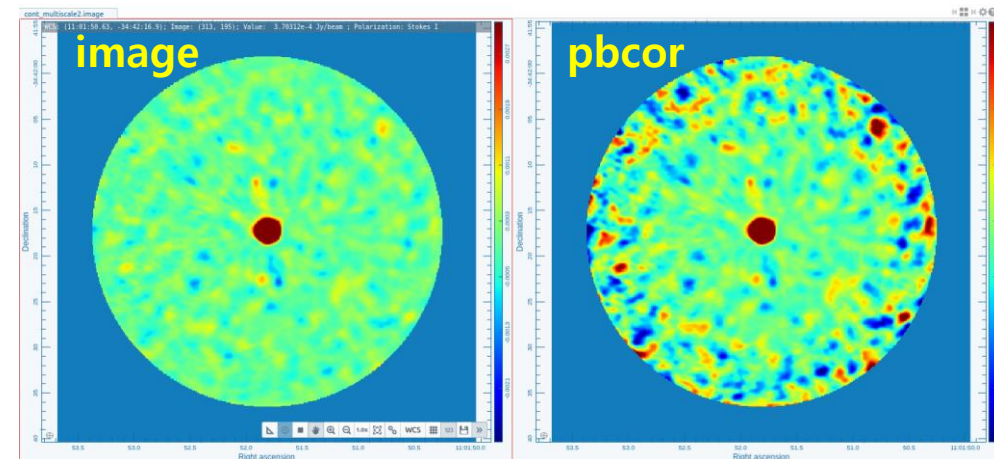


##.image



##.pbcor

The signal received incorporates the $A(l,m)$.
You should use the **pbcor** for science!!



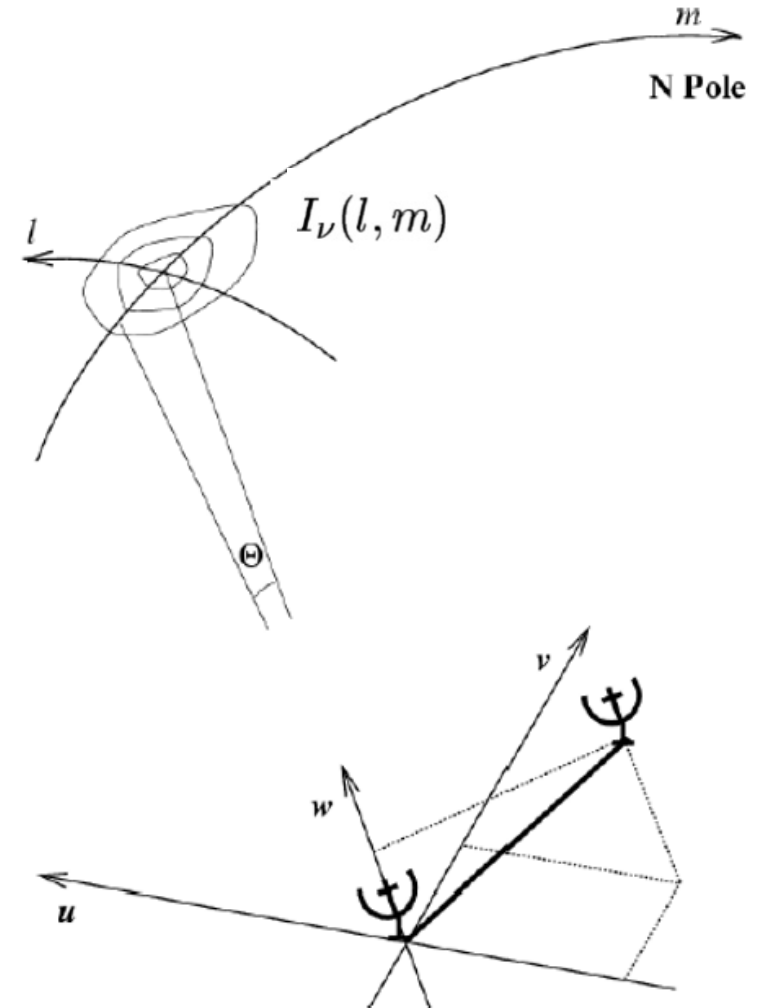
Visibility (V) and Sky Brightness (I)

$$V_{\nu}(u, v) = \iint I_{\nu}(l, m) e^{-2\pi i(ul + vm)} dl dm$$

$$I_{\nu}(l, m) = \iint V_{\nu}(u, v) e^{2\pi i(ul + vm)} du dv$$

V(u,v), the complex visibility function, is the 2D Fourier transform of I(l,m), the sky brightness distribution (van Cittert-Zernike theorem)

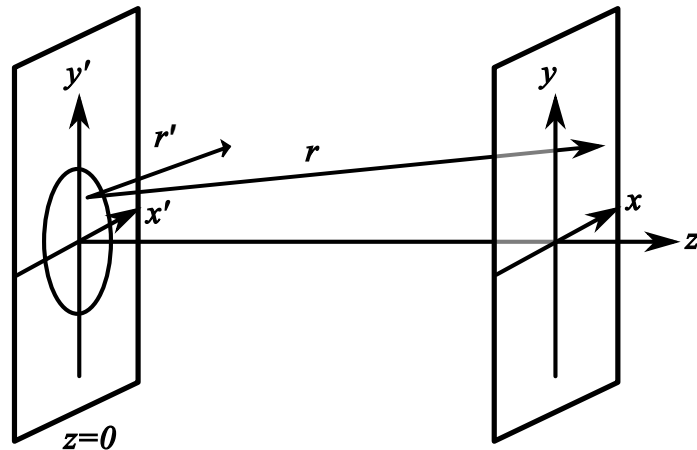
(u,v) : UV plane , distributions of Antenna pairs
(l,m) : R.A. and Dec. in sky



Fraunhofer diffraction pattern

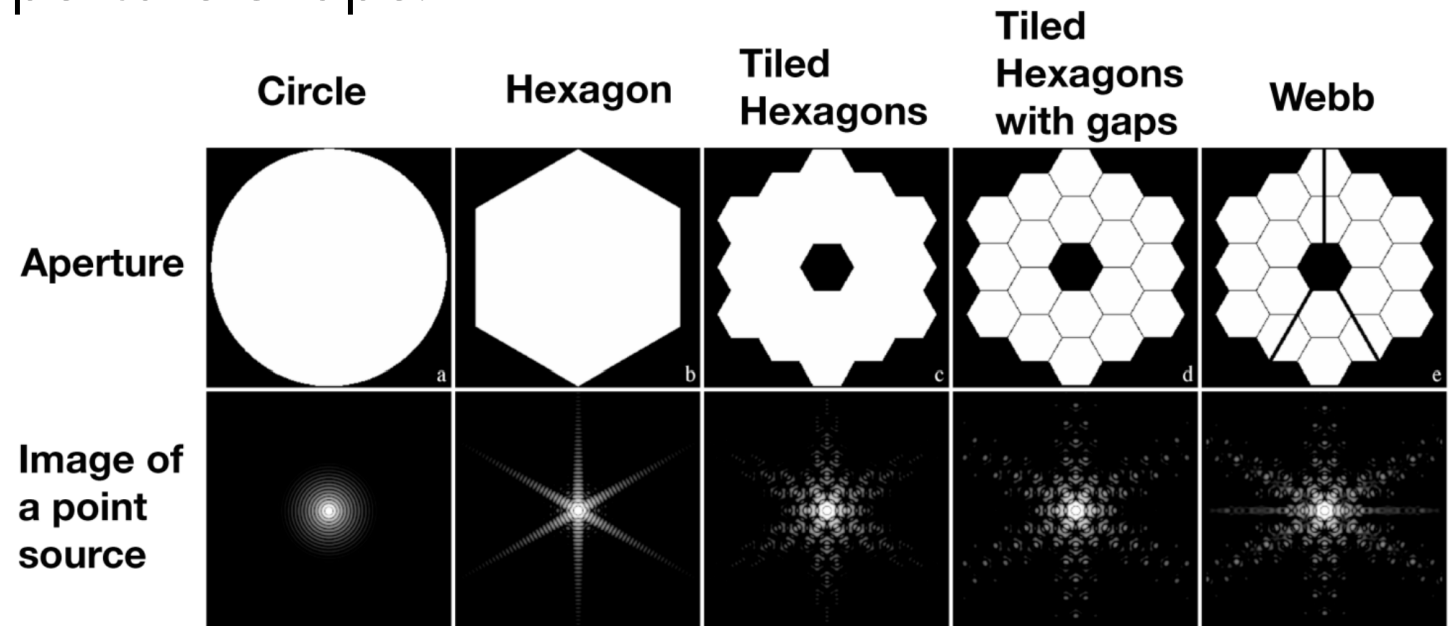
$$U(x, y, z) \propto \iint_{\text{Aperture}} A(x', y') e^{-i \frac{2\pi}{\lambda} (lx' + my')} dx' dy'$$

PSF is **Fourier Transform** of Aperture shape.



Credit: NASA/STScI

The very first finely-phased image ever released by NASA's James Webb Space Telescope shows a single image of a star, complete with six prominent diffraction spikes (and two less-prominent ones), with background stars and galaxies revealed behind it. The background galaxies were a surprise to astronomers; JWST is imaging the Universe at roughly double the performance precision it was design-specified for. Even images such as this, not originally designed for scientific purposes, may prove useful to astronomers studying the Universe as a unique and unexpected source of data.

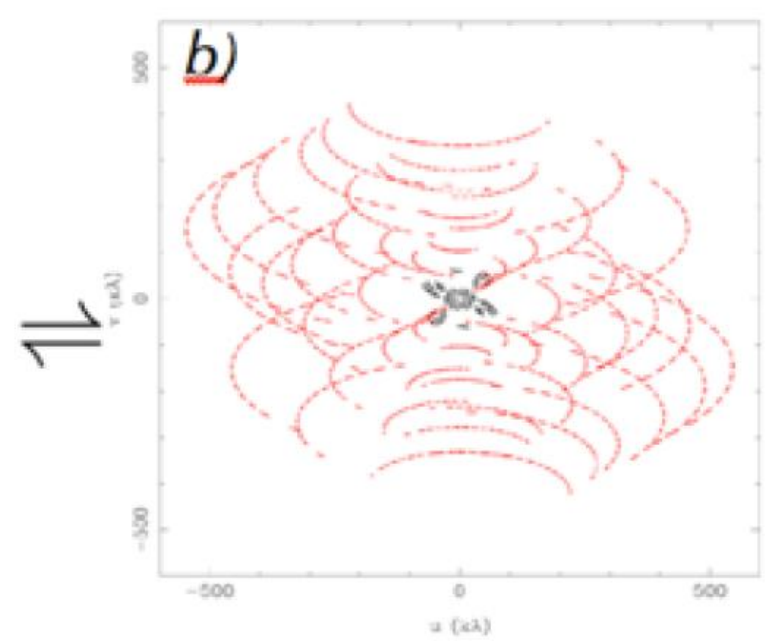
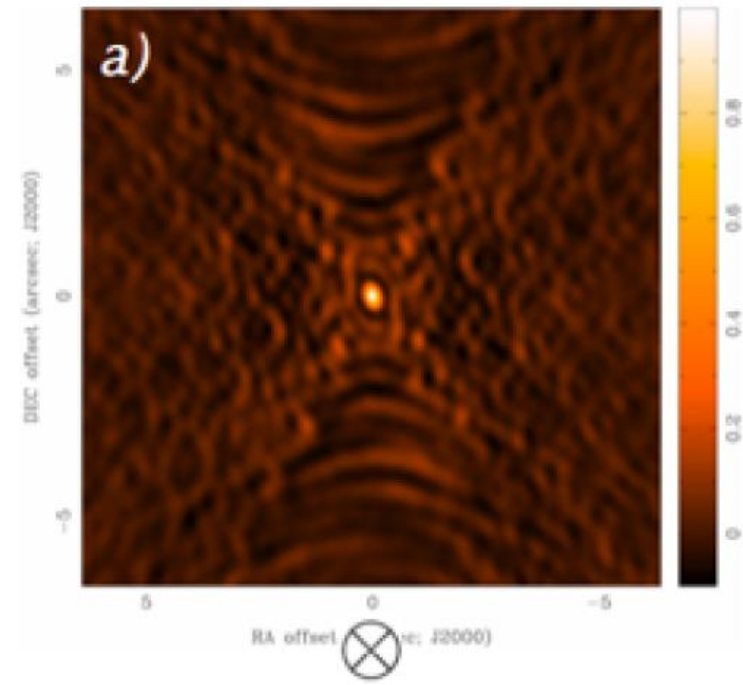


Credit: R. B. Makidon, S. Casertano, C. Cox & R. van der Marel, STScI/NASA/AURA

Caution: Although the PSF (dirty beam) is the Fourier transform of the aperture shape (UV distribution), the details are quite different.

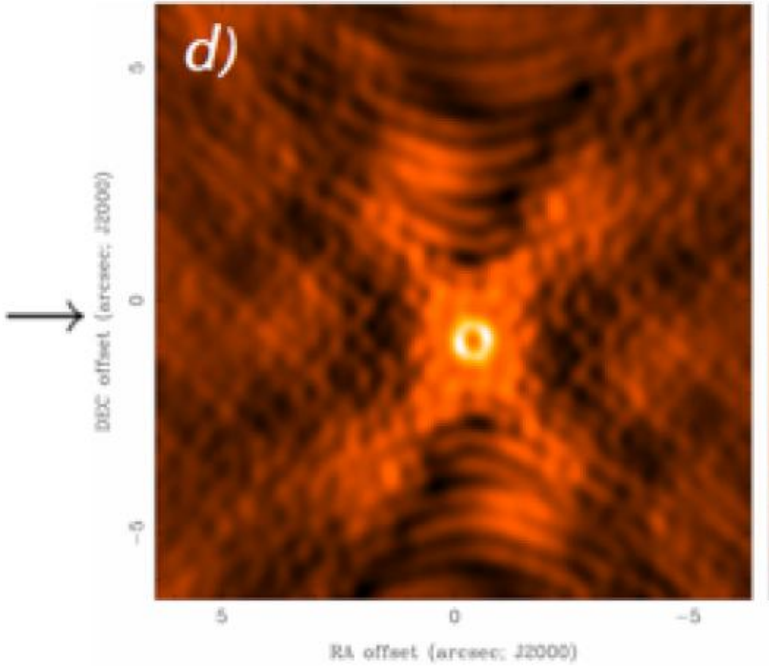
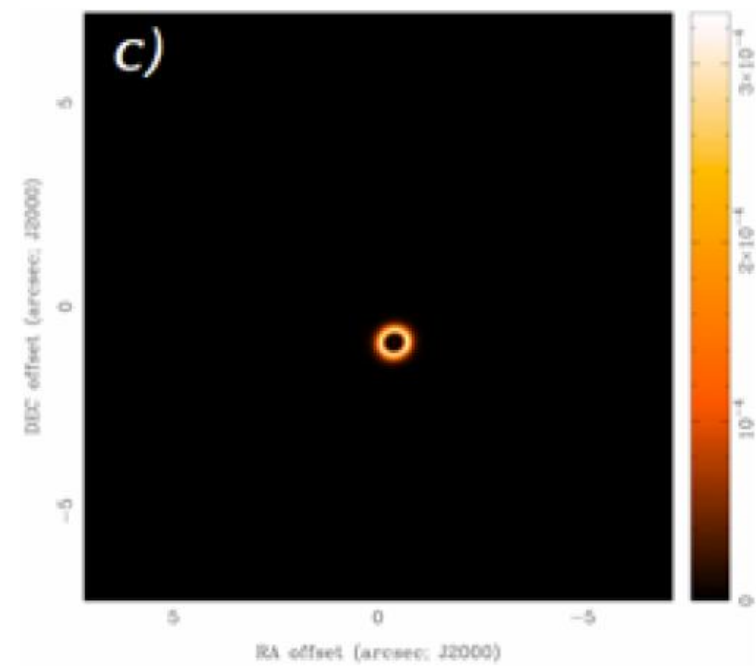
Dirty beam (PSF) & Dirty image (PSF-convolved image)

$b(l,m)$
(dirty
beam)



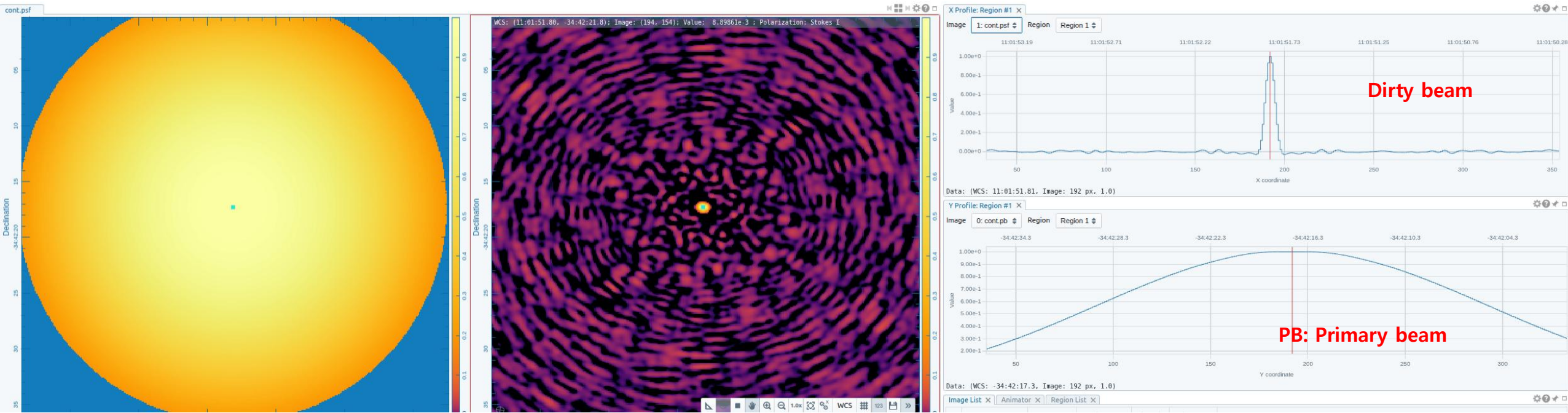
$B(u,v)$

$I(l,m)$



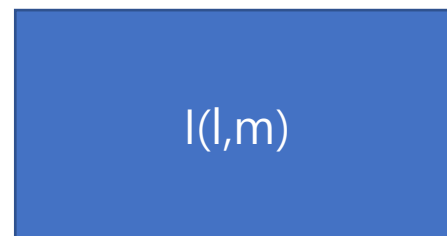
$I^D(l,m)$
(dirty
image)

Primary beam (single dish) vs. Dirty Beam

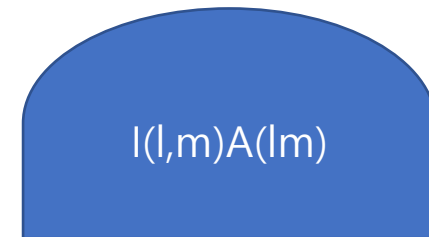


PB: Primary beam

Dirty beam

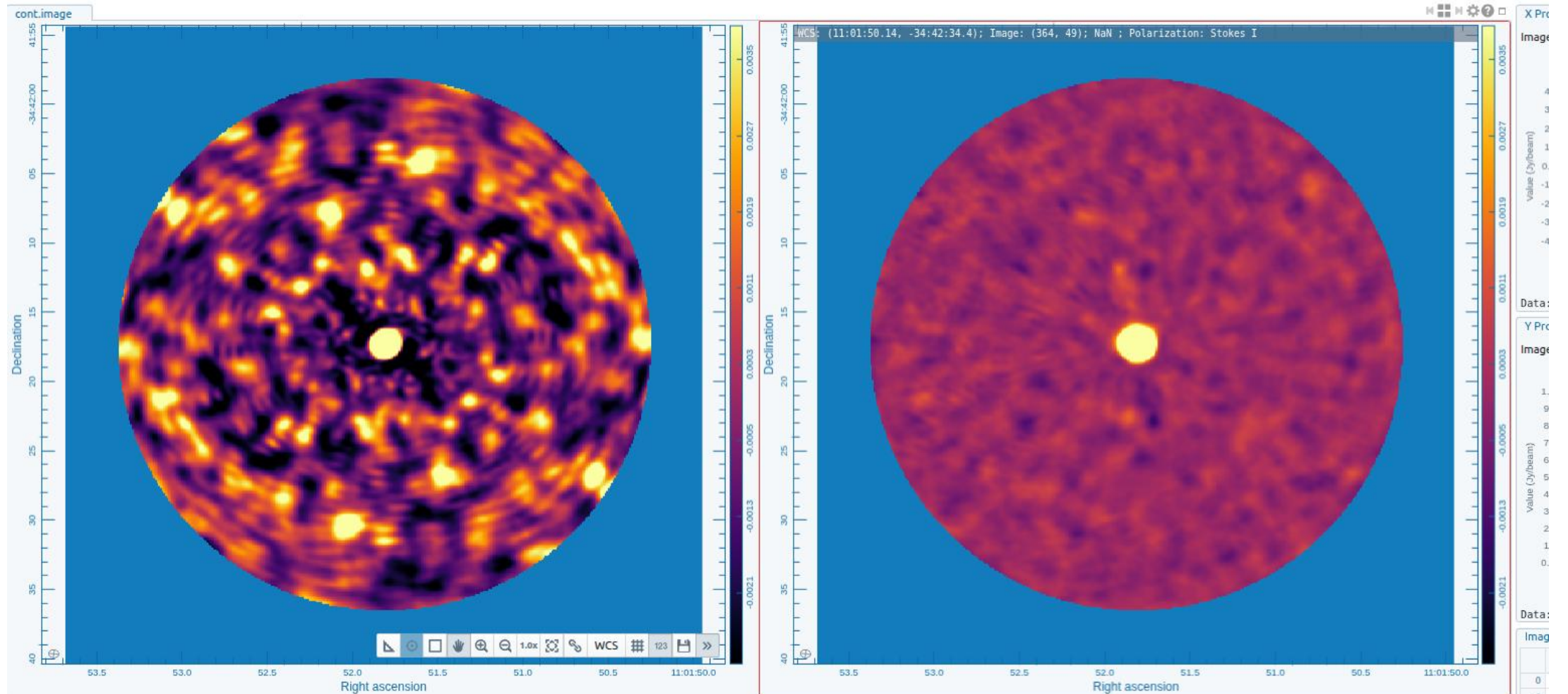


##.pbcor

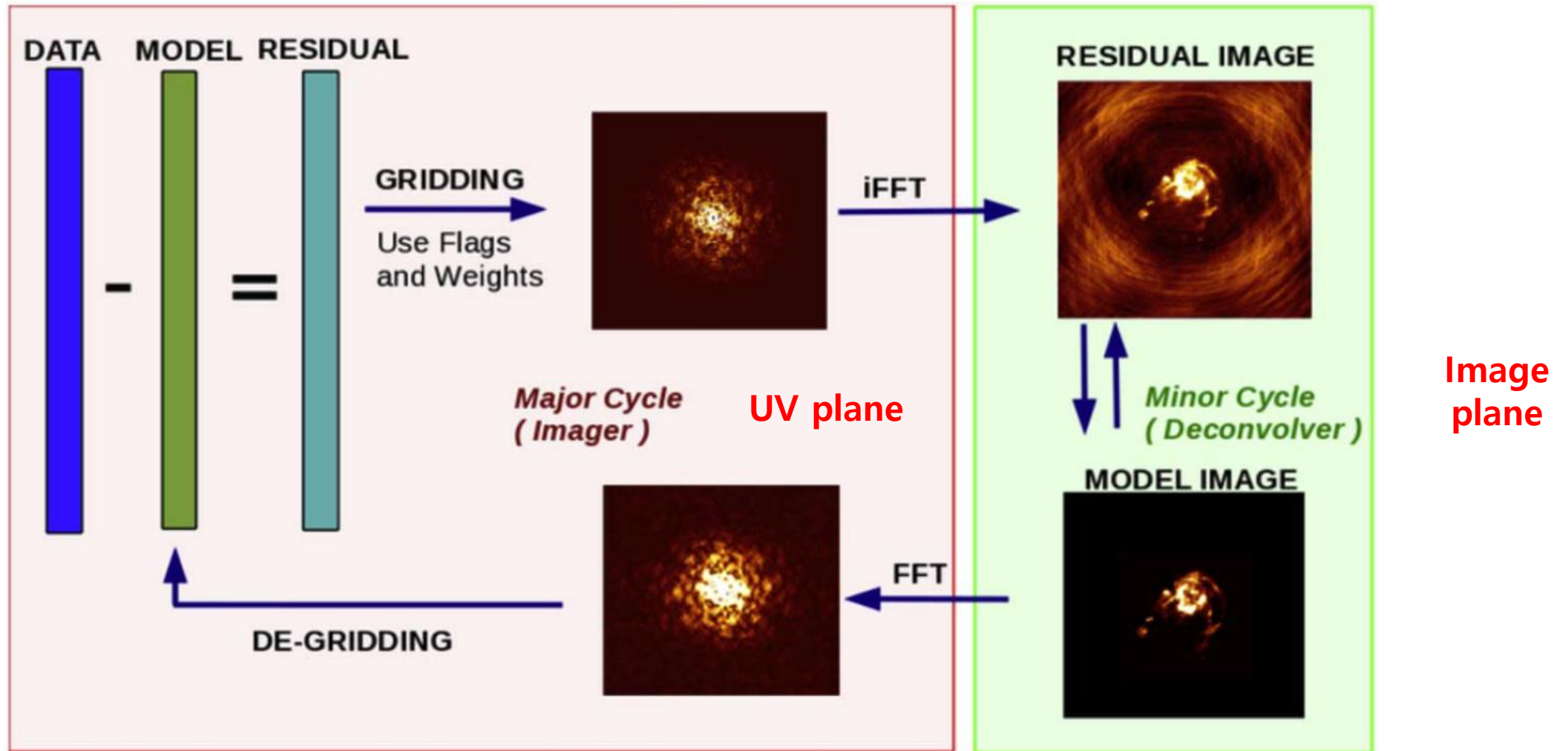


= ##.image / ##.pb

Dirty image vs. cleaned image

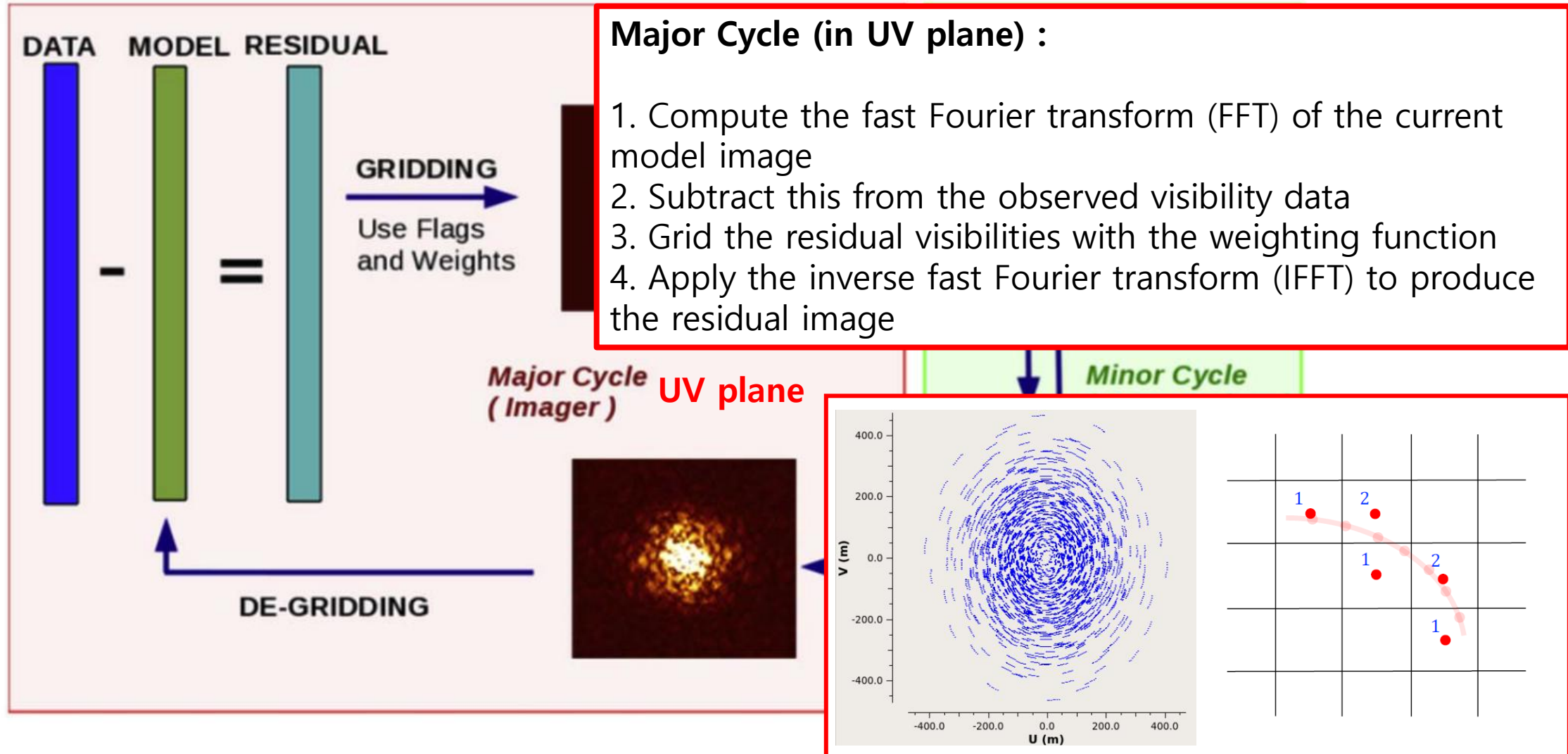


Deconvolution (tclean)



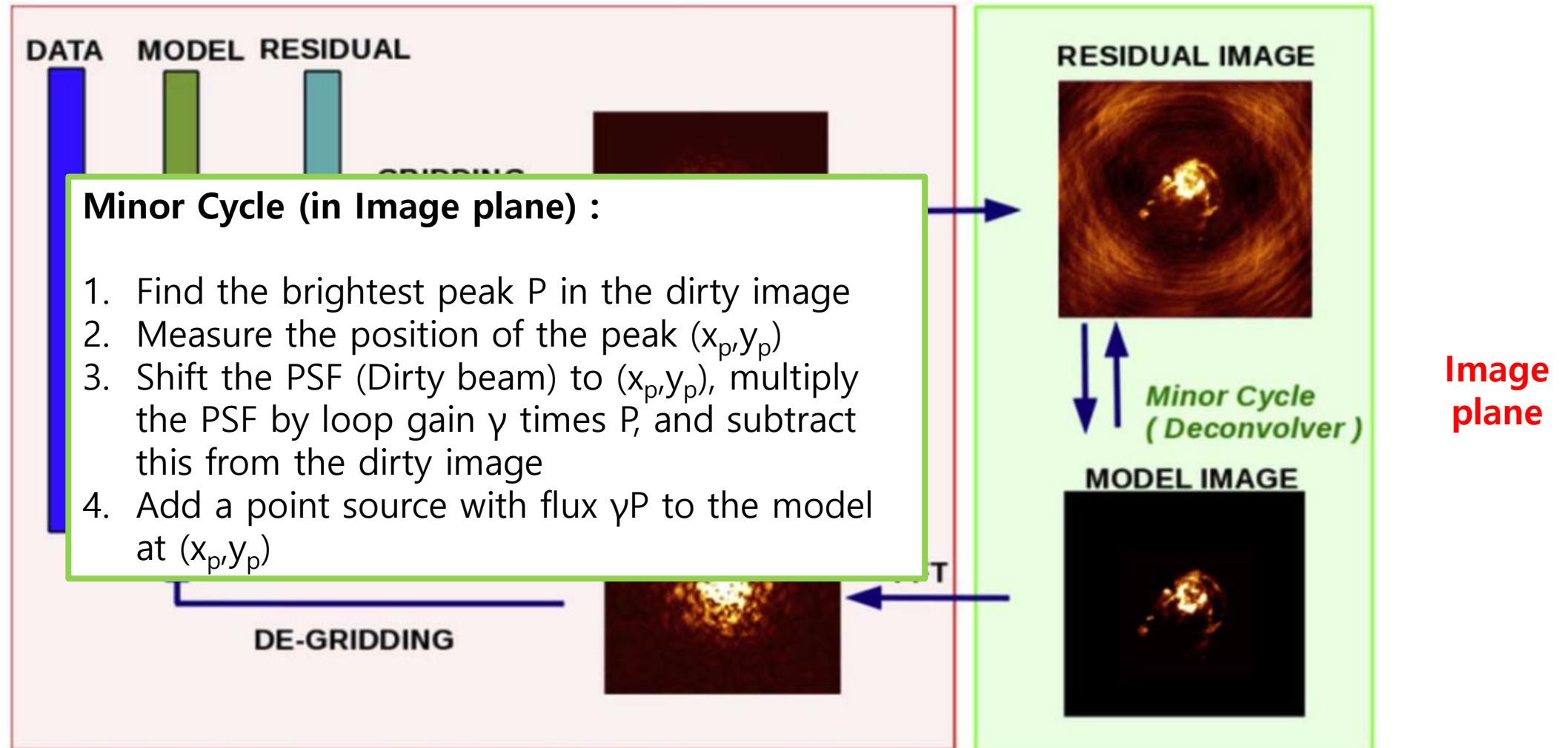
See 'An introduction to CLEAN' (<https://youtube.com/watch?v=fF3KetPUyFE>)

Deconvolution (tclean)



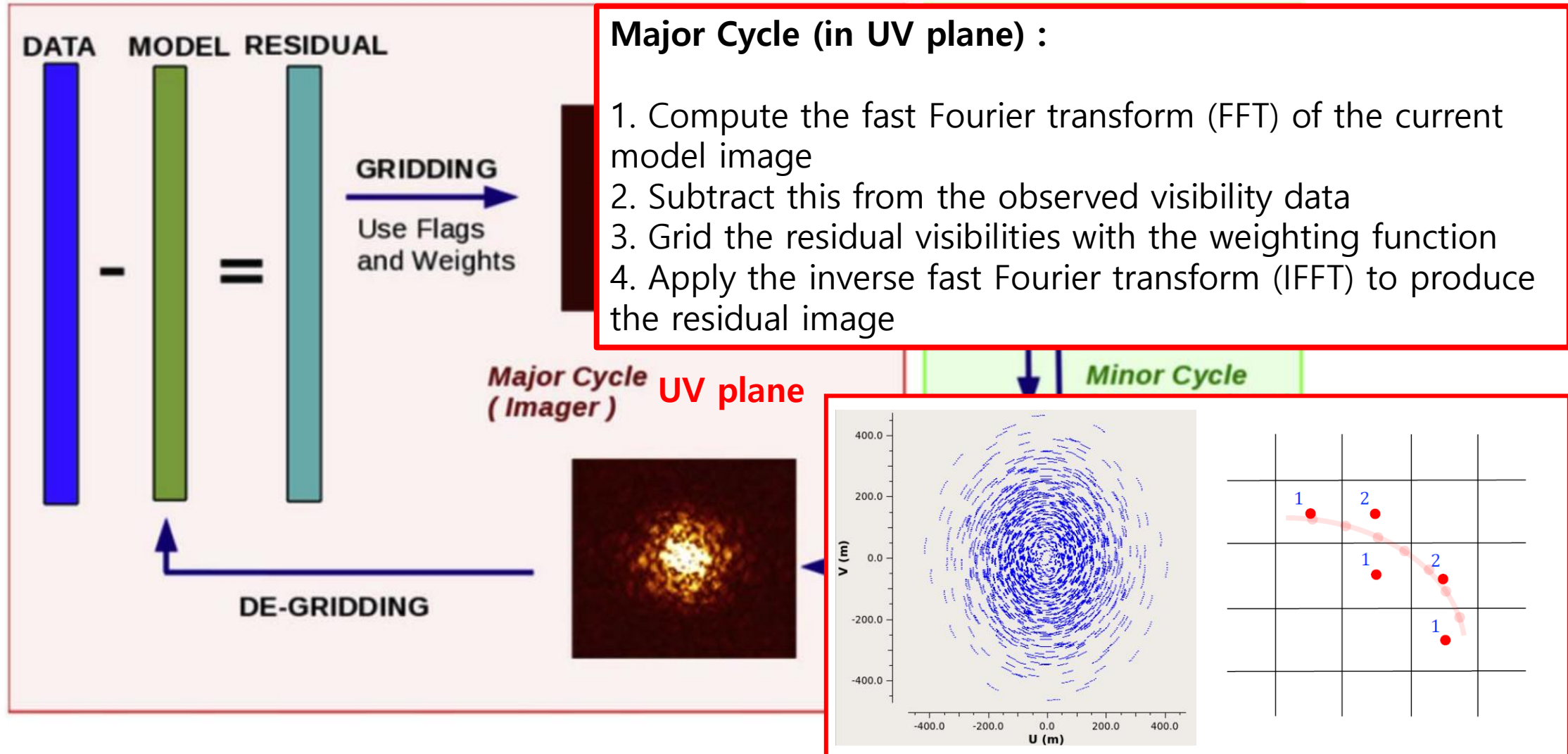
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Deconvolution (tclean)



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Deconvolution (tclean)

Cleaned Image
=
Model convolved
with main lobe
of PSF
+
Residual image

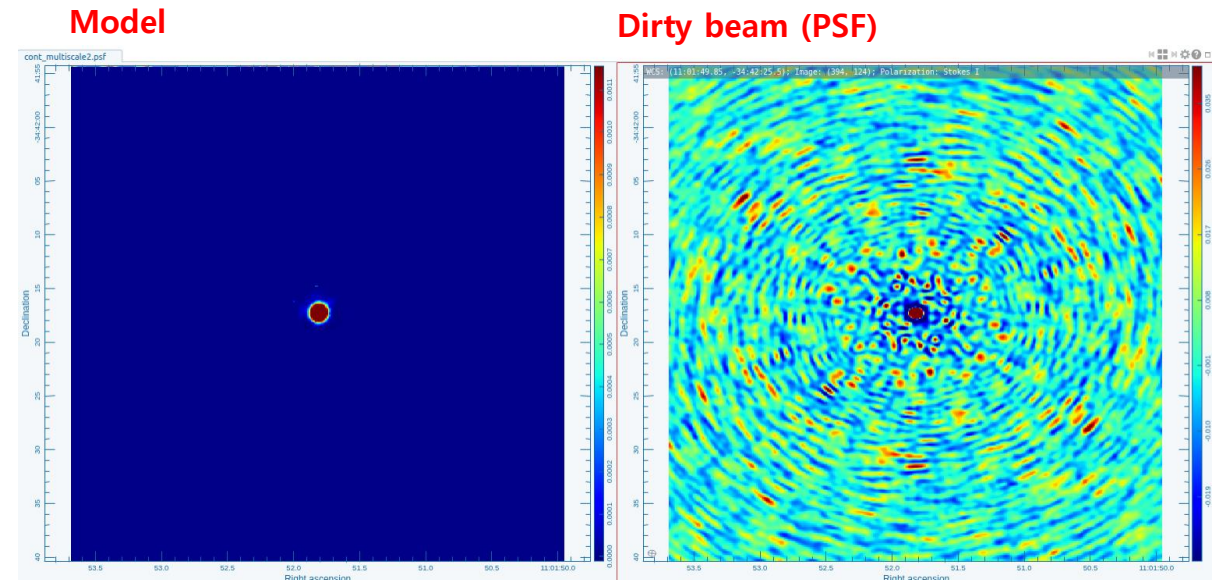
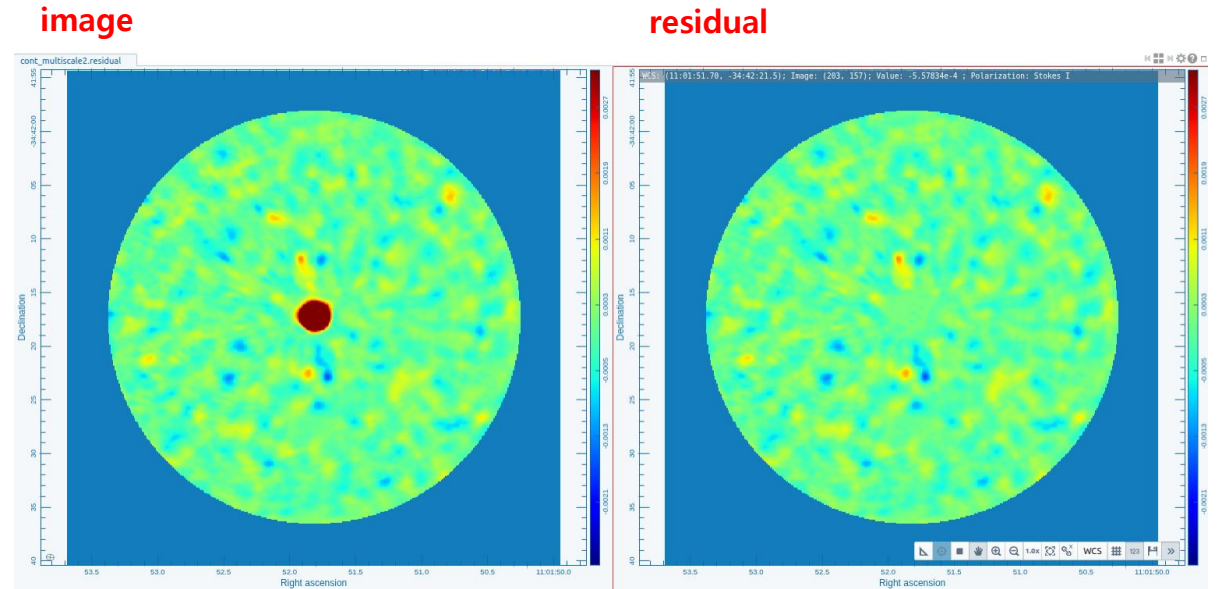


Image Parameter Selection

- pixel size : 5-8 times of minor axis of beam
- Image size : cover the Primary Beam
- specmode :
 - Continuum : 'mfs' (multi-frequency synthesis)
 - Line : 'cube'
- Deconvolver : 'Hogbom', 'Clark'(fast,simple), 'multi-scale'; 'mtmfs'
- Niter/threshold : how many to clean before stopping

Image Parameter Selection

- Weightings
 - Default is Briggs weighting with robust = 0.5
 - +2 : Natural & -2 : Uniform
 - Uniform : Higher resolution
 - Natural : Higher SNR or Accurate flux

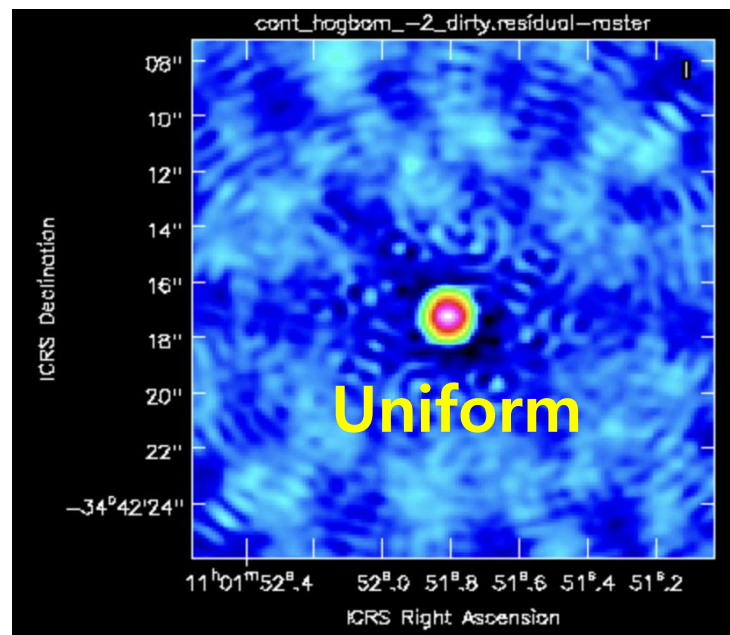
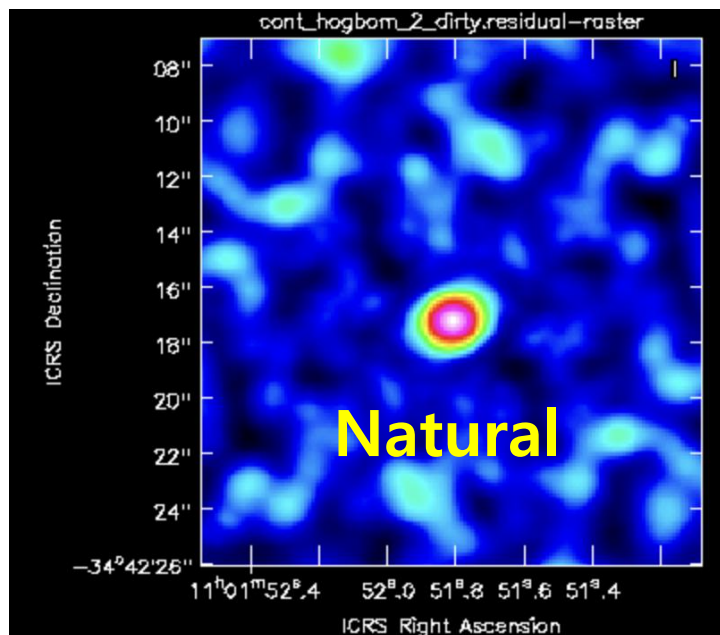


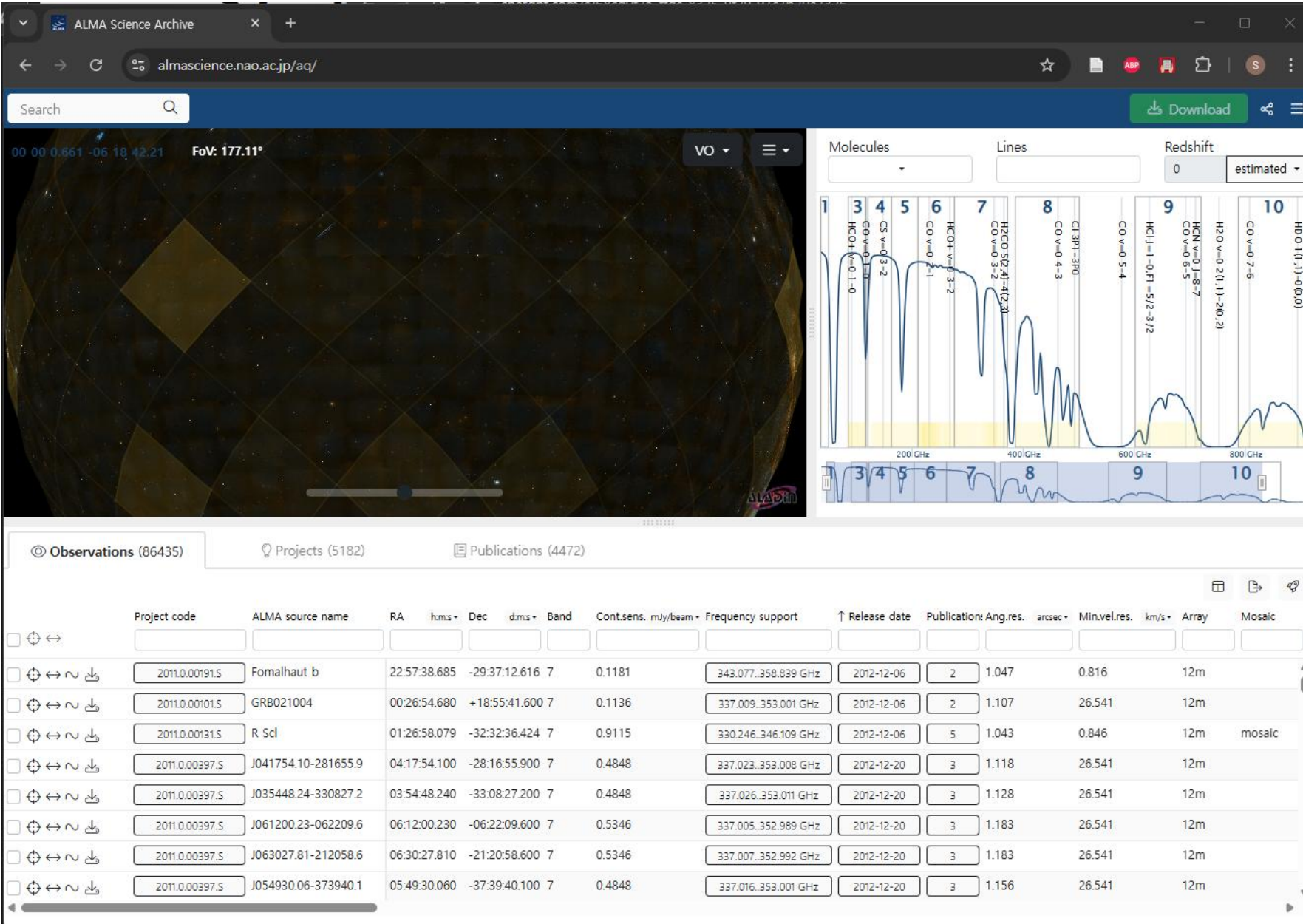
Image Parameter Selection

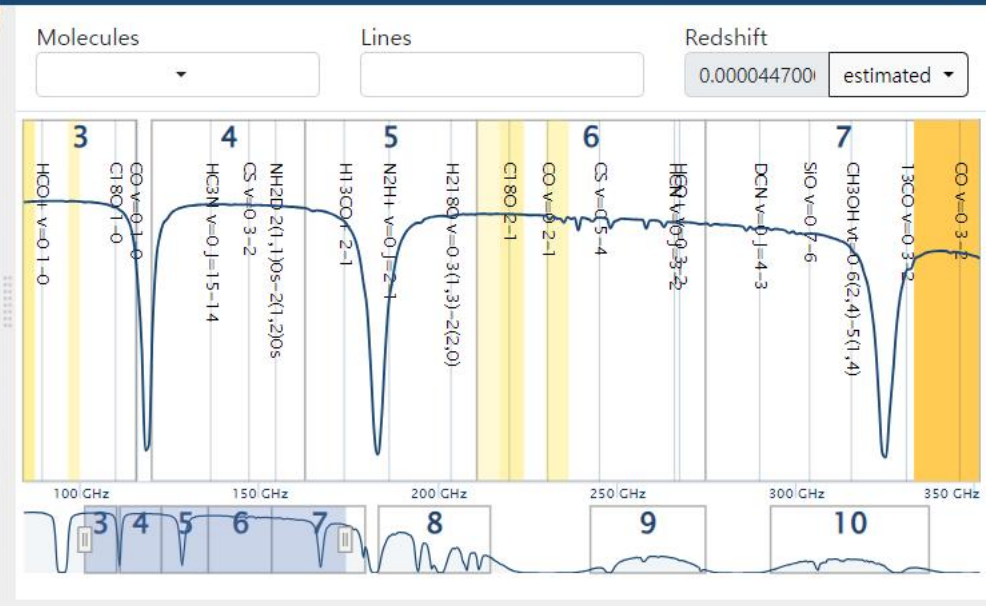
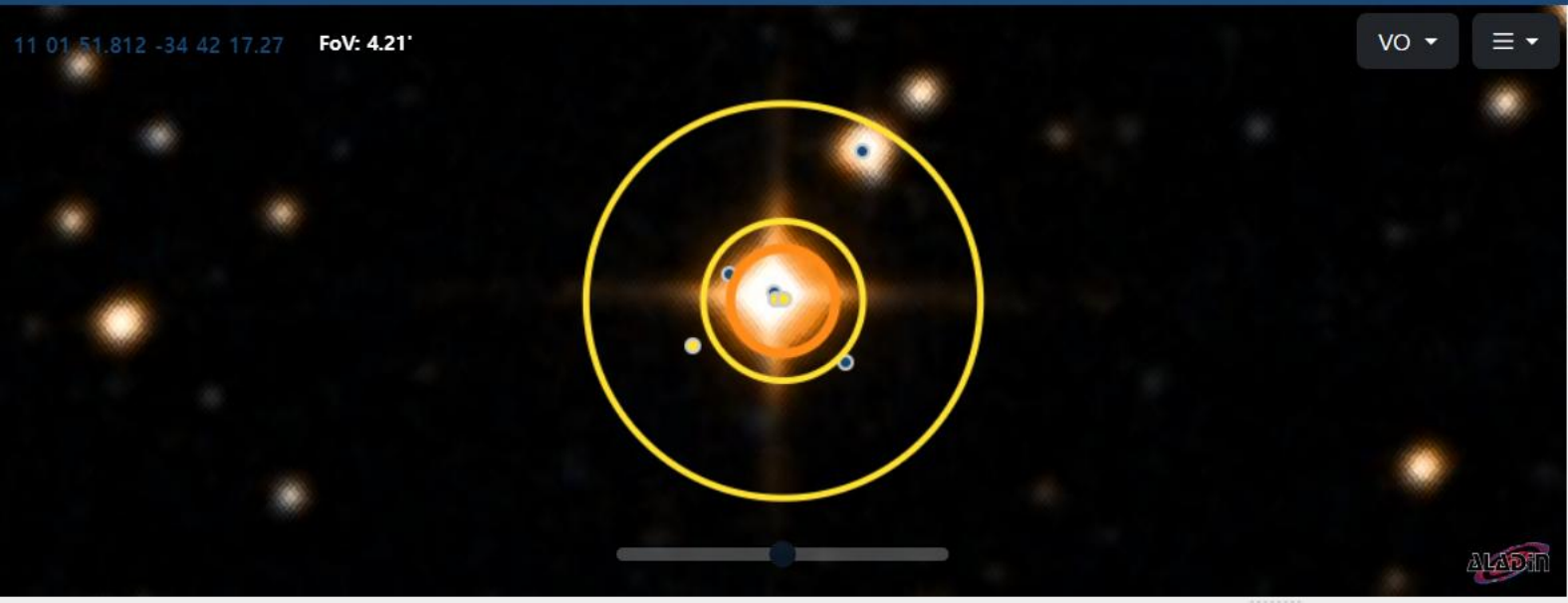
- Taper
 - Make larger beam size : gaussian smoothing
- UV range
 - If you compare the flux between observations, you need to
 - If a few short baseline make stripes or a few long baseline make tile-like features, you can exclude them.

Download Data

1. Download data (raw + auxiliary data) from ALMA Science Portal & Restore it using casa pipeline
2. Request the calibrated data via helpdesk
(even case 2, download the auxiliary data for weblog)

https://almascience.nao.ac.jp/aq/





Project code: 2016.1.00229.S Remove tab filters

	Project code	ALMA source name	RA	h:m:s	Dec	d:m:s	Band	Cont.sens.	mJy/beam	Frequency support	Release date	Publications	Ang.res.	arcsec	Min.vel.res.
☒	2016.1.00229.S	tw_hya	11:01:51.817		-34:42:17.270		7	0.0329		336.325..351.281 GHz	2018-03-03	4	0.293		0.209
☐	2016.1.00229.S	TW_Hya	11:01:51.813		-34:42:17.276		6	0.0224		217.108..235.289 GHz	2018-05-07	4	0.660		0.158
☐	2016.1.00229.S	TW_Hya	11:01:51.812		-34:42:17.263		3	0.0182		84.08..98.847 GHz	2018-09-06	4	0.318		0.186

Position

Source name

ALMA source name

RA Dec

Galactic

Target List

Angular Resolution

Max. Recoverable Scale

Energy

Frequency

Band

Project

Project code

Project Title

Publication

BibCode

Publication Title

Abstract

First Author

Authors

Observation

Observation Date

Polarisation Type

Member ous id

Object type

☐ Public data only

☐ Calibration observations

RA Dec

Info: Right Ascension and Declination

Description: Coordinate search with default radius of 10 arcmin or coordinate-range search. RA and Dec may be expressed in sexagesimal or in decimal degrees. An alternative search **radius** in arcmin can be added to the end separated by a comma. All observations that have footprints overlapping with the search region will be returned.

Units: Sexagesimal

Examples:

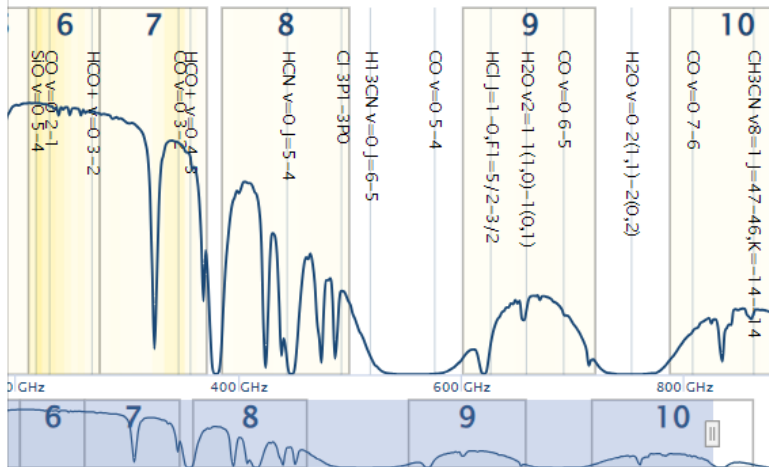
- 13:37:00.89 -29:51:59.8
- 83.633075 22.014494
- 04:31:38.425 18:13:57.242, 5
- 181.185 >-0.1928

Lines

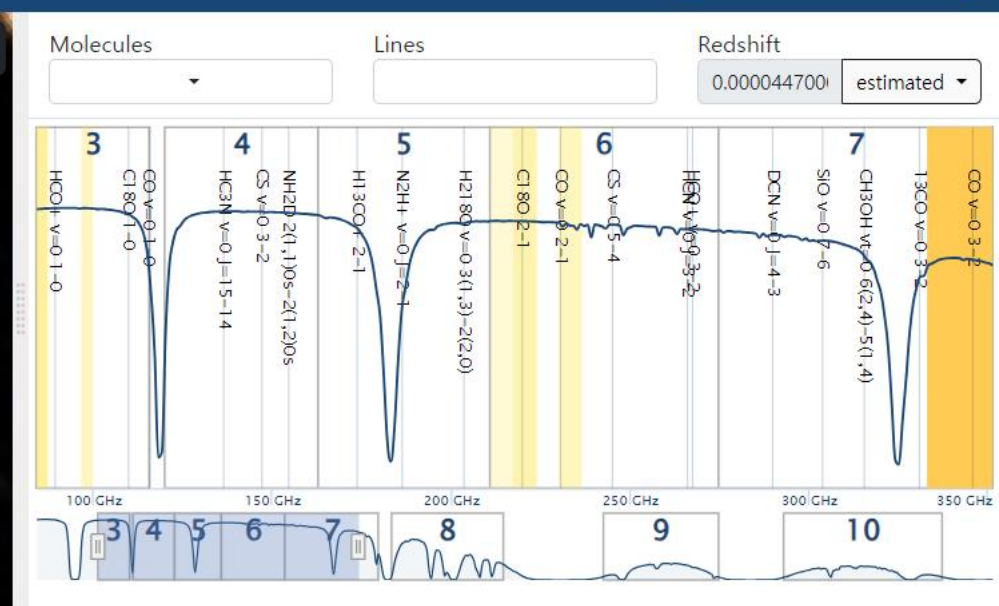
Redshift

0

estimated



Project code: 181.185 >-0.1928												
	Project code	RA	Dec	Band	Cont.sens.	Frequency support	Release date	Publications	Ang.res.	Min.vel.		
☐	2011.0.00191.S	Fomalhaut b	22:57:38.685	-29:37:12.616	7	0.1181	343.077..358.839 GHz	2012-12-06	2	1.047	0.816	
☐	2011.0.00101.S	GRB021004	00:26:54.680	+18:55:41.600	7	0.1136	337.009..353.001 GHz	2012-12-06	2	1.107	26.541	
☐	2011.0.00131.S	R Scl	01:26:58.079	-32:32:36.424	7	0.9115	330.246..346.109 GHz	2012-12-06	5	1.043	0.846	
☐	2011.0.00397.S	J035448.24-330827.2	03:54:48.240	-33:08:27.200	7	0.4848	337.026..353.011 GHz	2012-12-20	3	1.128	26.541	
☐	2011.0.00397.S	J041754.10-281655.9	04:17:54.100	-28:16:55.900	7	0.4848	337.023..353.008 GHz	2012-12-20	3	1.118	26.541	
☐	2011.0.00397.S	J063027.81-212058.6	06:30:27.810	-21:20:58.600	7	0.5346	337.007..352.992 GHz	2012-12-20	3	1.183	26.541	



Observations (3) Projects (4698) Publications (3803)

Project code: 2016.1.00229.S Remove tab filters

	Project code	ALMA source name	RA	Dec	Band	Cont.sens.	Frequency support	Release date	Publications	Ang.res.	Min.vel.res.
☒	2016.1.00229.S	tw_hya	11:01:51.817	-34:42:17.270	7	0.0329	336.325..351.281 GHz	2018-03-03	4	0.293	0.209
☐	2016.1.00229.S	TW_Hya	11:01:51.813	-34:42:17.276	6	0.0224	217.108..235.289 GHz	2018-05-07	4	0.660	0.158
☐	2016.1.00229.S	TW_Hya	11:01:51.812	-34:42:17.263	3	0.0182	84.08..98.847 GHz	2018-09-06	4	0.318	0.186

[Download 34 GB](#)[Open legacy Request Handler](#)[Login](#)

Project (1)

Group ObsUniSet (1)

Member ObsUniSet (1)

Source (1)

Collection (1)

Array (1)

File type (5)

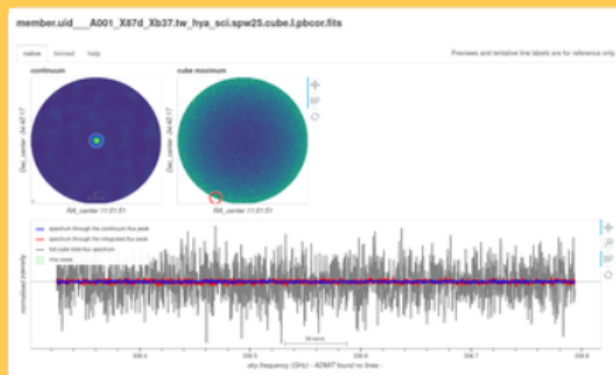
File class (3)

☒ Readme ☒ Product ☒ Auxiliary ☐ Raw ☐ Raw (semipass) ☐ External

	Name	Size	↑ Project	↑ GOUS
--	------	------	-----------	--------

☒	☒	member.uid_A001_X87d_Xb37.tw_hya_sci.spw27.cube.lpb.fits.gz	(product)	1 GB	2016.1.00229.S	uid://A001/X87d
-------------------------------------	-------------------------------------	---	-----------	------	----------------	-----------------

☒	☒	member.uid_A001_X87d_Xb37.tw_hya_sci.spw25.cube.lpbcor.fits	(product)	4 GB	2016.1.00229.S	uid://A001/X87d
-------------------------------------	-------------------------------------	---	-----------	------	----------------	-----------------



Band: 7
 Frequency range: 336.325..336.794
 Frequency resolution: 244.141 kHz
 Line sens. (10km/s): 0.599mJy/beam
 Line sens. (native): 0.093uJy/beam
 Polarizations: XX YY
 Array: 12m

☒	☒	member.uid_A001_X87d_Xb37.tw_hya_sci.spw19_25_27_29.cont.lpb.fits.gz	(product)	374 kB	2016.1.00229.S	uid://A001/X87d
-------------------------------------	-------------------------------------	--	-----------	--------	----------------	-----------------

☒	☒	member.uid_A001_X87d_Xb37.tw_hya_sci.spw29.mfs.lpbcor.fits	(product)	1 MB	2016.1.00229.S	uid://A001/X87d
-------------------------------------	-------------------------------------	--	-----------	------	----------------	-----------------



Band: 7
 Frequency range: 349.099..349.568
 Frequency resolution: 244.141 kHz
 Line sens. (10km/s): 0.593mJy/beam
 Line sens. (native): 0.094uJy/beam
 Polarizations: XX YY
 Array: 12m

ALMA Request Handler

Login

Anonymous User: Request #1417934760309

Request Title: [click to edit](#)

Download Selected

☒ readme ☒ product ☒ auxiliary ☐ raw ☐ raw (semipass) ☐ external

Project / OUSet / Executionblock	Updated	File	Size	Accessible	Actions
Request 1417934760309			114 GB		
Project 2016.1.00229.S					
Science Goal OUS uid://A001/X87d/Xb35					
Group OUS uid://A001/X87d/Xb36					
Member OUS uid://A001/X87d/Xb37	2020-07-15				
SB tw_hya_a_07_TM1					
☒ readme		member.uid_A001_X87d_Xb37_README.txt	17 kB	✓	
☐ product		2016.1.00229.S_uid_A001_X87d_Xb37_001_of_001.tar	18 GB	✓	
☒ auxiliary		2016.1.00229.S_uid_A001_X87d_Xb37_auxiliary.tar	275 MB	✓	
☒ raw		2016.1.00229.S_uid_A002_Xbb00cd_X32c.asdm.sdm.tar	113 GB	✓	

Each Member OUS (or SB) may have the following files available for download:

- readme** A text file with very basic information
- product** Final images and image cubes
- auxiliary** A file containing logs, quality assurance information, scripts, and calibration data
- raw** Raw visibility data
- external** Enhanced data products (including enhanced images or visibility data) created after the data delivery

In download_XXX.sh, check the server!!

```
LIST=( "  
https://almascience.nao.ac.jp/dataPortal/2016.1.00229.S_uid___A001_X87d_Xb3d_auxiliary.tar  
https://almascience.nao.ac.jp/dataPortal/2016.1.00229.S_uid___A002_Xbfcd9b_X70c.asdm.sdm.tar  
https://almascience.nao.ac.jp/dataPortal/member.uid___A001_X87d_Xb3d.README.txt  
")
```

```
[shlee@pluto TW_Hya]$ ls  
2016.1.00229.S  
2016.1.00229.S_uid___A001_X87d_Xb3d_auxiliary.tar  
2016.1.00229.S_uid___A002_Xbfcd9b_X70c.asdm.sdm.tar  
download-files_twhya.sh  
member.uid___A001_X87d_Xb3d.README.txt  
[shlee@pluto TW_Hya]$ cd 2016.1.00229.S/science_goal.uid___A001_X87d_Xb3b/group.  
uid___A001_X87d_Xb3c/member.uid___A001_X87d_Xb3d/  
[shlee@pluto member.uid___A001_X87d_Xb3d]$ ls  
calibrated calibration log product qa raw script  
[shlee@pluto member.uid___A001_X87d_Xb3d]$ ls script/  
casa-20240703-150200.log      ipython-20240703-150206.log      scriptForImaging.py  
casa_piperestorescript.py  PPR_uid___A001_X87d_Xb3e.xml    scriptForPI.py  
casa_pipescript.py          scriptForImagingPrep.py
```



```
| -- project_id/  
| | -- science_goal.ouss_id/  
| | | -- group.ouss_id/  
| | | | -- member.ouss_id/
```

Project - All observations associated with a specific proposal.

Science Goal OUS - All observations associated with a specific science goal in that proposal.

Group OUS – Associated observations within a Science Goal (e.g., observations of the same fields with the same spectral tunings but with **different arrays** or **array configurations**).

Member OUS – A specific set of observations of the same fields using the same tunings and array or array configuration.

Execution block – An individual “unit” of the observations needed for a Member OUS.

Atacama Large Millimeter/submillimeter Array (ALMA)

#####

Cycle: 4

Project code: 2016.1.00229.S

SB name: TW_Hya_a_06_TM1

PI name: Edwin Bergin

Project title: Unveiling the Gas Phase Kinetic Chemistry in Protoplanetary Disks

Configuration: Longest Baseline = 460.0 m

Proposed rms: 5.0 mJy / 0.122070 MHz (0.166 km/s)

Proposed beam size: 1.0 arcsec

CASA version used for reduction: 4.7.2

QA2 Result: PASS

Total Number of Member SBS in this OUS Group: 1

Comments from Reducer:

This scheduling block was calibrated and imaged using the pipeline version 38366 (C4-R2B) in CASA 4.7.2 (r39732).

The calibration appears reasonable, no additional flagging was required in stage 2 (hifa_flagdata).

The imaging is good, but there is some noise.

The continuum is good, but there is some noise. The PI may choose to do a more careful

l identification.

The continuum is good, but there is some noise. The PI may choose to do a more careful

Self-calibration.

All pipeline flags are good, and a broad mask, the PI may want to do a deeper clean

with careful masking to improve the images.

In script directory,

➤ casa(4.7.2) –pipeline

➤ Execfile('scriptForPI.py')

Request it via Helpdesk

- Only interferometric data (not single-dish data)
- Up to 5 SBs at one time.
- You can download it after 2 weeks.

【How to request】

1. Log in to Helpdesk <https://help.almascience.org/>
2. Create a new ticket by clicking "Submit Helpdesk Ticket"
3. Choose "Archive and Data Retrieval (EA)" as Department
4. Put "Request of calibrated MS" as Subject
5. Choose "Data request (calibrated MS, stale data, calibrator data, or suggestions)" as Sub-category
6. Provide the following information in the box of "Message", then "Submit" the ticket

Message

a. Your affiliation

[example: University of Tokyo] (so that we can check if the institute does not have a network issue with NAOJ etc.)

b. Project Code and SB name of the data

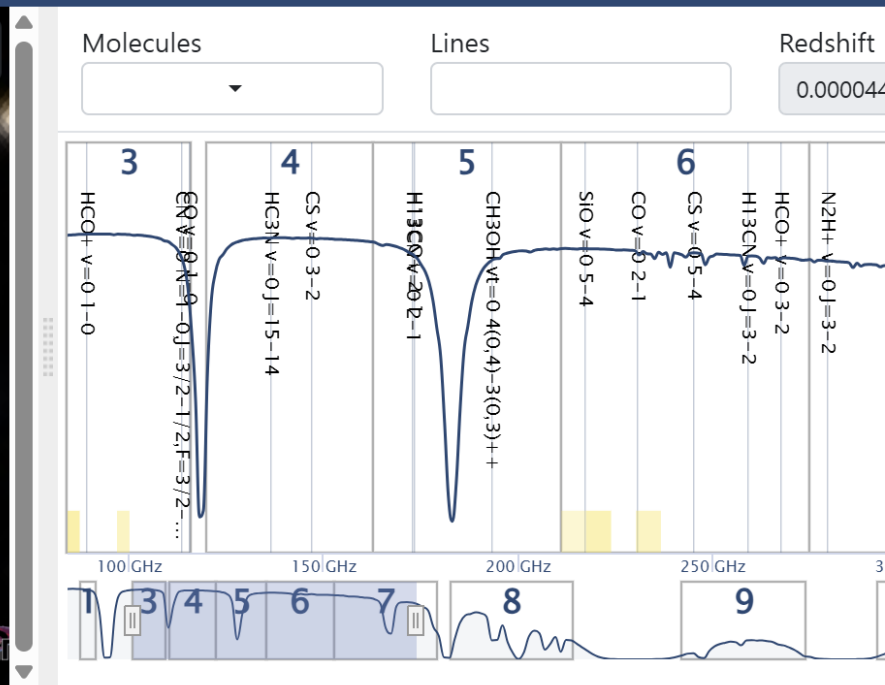
[example: please split Project Code and SB name with a space.

2013.1.00010.S NGC253_a_04_TM1

2013.1.00010.S NGC253_a_07_TM1

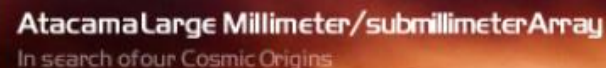
2014.1.00020.S NGC1068_a_06_TM1

]

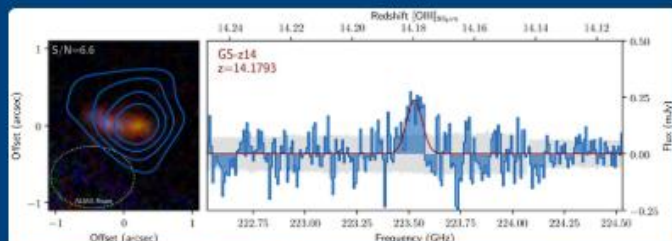


Project code: 2016.1.00229.S

Project code	ALMA source name	FOV	arcsec	Int.Time	s	Min.freq.res.	kHz	L80 BL	m	Obs.date	SB name	Line sens.(10 km/s)	mJy/beam	Group	ous id
2016.1.00229.S															
2016.1.00229.S	tw_hya	16.937		2963.520		244.1314		351.852		2016-11-23	tw_hya_a_07_TM1	0.544			uid://A001/X87d/Xb36
2016.1.00229.S	TW_Hya	25.743		2419.200		122.0798		237.619		2017-04-30	TW_Hya_a_06_TM1	0.539			uid://A001/X87d/Xb3c
2016.1.00229.S	TW_Hya	63.664		2612.736		61.0408		1221.902		2017-08-01	TW_Hya_a_03_TM1	0.480			uid://A001/X87d/Xb30



ALMA Detects [O III] 88 μ m Emission from a $z \sim 14$ Galaxy



(left) ALMA [O III] 88 μ m emission contours overlaid on a JWST NIRCAM image of the galaxy. (right) The ALMA spectrum of the [O III] emission extracted from the $>3\sigma$ emission region in the moment-zero map. A Gaussian fit to the line is also shown.

The sensitivity of ALMA is making it possible to conduct follow-up observations of high-redshift JWST galaxy candidates to (a) confirm their redshifts and (b) derive galaxy properties. In a recently published study led by Sander Schouws, the authors make use of ALMA to search for [O III] 88 μ m line emission from JADES-GS-

**ALMA Announces Observatory Project
for Configurations 8, 9 and 10**
Jul 11, 2025

Cycle 12 Proposal Submission Statistics
Jun 13, 2025

Cycle 12 OT Known Issues
Apr 17, 2025

ALMA DDT Proposals: Results from Cycles 1-10
Mar 24, 2025

More...

ALMA Data Reduction Workshop
NAOJ (26-28 May)
May 21, 2025

ALMA Cycle 12 Proposal preparation meeting (21 March 2005)
Mar 19, 2025

**East Asian ALMA Science Workshop
2025 (22-24 Sept.) & East Asian ALMA
Data Analysis Workshop 2025 (25-26
Sept.)
Mar 19, 2025**

FY2024 ALMA/45m/ASTE Users

[More...](#)

Current configuration: C-9/10

[More...](#)

The ALMA Science Portal is a one-stop source for information and tools aimed at the scientific community as a whole, including proposers, archive researchers and ALMA staff.

Quick Links



🔍 How can we help you today?



Knowledgebase

View all articles >



Submit Helpdesk Ticket

Get in touch for help >



My Tickets

View your tickets >



Face to Face Visit

Arrange a visit >

Submit Helpdesk Ticket

Please complete this form and one of our agents will reply to you by email as soon as possible.

Name *

Seokho Lee

Email

seokholee@kasi.re.kr

CC

Department *

Archive and Data Retrieval (EA)

Face to Face Support (EA)

Data Reduction (EA)

Observing Tool (EA)

Proposal Handling (EA)

[Archive and Data Retrieval \(EA\)](#)

Proposal Change Requests (EA)

Proposal Change Request Review (EA)

Project Triggers

Message *

Please complete this form and one of our agents will reply to you by email as soon as possible.

Name *

Seokho Lee

Email

seokholee@kasi.re.kr

CC

Department *

Archive and Data Retrieval (EA)

Subject *

Request of calibrated MS

Project ID *

Sub-category *

Data Request (calibrated MS, stale data, calibrator data, or suggestions)

Message *

- Your affiliation
- Project Code and SB name of the data (up to 5 SBs)

Add attachment

Weblog

We can find important information and parameters


```
[shlee@pluto member.uid___A001_X87d_
[shlee@pluto qa]$ ls
uid___A001_X87d_Xb3d.weblog.tgz
[shlee@pluto qa]$ tar -xzf uid___A00
[shlee@pluto qa]$ cd pipeline-201705
[shlee@pluto html]$ ls
casa-20170503-212520.log      stage10
casa-20170504-123945.log      stage11
casa_commands.log           stage12
casa_piperestorescript.py     stage13
casa_pipescript.py           stage14
index.html                   stage15
resources                     stage16
session_1                     stage17
stage1                         stage18
```

```
-- project_id/
|-- science_goal.ouss_id/
|-- group.ouss_id/
|-- member.ouss_id/

|-- README (1)
|-- product/ (2)
|-- calibration/ (3)
|-- qa/ (4)
|-- script/ (5)
|-- log/ (6)
|-- raw/ (7)
```

uid_____XXXXXX.weblog.tgz
pipeline-XXXXXX/html (directory)

```
[shlee@pluto html]$ casa --pipeline
```

```
CASA <1>: h_weblog()
```

```
2024-07-03 20:06:57 INFO: Found weblogs at:
t1-1.html
```

```
2024-07-03 20:07:02 INFO: Serving web log on 127.0.0.1 port 30000 (http://127.0.0.1:30000/) ...
```

```
2024-07-03 20:07:02 INFO: Opening http://127.0.0.1:30000/t1-1.html
```

Open Weblog

<https://help.almascience.org/kb/articles/what-is-the-best-way-to-view-the-weblog>

The screenshot shows a web browser window with the URL `help.almascience.org/kb/articles/what-is-the-best-way-to-view-the-weblog`. The page has a blue header with a search bar containing the text "How can we help you today?". Below the header is a breadcrumb trail: "Help Center > Knowledgebase > General ALMA Queries > What is the best way to view the weblog? TOO Search Sci Portal". The main content area has a title "What is the best way to view the weblog?" with a lightbulb icon. Below the title is a small circle with "SW" and the text "Last updated: Jan 25, 2022 by Sarah Wood". To the right of the title are icons for printing and downloading. The main text area contains three paragraphs: "For Mac OS and linux users the recommended internet browser to use for full functionality of the pipeline weblog viewing is Firefox.", "Since 2021 the ALMA pipeline (2021.2.0.128) using CASA 6.2.1-7, weblogs are also viewable with Chrome and Safari browsers. However, due to various browser security options, total functionality of the weblogs may not be available by default.", and "Weblog interface errors can include not opening or loading: linked files, the 'by topic' or 'by task' pages, sub-plots and sub-pages - due to an inability to find the correct reference html page links. There may also be issue to produce all plots, and or the radio direction buttons and side bar links. Note, for older Pipeline datasets using e.g. < CASA 5.6.1, Firefox remains the advised browser to use." followed by "Please take note: sometimes there can be errors when using a browser to open the weblog." On the right side, there is a "Subscribe" button with a dropdown arrow, a "Labels" section with buttons for "error", "firefox", "qa2", and "weblog", and an "Author" section listing "Sarah Wood" and a "Date Created" section listing "Sep 23, 2019".

help.almascience.org/kb/articles/what-is-the-best-way-to-view-the-weblog

How can we help you today?

Help Center > Knowledgebase > General ALMA Queries > What is the best way to view the weblog? TOO Search Sci Portal

What is the best way to view the weblog?

SW

Last updated: Jan 25, 2022 by Sarah Wood

For Mac OS and linux users the recommended internet browser to use for full functionality of the pipeline weblog viewing is Firefox.

Since 2021 the ALMA pipeline (2021.2.0.128) using CASA 6.2.1-7, weblogs are also viewable with Chrome and Safari browsers. However, due to various browser security options, total functionality of the weblogs may not be available by default.

Weblog interface errors can include not opening or loading: linked files, the 'by topic' or 'by task' pages, sub-plots and sub-pages - due to an inability to find the correct reference html page links. There may also be issue to produce all plots, and or the radio direction buttons and side bar links. Note, for older Pipeline datasets using e.g. < CASA 5.6.1, Firefox remains the advised browser to use.

Please take note: sometimes there can be errors when using a browser to open the weblog.

Subscribe

Labels

error firefox qa2 weblog

Author

Sarah Wood

Date Created

Sep 23, 2019

Open weblog

ii) Use a python3 call, external to a CASA session:

Outside of CASA one can also create the http server in which to view a local weblog. From the command line simply type:

```
python3 -m http.server 8080 --bind 127.0.0.1
```

The weblog can then be accessed in a web browser via the URL:

```
http://127.0.0.1:8080/"location_of_PL_weblog"/html/index.html
```

Note, this method requires python3, for which the version delivered with CASA can be used by setting it as an alias or by calling the full path. On MacOS this is found in "/Applications/CASA.app/Contents/MacOS/python3", or on Linux systems
"install_path/casa-6.2.1-7-pipeline-2021.2.0.128/bin/python3"

Observation Overview

Project	uid://A001/X5ac/Xba
Principal Investigator	ebergin
OUS Status Entity id	uid://A001/X87d/Xb3d
Observation Start	2017-04-30 00:27:22 UTC
Observation End	2017-04-30 01:32:42 UTC

Pipeline Summary

Pipeline Version	r39732 (Pipeline-Cycle4-R2-B) (documentation)
CASA Version	4.7.2 r39762
Pipeline Start	2017-05-03 21:25:41 UTC
Execution Duration	1 day, 0:32:00

Observation Summary

Measurement Set	Receivers	Size
Observing Unit Set Status: uid://A001/X87d/Xb3d Scheduling Block ID: uid://A001/X87d/Xb2d		
Session: session_1		
uid__A002_Xbfc9b_X70c.ms	ALMA Band 6 40 2017-04-30 00:27:21 2017-04-30 01:32:41 0:40:32 15.1 m 460.0 m 189.5 m 57.5 GB	
uid__A002_Xbfc9b_X70c_target.ms	ALMA Band 6 40 2017-04-30 00:41:17 2017-04-30 01:31:21 0:40:28 15.1 m 460.0 m 189.5 m 40.0 GB	

When run 'tclean',
If Calibration using < CASA 4.3, then should use the same version
Else If Calibration using >= CASA 4.3, It is OK using recent version,
but recommend that using the same version.

Observation Overview

Project	uid://A001/X5ac/Xba
Principal Investigator	ebergin
OUS Status Entity id	uid://A001/X87d/Xb3d
Observation Start	2017-04-30 00:27:22 UTC
Observation End	2017-04-30 01:32:42 UTC

Pipeline Summary

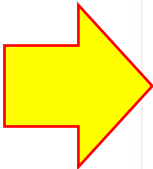
Pipeline Version	r39732 (Pipeline-Cycle4-R2-B) (documentation)
CASA Version	4.7.2 r39762
Pipeline Start	2017-05-03 21:25:41 UTC
Execution Duration	1 day, 0:32:00

Observation Summary

Measurement Set	Receivers	Num Antennas	Time (UTC)			Baseline Length			Size
			Start	End	On Source	Min	Max	RMS	
Observing Unit Set St	A001/X87d/Xb3d Scheduling Block ID: uid://A001/X87d/Xb2d								
Session: session_1									
uid__A002_Xbfcd9b_X70c.ms	ALMA Band 6	40	2017-04-30 00:27:21	2017-04-30 01:32:41	0:40:32	15.1 m	460.0 m	189.5 m	57.5 GB
uid__A002_Xbfcd9b_X70c_target.ms	ALMA Band 6	40	2017-04-30 00:41:17	2017-04-30 01:31:21	0:40:28	15.1 m	460.0 m	189.5 m	40.0 GB



Session: session_1
uid__A002_Xbfcd9b_X70c.ms
uid__A002_Xbfcd9b_X70c_target.ms



Overview of 'uid__A002_Xbfcd9b_X70c.ms'

Observation Execution Time

Start Time	2017-04-30 00:27:21
End Time	2017-04-30 01:32:41
Total Time on Source	0:56:56
Total Time on Science Target	0:40:32

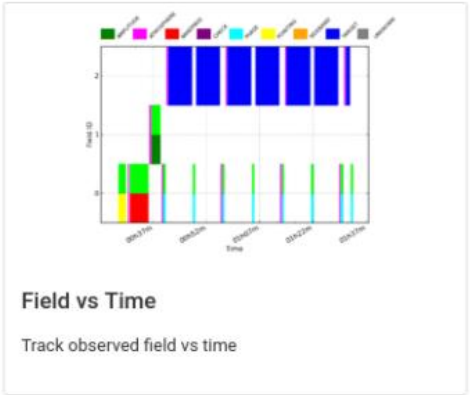
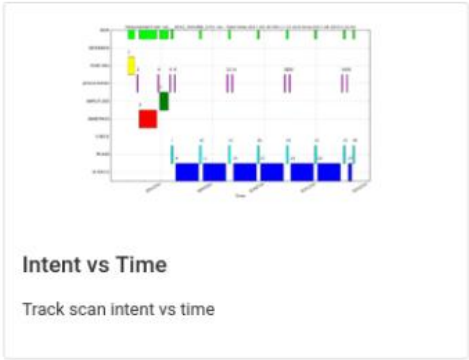
LISTOBS OUTPUT

Spatial Setup

Science Targets	'TW_Hya'
Calibrators	'J1037-2934' and 'J1107-4449'

Antenna Setup

Min Baseline	15.1 m
Max Baseline	460.0 m
Number of Baselines	780
Number of Antennas	40



Spectral Setup

All Bands	'ALMA Band 6' and 'WVR'
Science Bands	'ALMA Band 6'

Sky Setup

Min Elevation	67.28 degrees
Max Elevation	83.36 degrees

```

a ##### Begin Task: listobs #####
a listobs( vis='calibrated_final.ms', selectdata=True, spw='', field='', antenna='', uvrange='', timerange='', correlation='', scan='', intent='', feed='', ar
y #####
+ MeasurementSet Name: /scratch/alma/shlee/alma_summer/TW_Hya/2016.1.00229.S/science_goal.uid___A001_X87d_Xb3b/group.uid___A001_X87d_Xb3c/member.
+ #####
+ Observer: ebergin Project: uid://A001/X5ac/Xba
+ Observation: ALMA
+ Computing scan and subscan properties...
y Data records: 1640000 Total elapsed time = 3065.14 seconds
+ Observed from 30-Apr-2017/00:41:17.1 to 30-Apr-2017/01:32:22.3 (UTC)
y
+ ObservationID = 0 ArrayID = 0
+ Date Timerange (UTC) Scan FldId FieldName nRows SpwIds Average Interval(s) ScanIntent
+ 30-Apr-2017/00:41:17.1 - 00:47:51.8 9 2 TW_Hya 266500 [0,1,2,3,4] [6.05, 6.05, 6.05, 6.05, 6.05] [OBSERVE_TARGET#ON_SOURCE]
+ 00:49:09.4 - 00:55:44.0 11 2 TW_Hya 266500 [0,1,2,3,4] [6.05, 6.05, 6.05, 6.05, 6.05] [OBSERVE_TARGET#ON_SOURCE]
+ 00:58:02.6 - 01:04:37.3 15 2 TW_Hya 266500 [0,1,2,3,4] [6.05, 6.05, 6.05, 6.05, 6.05] [OBSERVE_TARGET#ON_SOURCE]
+ 01:05:53.4 - 01:12:28.1 17 2 TW_Hya 266500 [0,1,2,3,4] [6.05, 6.05, 6.05, 6.05, 6.05] [OBSERVE_TARGET#ON_SOURCE]
+ 01:14:41.0 - 01:21:15.6 21 2 TW_Hya 266500 [0,1,2,3,4] [6.05, 6.05, 6.05, 6.05, 6.05] [OBSERVE_TARGET#ON_SOURCE]
+ 01:22:31.5 - 01:29:06.2 23 2 TW_Hya 266500 [0,1,2,3,4] [6.05, 6.05, 6.05, 6.05, 6.05] [OBSERVE_TARGET#ON_SOURCE]
+ 01:31:21.8 - 01:32:22.3 27 2 TW_Hya 41000 [0,1,2,3,4] [6.05, 6.05, 6.05, 6.05, 6.05] [OBSERVE_TARGET#ON_SOURCE]
y (nRows = Total number of rows per scan)
y Fields: 1
+ ID Code Name RA Decl Epoch SrcId nRows
+ 2 none TW_Hya 11:01:51.812563 -34.42.17.27561 ICRS 2 1640000
y Spectral Windows: (5 unique spectral windows and 1 unique polarization setups)
+ SpwID Name #Chans Frame Ch0 (MHz) ChanWid(kHz) TotBW(kHz) CtrFreq(MHz) BBC Num Corrs
+ 0 X2074619945#ALMA_RB_06#BB_2#SW-01#FULL_RES 128 TOPO 219075.479 -15625.000 2000000.0 218083.2912 2 XX YY
+ 1 X2074619945#ALMA_RB_06#BB_4#SW-01#FULL_RES 128 TOPO 233286.533 15625.000 2000000.0 234278.7202 4 XX YY
+ 2 X2074619945#ALMA_RB_06#BB_1#SW-01#FULL_RES 1920 TOPO 220437.264 -61.035 117187.5 220378.7008 1 XX YY
+ 3 X2074619945#ALMA_RB_06#BB_1#SW-02#FULL_RES 1920 TOPO 219598.946 -61.035 117187.5 219540.3830 1 XX YY
+ 4 X2074619945#ALMA_RB_06#BB_3#SW-01#FULL_RES 3840 TOPO 231082.493 61.035 234375.0 231199.6500 3 XX YY
y Sources: 15
+ ID Name SpwId RestFreq(MHz) SysVel (km/s)
+ 0 J1037-2934 0 218103.1 0
+ 0 J1037-2934 1 234300 0
+ 0 J1037-2934 2 220398.6842 0
+ 0 J1037-2934 3 219560.358 0
+ 0 J1037-2934 4 231220.686 0
+ 1 J1107-4449 2 220398.6842 0
+ 1 J1107-4449 3 219560.358 0
+ 1 J1107-4449 0 218103.1 0
+ 1 J1107-4449 4 231220.686 0
+ 1 J1107-4449 1 234300 0
+ 2 TW_Hya 2 220398.6842 13.4
+ 2 TW_Hya 3 219560.358 13.4
+ 2 TW_Hya 0 218103.1 13.4
+ 2 TW_Hya 4 231220.686 13.4
+ 2 TW_Hya 1 234300 13.4
y Antennas: 40:
+ ID Name Station Diam. Long. Lat. Offset from array center (m) ITRF Geocentric coordinates (m)

```

Listobs in casalogger

Check the field ID and Spw ID

```

=====
MeasurementSet Name: /scratch/alma/shlee/alma_summer/TW_Hya/2016.1.00229.S/science_goal.uid___A001_X87d_Xb3b/group.uid___A001_X87d_Xb3c/member.uid___A001_X87d_Xb3d/calibrated/calibrated.ms MS Version 2
=====
Observer: ebergin Project: uid://A001/X5ac/Xba
Observation: ALMA
Data records: 2455900 Total elapsed time = 3791.95 seconds
Observed from 30-Apr-2017/00:30:00.9 to 30-Apr-2017/01:33:12.8 (UTC)

ObservationID = 0 ArrayID = 0
Date Timerange (UTC) Scan FldId FieldName nRows SpwIds Average Interval(s) ScanIntent
30-Apr-2017/00:30:00.9 - 00:30:17.0 2 0 J1037-2934 34440 [0,1] [0.576, 0.576] [CALIBRATE_ATMOSPHERE#AMBIENT,CALIBRATE_ATMOSPHERE#HOT,CALIBRATE_ATMOSPHERE#OFF_SOURCE,CALIBRATE_WVR#AMBIENT,CALIBRATE_WVR#HOT,CALIBRATE_WVR#OFF_SOURCE]
00:30:37.3 - 00:35:39.7 3 0 J1037-2934 205000 [0,1,2,3,4] [6.05, 6.05, 6.05, 6.05, 6.05, 6.05] [CALIBRATE_BANDPASS#ON_SOURCE,CALIBRATE_WVR#ON_SOURCE]
00:36:04.2 - 00:36:20.3 4 1 J1107-4449 34440 [0,1] [0.576, 0.576] [CALIBRATE_ATMOSPHERE#AMBIENT,CALIBRATE_ATMOSPHERE#HOT,CALIBRATE_ATMOSPHERE#OFF_SOURCE,CALIBRATE_WVR#AMBIENT,CALIBRATE_WVR#HOT,CALIBRATE_WVR#OFF_SOURCE]
00:36:36.3 - 00:39:07.5 5 1 J1107-4449 102500 [0,1,2,3,4] [6.05, 6.05, 6.05, 6.05, 6.05] [CALIBRATE_FLUX#ON_SOURCE,CALIBRATE_WVR#ON_SOURCE]
00:39:30.9 - 00:39:47.0 6 0 J1037-2934 34440 [0,1] [0.576, 0.576] [CALIBRATE_ATMOSPHERE#AMBIENT,CALIBRATE_ATMOSPHERE#HOT,CALIBRATE_ATMOSPHERE#OFF_SOURCE,CALIBRATE_WVR#AMBIENT,CALIBRATE_WVR#HOT,CALIBRATE_WVR#OFF_SOURCE]
00:39:57.4 - 00:40:27.6 7 0 J1037-2934 20500 [0,1,2,3,4] [6.05, 6.05, 6.05, 6.05, 6.05] [CALIBRATE_PHASE#ON_SOURCE,CALIBRATE_WVR#ON_SOURCE]
00:40:49.7 - 00:41:05.8 8 2 TW_Hya 34440 [0,1] [0.576, 0.576] [CALIBRATE_ATMOSPHERE#AMBIENT,CALIBRATE_ATMOSPHERE#HOT,CALIBRATE_ATMOSPHERE#OFF_SOURCE,CALIBRATE_WVR#AMBIENT,CALIBRATE_WVR#HOT,CALIBRATE_WVR#OFF_SOURCE]
00:41:17.1 - 00:47:51.8 9 2 TW_Hya 266500 [0,1,2,3,4] [6.05, 6.05, 6.05, 6.05, 6.05] [OBSERVE_TARGET#ON_SOURCE]
00:48:13.7 - 00:48:43.9 10 0 J1037-2934 20500 [0,1,2,3,4] [6.05, 6.05, 6.05, 6.05, 6.05] [CALIBRATE_PHASE#ON_SOURCE,CALIBRATE_WVR#ON_SOURCE]
00:49:09.4 - 00:55:44.0 11 2 TW_Hya 266500 [0,1,2,3,4] [6.05, 6.05, 6.05, 6.05, 6.05] [OBSERVE_TARGET#ON_SOURCE]
00:56:06.1 - 00:56:22.3 12 0 J1037-2934 34440 [0,1] [0.576, 0.576] [CALIBRATE_ATMOSPHERE#AMBIENT,CALIBRATE_ATMOSPHERE#HOT,CALIBRATE_ATMOSPHERE#OFF_SOURCE,CALIBRATE_WVR#AMBIENT,CALIBRATE_WVR#HOT,CALIBRATE_WVR#OFF_SOURCE]
00:56:37.1 - 00:57:07.3 13 0 J1037-2934 20500 [0,1,2,3,4] [6.05, 6.05, 6.05, 6.05, 6.05] [CALIBRATE_PHASE#ON_SOURCE,CALIBRATE_WVR#ON_SOURCE]
00:57:32.3 - 00:57:48.4 14 2 TW_Hya 34440 [0,1] [0.576, 0.576] [CALIBRATE_ATMOSPHERE#AMBIENT,CALIBRATE_ATMOSPHERE#HOT,CALIBRATE_ATMOSPHERE#OFF_SOURCE,CALIBRATE_WVR#AMBIENT,CALIBRATE_WVR#HOT,CALIBRATE_WVR#OFF_SOURCE]
00:58:02.6 - 01:04:37.3 15 2 TW_Hya 266500 [0,1,2,3,4] [6.05, 6.05, 6.05, 6.05, 6.05] [OBSERVE_TARGET#ON_SOURCE]
01:04:59.6 - 01:05:29.8 16 0 J1037-2934 20500 [0,1,2,3,4] [6.05, 6.05, 6.05, 6.05, 6.05] [CALIBRATE_PHASE#ON_SOURCE,CALIBRATE_WVR#ON_SOURCE]
01:05:53.4 - 01:12:28.1 17 2 TW_Hya 266500 [0,1,2,3,4] [6.05, 6.05, 6.05, 6.05, 6.05] [OBSERVE_TARGET#ON_SOURCE]
01:12:50.1 - 01:13:06.2 18 0 J1037-2934 34440 [0,1] [0.576, 0.576] [CALIBRATE_ATMOSPHERE#AMBIENT,CALIBRATE_ATMOSPHERE#HOT,CALIBRATE_ATMOSPHERE#OFF_SOURCE,CALIBRATE_WVR#AMBIENT,CALIBRATE_WVR#HOT,CALIBRATE_WVR#OFF_SOURCE]
01:13:21.1 - 01:13:51.3 19 0 J1037-2934 20500 [0,1,2,3,4] [6.05, 6.05, 6.05, 6.05, 6.05] [CALIBRATE_PHASE#ON_SOURCE,CALIBRATE_WVR#ON_SOURCE]
01:14:14.7 - 01:14:30.8 20 2 TW_Hya 34440 [0,1] [0.576, 0.576] [CALIBRATE_ATMOSPHERE#AMBIENT,CALIBRATE_ATMOSPHERE#HOT,CALIBRATE_ATMOSPHERE#OFF_SOURCE,CALIBRATE_WVR#AMBIENT,CALIBRATE_WVR#HOT,CALIBRATE_WVR#OFF_SOURCE]
01:14:41.0 - 01:21:15.6 21 2 TW_Hya 266500 [0,1,2,3,4] [6.05, 6.05, 6.05, 6.05, 6.05] [OBSERVE_TARGET#ON_SOURCE]
01:21:37.1 - 01:22:07.3 22 0 J1037-2934 20500 [0,1,2,3,4] [6.05, 6.05, 6.05, 6.05, 6.05] [CALIBRATE_PHASE#ON_SOURCE,CALIBRATE_WVR#ON_SOURCE]
01:22:31.5 - 01:29:06.2 23 2 TW_Hya 266500 [0,1,2,3,4] [6.05, 6.05, 6.05, 6.05, 6.05] [OBSERVE_TARGET#ON_SOURCE]
01:29:27.3 - 01:29:43.4 24 0 J1037-2934 34440 [0,1] [0.576, 0.576] [CALIBRATE_ATMOSPHERE#AMBIENT,CALIBRATE_ATMOSPHERE#HOT,CALIBRATE_ATMOSPHERE#OFF_SOURCE,CALIBRATE_WVR#AMBIENT,CALIBRATE_WVR#HOT,CALIBRATE_WVR#OFF_SOURCE]
01:29:57.7 - 01:30:27.9 25 0 J1037-2934 20500 [0,1,2,3,4] [6.05, 6.05, 6.05, 6.05, 6.05] [CALIBRATE_PHASE#ON_SOURCE,CALIBRATE_WVR#ON_SOURCE]
01:30:51.2 - 01:31:07.3 26 2 TW_Hya 34440 [0,1] [0.576, 0.576] [CALIBRATE_ATMOSPHERE#AMBIENT,CALIBRATE_ATMOSPHERE#HOT,CALIBRATE_ATMOSPHERE#OFF_SOURCE,CALIBRATE_WVR#AMBIENT,CALIBRATE_WVR#HOT,CALIBRATE_WVR#OFF_SOURCE]
01:31:21.8 - 01:32:22.3 27 2 TW_Hya 41000 [0,1,2,3,4] [6.05, 6.05, 6.05, 6.05, 6.05] [OBSERVE_TARGET#ON_SOURCE]
01:32:42.6 - 01:33:12.8 28 0 J1037-2934 20500 [0,1,2,3,4] [6.05, 6.05, 6.05, 6.05, 6.05] [CALIBRATE_PHASE#ON_SOURCE,CALIBRATE_WVR#ON_SOURCE]
(nRows = Total number of rows per scan)

Fields: 3
ID Code Name RA Decl Epoch SrcId nRows
0 none J1037-2934 10:37:16.080000 -29.34:02.81300 ICRS 0 541200
1 none J1107-4449 11:07:08.694000 -44.49:07.61800 ICRS 1 136940
2 none TW_Hya 11:01:51.812563 -34.42:17.27561 ICRS 2 1777760

Spectral Windows: (5 unique spectral windows and 1 unique polarization setups)
SpwID Name #Chans Frame Ch0(MHz) ChanWid(kHz) TotBW(kHz) CtrFreq(MHz) BBC Num Corrs
0 X2074619945#ALMA_RB_06#BB_2#SW-01#FULL_RES 128 TOPO 219075.479 -15625.000 2000000.0 218083.2912 2 XX YY
1 X2074619945#ALMA_RB_06#BB_4#SW-01#FULL_RES 128 TOPO 233286.533 15625.000 2000000.0 234278.7202 4 XX YY
2 X2074619945#ALMA_RB_06#BB_1#SW-01#FULL_RES 1920 TOPO 228437.264 -61.035 117187.5 220378.7008 1 XX YY
3 X2074619945#ALMA_RB_06#BB_1#SW-02#FULL_RES 1920 TOPO 219598.946 -61.035 117187.5 219540.3830 1 XX YY
4 X2074619945#ALMA_RB_06#BB_3#SW-01#FULL_RES 3840 TOPO 231082.493 61.035 234375.0 231199.6500 3 XX YY

Sources: 15
ID Name SpwID RestFreq(MHz) SysVel(km/s)
0 J1037-2934 0 218103.1 0
0 J1037-2934 1 234300 0
0 J1037-2934 2 220398.6842 0
0 J1037-2934 3 219560.358 0
0 J1037-2934 4 231220.686 0
1 J1107-4449 2 220398.6842 0
1 J1107-4449 3 219560.358 0
1 J1107-4449 0 218103.1 0
1 J1107-4449 4 231220.686 0
1 J1107-4449 1 234300 0
2 TW_Hya 2 220398.6842 13.4
2 TW_Hya 3 219560.358 13.4
2 TW_Hya 0 2181
```


listobs ('calibrated_final.ms')

```
*****
MeasurementSet Name: /scratch/alma/shlee/alma_summer/TW_Hya/2016.1.00229.5/science_goal.uid__A001_X87d_Xb3b/group.uid__A001_X87d_Xb3c/member.uid__A001_X87d_Xb3d/calibrated/calibrated_final.ms MS Version 2
*****
Observer: ebergin Project: uid://A001/X5ac/Xba
Observation: ALMA
Data records: 1640000 Total elapsed time = 3065.14 seconds
Observed from 30-Apr-2017/00:41:17.1 to 30-Apr-2017/01:32:22.3 (UTC)

ObservationID = 0 ArrayID = 0
Date Timerange (UTC) Scan FldId FieldName nRows SpwIds Average Interval(s) ScanIntent
30-Apr-2017/00:41:17.1 - 00:47:51.8 9 2 TW_Hya 266500 [0,1,2,3,4] [6.05, 6.05, 6.05, 6.05, 6.05] [OBSERVE_TARGET#ON_SOURCE]
00:49:09.4 - 00:55:44.0 11 2 TW_Hya 266500 [0,1,2,3,4] [6.05, 6.05, 6.05, 6.05, 6.05] [OBSERVE_TARGET#ON_SOURCE]
00:58:02.6 - 01:04:37.3 15 2 TW_Hya 266500 [0,1,2,3,4] [6.05, 6.05, 6.05, 6.05, 6.05] [OBSERVE_TARGET#ON_SOURCE]
01:05:53.4 - 01:12:28.1 17 2 TW_Hya 266500 [0,1,2,3,4] [6.05, 6.05, 6.05, 6.05, 6.05] [OBSERVE_TARGET#ON_SOURCE]
01:14:41.0 - 01:21:15.6 21 2 TW_Hya 266500 [0,1,2,3,4] [6.05, 6.05, 6.05, 6.05, 6.05] [OBSERVE_TARGET#ON_SOURCE]
01:22:31.5 - 01:29:06.2 23 2 TW_Hya 266500 [0,1,2,3,4] [6.05, 6.05, 6.05, 6.05, 6.05] [OBSERVE_TARGET#ON_SOURCE]
01:31:21.8 - 01:32:22.3 27 2 TW_Hya 41000 [0,1,2,3,4] [6.05, 6.05, 6.05, 6.05, 6.05] [OBSERVE_TARGET#ON_SOURCE]
(nRows = Total number of rows per scan)

Fields: 1
ID Code Name RA Decl Epoch SrcId nRows
2 none TW_Hya 11:01:51.812563 -34.42.17.27561 ICRS 2 1640000

Spectral Windows: (5 unique spectral windows and 1 unique polarization setups)
SpwID Name #Chans Frame Ch0 (MHz) ChanWid (kHz) TotBW (kHz) CtrFreq (MHz) BBC Num Corrs
0 X2074619945#ALMA_RB_06#BB_2#SW-01#FULL_RES 128 TOP0 219075.479 -15625.000 2000000.0 218083.2912 2 XX YY
1 X2074619945#ALMA_RB_06#BB_4#SW-01#FULL_RES 128 TOP0 233286.533 15625.000 2000000.0 234278.7202 4 XX YY
2 X2074619945#ALMA_RB_06#BB_1#SW-01#FULL_RES 1920 TOP0 220437.264 -61.035 117187.5 220378.7008 1 XX YY
3 X2074619945#ALMA_RB_06#BB_1#SW-02#FULL_RES 1920 TOP0 219598.946 -61.035 117187.5 219540.3830 1 XX YY
4 X2074619945#ALMA_RB_06#BB_3#SW-01#FULL_RES 3840 TOP0 231082.493 61.035 234375.0 231199.6500 3 XX YY

Sources: 15
ID Name SpwID RestFreq (MHz) SysVel (km/s)
0 J1037-2934 0 218103.1 0
0 J1037-2934 1 234300 0
0 J1037-2934 2 220398.6842 0
0 J1037-2934 3 219560.358 0
0 J1037-2934 4 231220.686 0
1 J1107-4449 2 220398.6842 0
1 J1107-4449 3 219560.358 0
1 J1107-4449 0 218103.1 0
1 J1107-4449 4 231220.686 0
1 J1107-4449 1 234300 0
2 TW_Hya 2 220398.6842 13.4
2 TW_Hya 3 219560.358 13.4
2 TW_Hya 0 218103.1 13.4
2 TW_Hya 4 231220.686 13.4
2 TW_Hya 1 234300 13.4

Antennas: 40:
ID Name Station Diam. Long. Lat. Offset from array center (m) ITRF Geocentric coordinates (m)
East North Elevation x y z
0 DA41 A004 12.0 m -067.45.15.9 -22.53.28.0 52.6608 -704.4170 21.7721 2225094.796411 -5440052.421403 -2481687.277071
1 DA42 A064 12.0 m -067.45.14.7 -22.53.31.4 85.6572 -808.0277 21.5201 2225109.989466 -5440002.411752 -2481782.629929
2 DA43 A005 12.0 m -067.45.14.8 -22.53.28.7 83.3310 -725.0764 21.7245 2225120.123607 -5440033.332230 -2481706.291028
3 DA46 A034 12.0 m -067.45.18.4 -22.53.27.8 -18.8313 -698.8163 22.1778 2225029.594375 -5440081.847514 -2481682.275280
```



Session: session_1

uid__A002_Xbfcd9b_X70c.ms

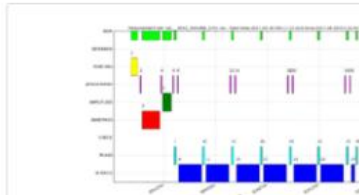
uid__A002_Xbfcd9b_X70c_target.ms

Overview of 'uid__A002_Xbfcd9b_X70c.ms'

Observation Execution Time

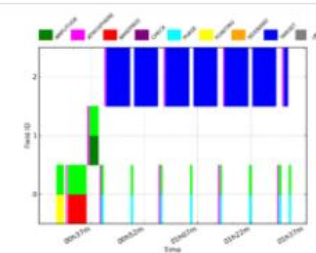
Start Time	2017-04-30 00:27:21
End Time	2017-04-30 01:32:41
Total Time on Source	0:56:56
Total Time on Science Target	0:40:32

LISTOBS OUTPUT



Intent vs Time

Track scan intent vs time



Field vs Time

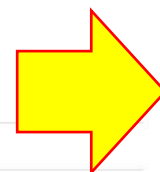
Track observed field vs time

Spatial Setup

Science Targets	'TW_Hya'
Calibrators	'J1037-2934' and 'J1107-4449'

Antenna Setup

Min Baseline	15.1 m
Max Baseline	460.0 m
Number of Baselines	780
Number of Antennas	40



Spectral Setup

All Bands	'ALMA Band 6' and 'WVR'
Science Bands	'ALMA Band 6'

Sky Setup

Min Elevation	67.28 degrees
Max Elevation	83.36 degrees

All Windows

ID	Frequency (TOPO)			Bandwidth (TOPO)	Channels (TOPO)			Correlator Axis	Band	Intents
	Start	Centre	End		Number	Frequency Width	Velocity Width			
0	220.538 GHz	221.538 GHz	222.538 GHz	2.000 GHz	1	2.000 GHz	2706.465 km/s	XX, YY	ALMA Band 6	POINTING, WVR
1	222.538 GHz	223.538 GHz	224.538 GHz	2.000 GHz	1	2.000 GHz	2682.251 km/s	XX, YY	ALMA Band 6	POINTING, WVR
2	236.538 GHz	237.538 GHz	238.538 GHz	2.000 GHz	1	2.000 GHz	2524.164 km/s	XX, YY	ALMA Band 6	POINTING, WVR
3	238.538 GHz	239.538 GHz	240.538 GHz	2.000 GHz	1	2.000 GHz	2503.089 km/s	XX, YY	ALMA Band 6	POINTING, WVR
4	183.800 GHz	187.550 GHz	191.300 GHz	7.500 GHz	4	1.500 GHz	2397.700 km/s	XX	WVR	AMPLITUDE, BANDPASS, PHASE, POINTING, TARGET, WVR
5	220.538 GHz	221.538 GHz	222.538 GHz	2.000 GHz	128	15.625 MHz	21.144 km/s	XX, YY	ALMA Band 6	POINTING, WVR
6	220.624 GHz	221.515 GHz	222.405 GHz	1.781 GHz	1	1.781 GHz	2410.701 km/s	XX, YY	ALMA Band 6	POINTING, WVR
7	222.538 GHz	223.538 GHz	224.538 GHz	2.000 GHz	128	15.625 MHz	20.955 km/s	XX, YY	ALMA Band 6	POINTING, WVR
8	222.624 GHz	223.515 GHz	224.405 GHz	1.781 GHz	1	1.781 GHz	2410.701 km/s	XX, YY	ALMA Band 6	POINTING, WVR
9	236.538 GHz	237.538 GHz	238.538 GHz	2.000 GHz	128	15.625 MHz	20.955 km/s	XX, YY	ALMA Band 6	POINTING, WVR
10	236.624 GHz	237.515 GHz	238.405 GHz	1.781 GHz	1	1.781 GHz	2410.701 km/s	XX, YY	ALMA Band 6	POINTING, WVR
11	238.538 GHz	239.538 GHz	240.538 GHz	2.000 GHz	128	15.625 MHz	19.555 km/s	XX, YY	ALMA Band 6	POINTING, WVR
12	238.624 GHz	239.515 GHz	240.405 GHz	1.781 GHz	1	1.781 GHz	2229.532 km/s	XX, YY	ALMA Band 6	POINTING, WVR
13	218.960 GHz	219.960 GHz	220.960 GHz	2.000 GHz	1	2.000 GHz	2725.887 km/s	XX, YY	ALMA Band 6	AMPLITUDE, ATMOSPHERE, BANDPASS, PHASE, TARGET, WVR
14	217.083 GHz	218.083 GHz	219.083 GHz	2.000 GHz	1	2.000 GHz	2749.339 km/s	XX, YY	ALMA Band 6	AMPLITUDE, ATMOSPHERE, BANDPASS, PHASE, TARGET, WVR
15	230.959 GHz	231.959 GHz	232.959 GHz	2.000 GHz	1	2.000 GHz	2584.875 km/s	XX, YY	ALMA Band 6	AMPLITUDE, ATMOSPHERE, BANDPASS, PHASE, TARGET, WVR
16	233.279 GHz	234.279 GHz	235.279 GHz	2.000 GHz	1	2.000 GHz	2559.280 km/s	XX, YY	ALMA Band 6	AMPLITUDE, ATMOSPHERE, BANDPASS, PHASE, TARGET, WVR
17	218.960 GHz	219.960 GHz	220.960 GHz	2.000 GHz	128	15.625 MHz	21.296 km/s	XX, YY	ALMA Band 6	ATMOSPHERE
18	219.014 GHz	219.952 GHz	220.889 GHz	1.875 GHz	1	1.875 GHz	2555.610 km/s	XX, YY	ALMA Band 6	ATMOSPHERE
19	217.083 GHz	218.083 GHz	219.083 GHz	2.000 GHz	128	15.625 MHz	21.479 km/s	XX, YY	ALMA Band 6	AMPLITUDE, ATMOSPHERE, BANDPASS, PHASE, TARGET, WVR
20	217.169 GHz	218.068 GHz	218.966 GHz	1.797 GHz	1	1.797 GHz	2470.286 km/s	XX, YY	ALMA Band 6	AMPLITUDE, ATMOSPHERE, BANDPASS, PHASE, TARGET, WVR
21	230.959 GHz	231.959 GHz	232.959 GHz	2.000 GHz	128	15.625 MHz	20.194 km/s	XX, YY	ALMA Band 6	ATMOSPHERE
22	231.014 GHz	231.951 GHz	232.889 GHz	1.875 GHz	1	1.875 GHz	2423.402 km/s	XX, YY	ALMA Band 6	ATMOSPHERE
23	233.279 GHz	234.279 GHz	235.279 GHz	2.000 GHz	128	15.625 MHz	19.994 km/s	XX, YY	ALMA Band 6	AMPLITUDE, ATMOSPHERE, BANDPASS, PHASE, TARGET, WVR
24	233.365 GHz	234.263 GHz	235.162 GHz	1.797 GHz	1	1.797 GHz	2299.507 km/s	XX, YY	ALMA Band 6	AMPLITUDE, ATMOSPHERE, BANDPASS, PHASE, TARGET, WVR
25	220.320 GHz	220.379 GHz	220.437 GHz	117.188 MHz	1920	61.035 kHz	83.029 m/s	XX, YY	ALMA Band 6	AMPLITUDE, BANDPASS, PHASE, TARGET, WVR
26	220.320 GHz	220.379 GHz	220.437 GHz	117.188 MHz	1	117.188 MHz	159.416 km/s	XX, YY	ALMA Band 6	AMPLITUDE, BANDPASS, PHASE, TARGET, WVR
27	219.482 GHz	219.540 GHz	219.599 GHz	117.188 MHz	1920	61.035 kHz	83.346 m/s	XX, YY	ALMA Band 6	AMPLITUDE, BANDPASS, PHASE, TARGET, WVR
28	219.482 GHz	219.540 GHz	219.599 GHz	117.188 MHz	1	117.188 MHz	160.025 km/s	XX, YY	ALMA Band 6	AMPLITUDE, BANDPASS, PHASE, TARGET, WVR
29	231.082 GHz	231.200 GHz	231.317 GHz	234.375 MHz	3840	61.035 kHz	79.143 m/s	XX, YY	ALMA Band 6	AMPLITUDE, BANDPASS, PHASE, TARGET, WVR
30	231.082 GHz	231.200 GHz	231.317 GHz	234.375 MHz	1	234.375 MHz	303.910 km/s	XX, YY	ALMA Band 6	AMPLITUDE, BANDPASS, PHASE, TARGET, WVR

Science Windows

All Windows

Science Windows

ID	Frequency (TOPO)			Bandwidth (TOPO)	Channels (TOPO)			Correlator Axis	Band
	Start	Centre	End		Number	Frequency Width	Velocity Width		
19	217.083 GHz	218.083 GHz	219.083 GHz	2.000 GHz	128	15.625 MHz	21.479 km/s	XX, YY	ALMA Band 6
23	233.279 GHz	234.279 GHz	235.279 GHz	2.000 GHz	128	15.625 MHz	19.994 km/s	XX, YY	ALMA Band 6
25	220.320 GHz	220.379 GHz	220.437 GHz	117.188 MHz	1920	61.035 kHz	83.029 m/s	XX, YY	ALMA Band 6
27	219.482 GHz	219.540 GHz	219.599 GHz	117.188 MHz	1920	61.035 kHz	83.346 m/s	XX, YY	ALMA Band 6
29	231.082 GHz	231.200 GHz	231.317 GHz	234.375 MHz	3840	61.035 kHz	79.143 m/s	XX, YY	ALMA Band 6

Spectral Windows with Science Intent in uid___A002_Xbfcd9b_X70c.ms



Session: session_1

uid___A002_Xe1f219_X78a6.ms

Spectral Setup Details

BACK

Science Windows

All Windows

Science Windows

Real ID	Virtual ID	Name	Type	Frequency (TOPO)			Bandwidth (TOPO)	Transitions	Channels (TOPO)			Correlator Axis	Band	Band Type
				Start	Centre	End			Number	Frequency Width	Velocity Width			
23	23	X176064364#ALMA_RB_07#BB_4#SW-01	TDM	302.471 GHz	303.471 GHz	304.471 GHz	2.000 GHz	ContForCal(ID=0)	128	15.625 MHz	15.436 km/s	XX, YY	ALMA Band 7	TSB
25	25	X176064364#ALMA_RB_07#BB_1#SW-01	FDM	316.665 GHz	316.782 GHz	316.899 GHz	234.375 MHz	D2O_1(1,0)-1(0,1)(ID=4104568)	1920	122.070 kHz	115.523 m/s	XX, YY	ALMA Band 7	TSB
27	27	X176064364#ALMA_RB_07#BB_2#SW-01	FDM	315.831 GHz	316.066 GHz	316.300 GHz	468.750 MHz	13CH3OH_v_t=0_10(-1,10)-9(0,9)(ID=575176), 13CH3OH_v_t=1_4(1,4)-5(2,3)_++(ID=3764462)	960	488.281 kHz	463.141 m/s	XX, YY	ALMA Band 7	TSB
29	29	X176064364#ALMA_RB_07#BB_2#SW-02	FDM	315.767 GHz	316.001 GHz	316.235 GHz	468.750 MHz	13CH3OH_v_t=1_4(1,4)-5(2,3)_++(ID=3764462), 13CH3OH_v_t=0_10(-1,10)-9(0,9)(ID=575176)	960	488.281 kHz	463.236 m/s	XX, YY	ALMA Band 7	TSB
31	31	X176064364#ALMA_RB_07#BB_3#SW-01	FDM	301.505 GHz	301.739 GHz	301.973 GHz	468.750 MHz	13CH3OH_v_t=0_8_(2,6)-7_-(-2,6)(ID=575128)	1920	244.141 kHz	242.566 m/s	XX, YY	ALMA Band 7	TSB

Spectral Windows with Science Intent in uid___A002_Xe1f219_X78a6.ms

You can find which spw targeted line is in.

Task Summaries

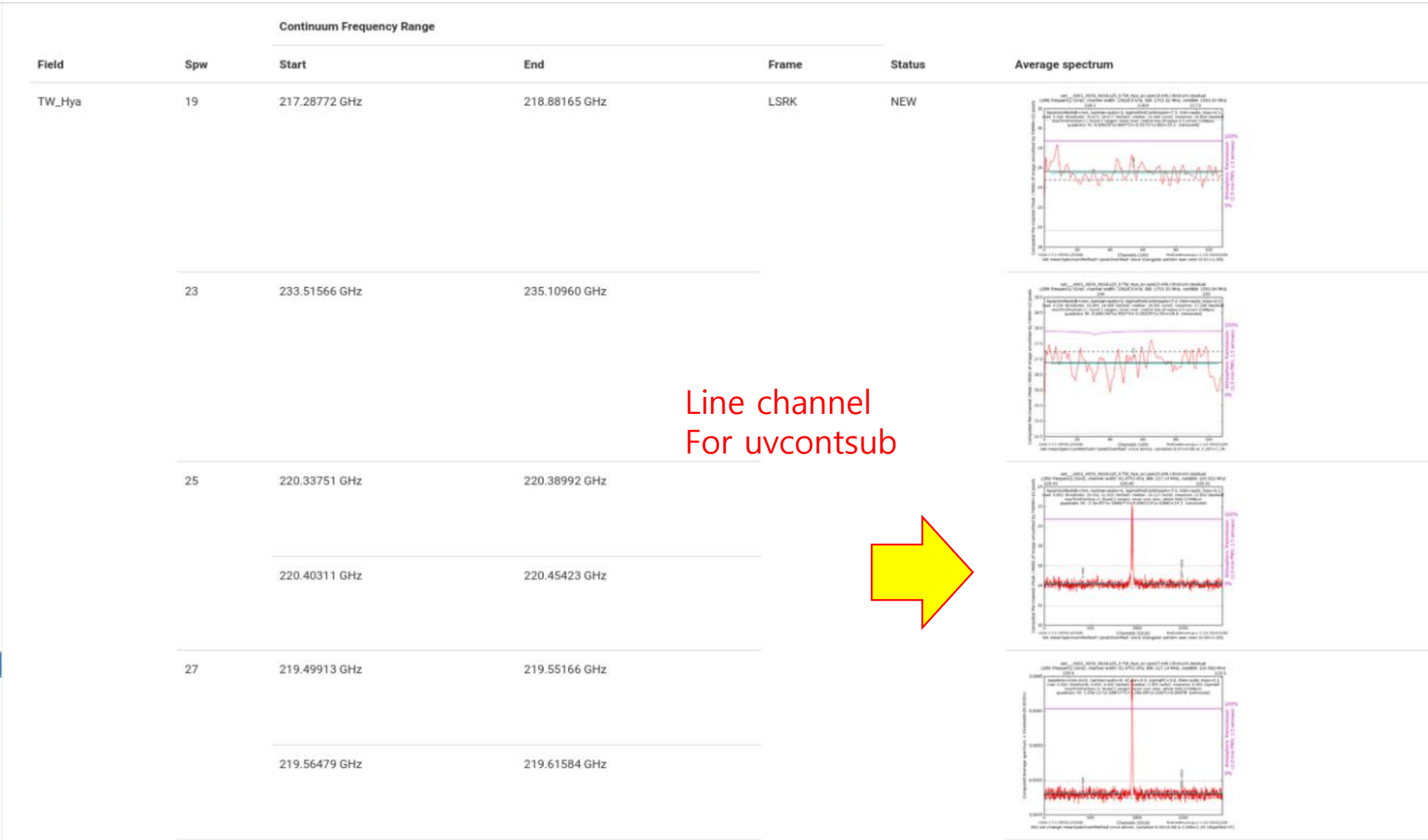
Task	QA Score
1. hifa_importdata : Register measurement sets with the pipeline	1.00
2. hifa_flagdata : ALMA deterministic flagging	1.00
3. hifa_fluxcalflag : Flag spectral features in solar system flux calibrators	1.00
4. hif_newflagphase : Flag channels in raw data	1.00
5. hif_refant : Select reference antennas	1.00
6. hifa_tsyscal : Calculate Tsys calibration	1.00
7. hifa_tsysflag : Flag Tsys calibration	0.99
8. hifa_antpos : Correct for antenna position offsets	0.90
9. hifa_uvvgcalflag : Calculate and flag WVR calibration	1.00
10. hif_lowgainflag : Flag antennas with low gain	1.00
11. hif_gainflag : Flag antennas with gain outliers	1.00
12. hif_setij : Set calibrator model visibilities	1.00
13. hifa_bandpass : Phase-up bandpass calibration	1.00
14. hifa_spwphaseup : Spw phase offsets calibration	1.00
15. hifa_gfluxscale : Transfer fluxscale from amplitude calibrator	1.00
16. hifa_timegaincal : Gain calibration	1.00
17. hif_applycal : Apply calibrations from context	1.00
18. hif_makeimlist : Set-up image parameters for calibrator imaging	1.00
19. hif_makeimages : Make calibrator images	1.00
20. hif_checkproductsize : Check product size	1.00
21. hif_exportdata : Prepare pipeline data products for export	1.00
22. hif_mstransform : Create science target MS	1.00
23. hifa_flagtargets : ALMA Target flagging	1.00
24. hif_makeimlist : Set-up image parameters for target per-spw continuum imaging	1.00
25. hif_fndcont : Detect continuum frequency ranges	1.00
26. hif_uvcontfit : UV continuum fitting	1.00
27. hif_uvcontsub : UV continuum subtraction	1.00
28. hif_makeimages : Make target per-spw continuum images	1.00
29. hif_makeimlist : Set-up image parameters for target aggregate continuum imaging	1.00
30. hif_makeimages : Make target aggregate continuum images	1.00
31. hif_makeimlist : Set-up image parameters for target cube imaging	1.00
32. hif_makeimages : Make target cubes	1.00

Calibration process

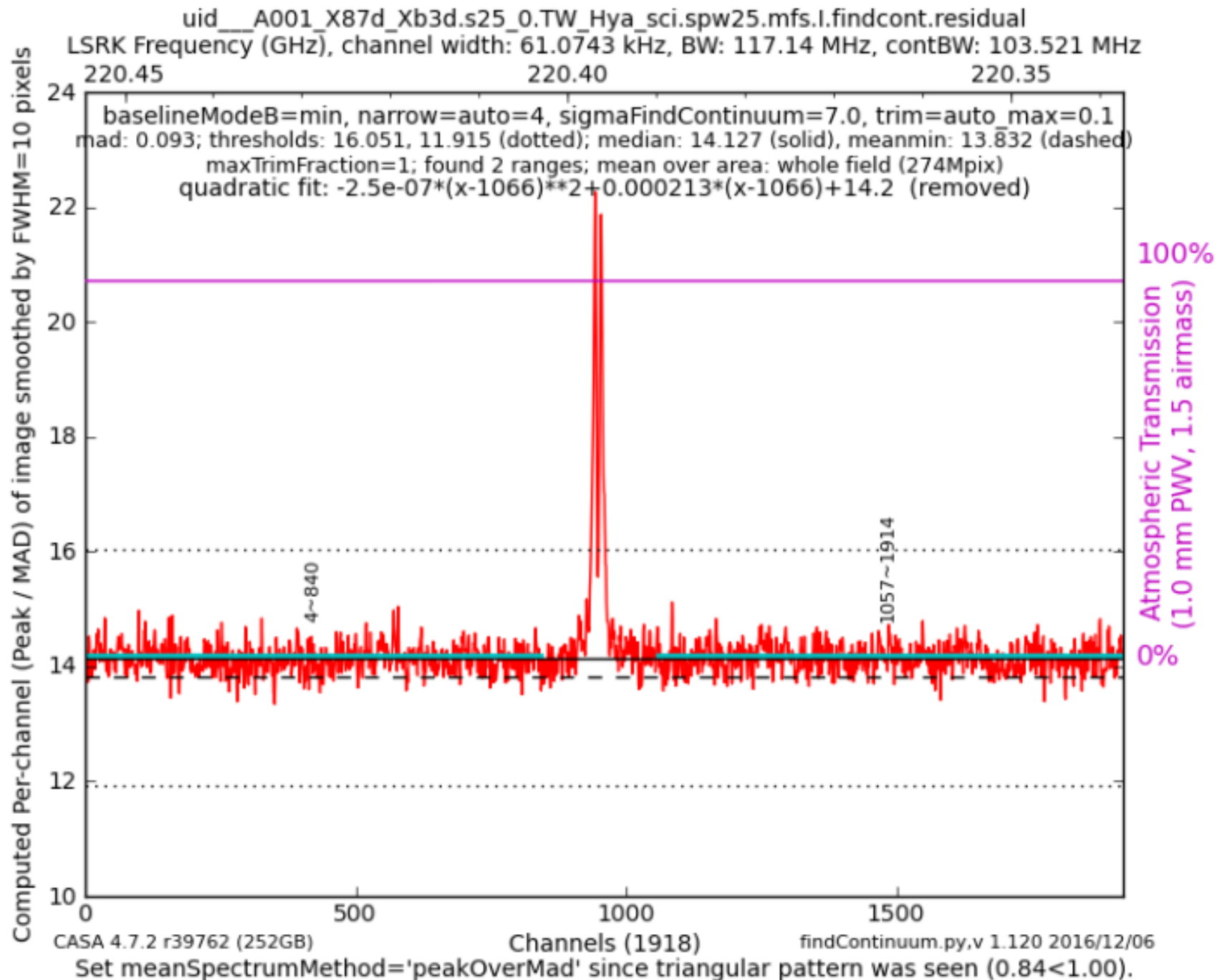
Imagig process

CASA logs and scripts
• View, view in new tab or download casa-20170303-212320.log (8.5 MB) • View, view in new tab or download casa-20170304-123945.log (5.4 MB) • View, view in new tab or download casa_commands.log (93.9 KB) • View, view in new tab or download casa_pipescript.py (1.5 KB) • View, view in new tab or download casa_piperestorescript.py (163 bytes)

- Tasks in execution order
1. hifa_importdata
 2. hifa_flagdata
 3. hifa_fluxcalflag
 4. hif_rawflagchans
 5. hif_refant
 6. hifa_tsyscal
 7. hifa_tsysflag
 8. hifa_antpos
 9. hifa_wvr calflag
 10. hif_lowgainflag
 11. hif_gainflag
 12. hif_settj
 13. hifa_bandpass
 14. hifa_spwphaseup
 15. hifa_gfluxscale
 16. hifa_timegaincal
 17. hif_applycal
 18. hif_makeimlist
 19. hif_makeimages
 20. hif_checkproductsizes
 21. hif_exportdata
 22. hif_mstransform
 23. hifa_flagtargets
 24. hif_makeimlist
 25. hif_findcont
 26. hif_uvcontfit
 27. hif_uvcontsub
 28. hif_makeimages
 29. hif_makeimlist
 30. hif_makeimages
 31. hif_makeimlist



Line channels
for uvcontsub

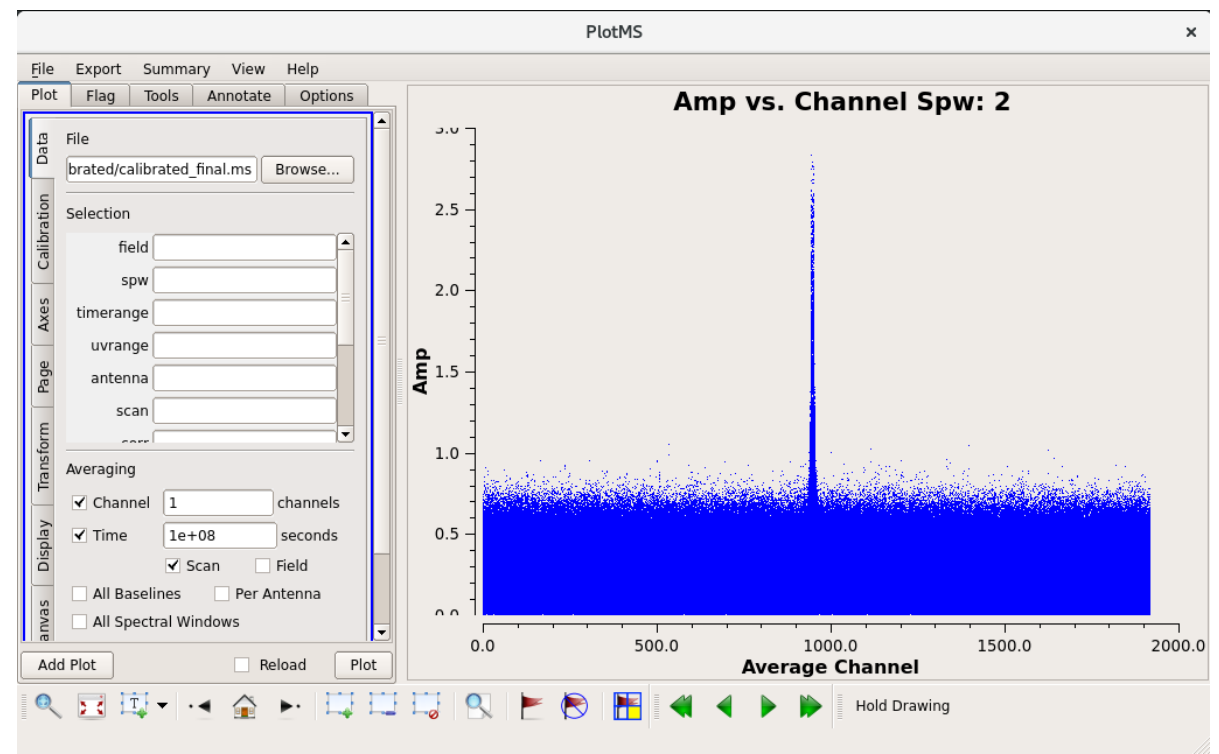
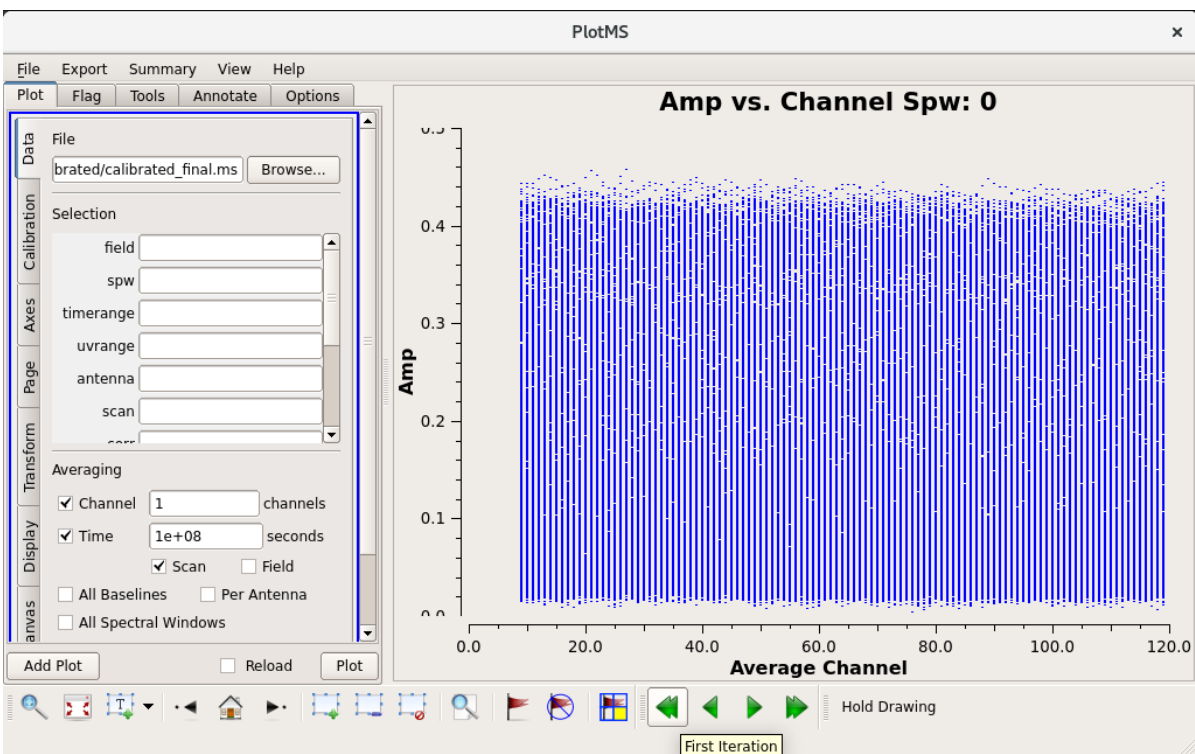


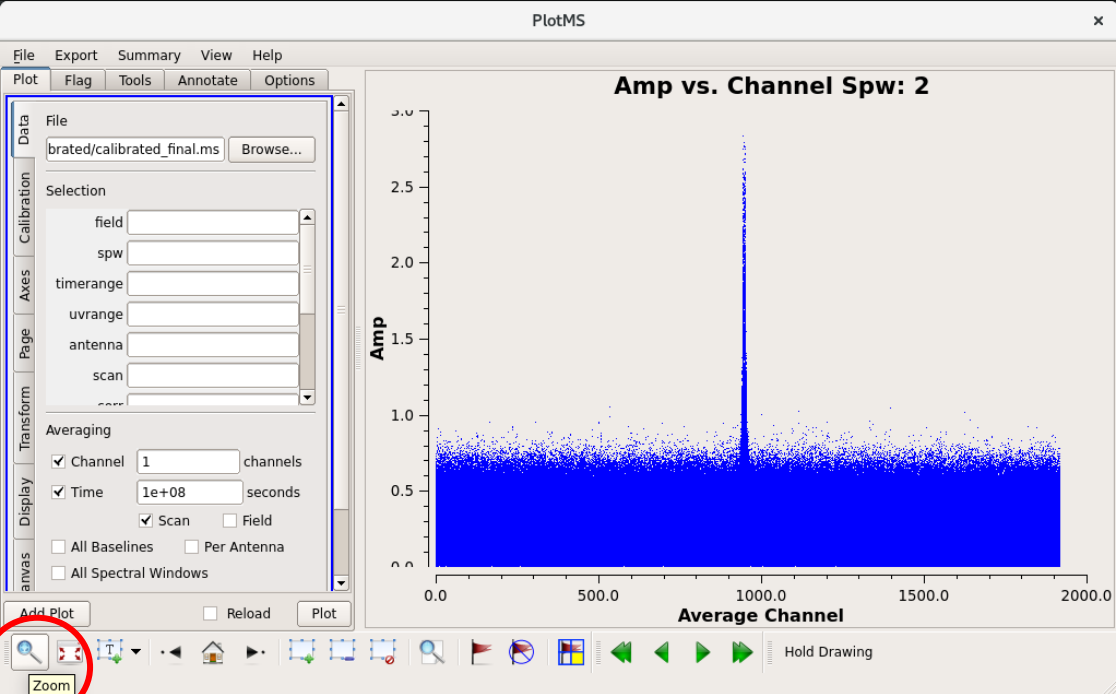
```

plotms(vis=finalvis, xaxis='channel', yaxis='amplitude',
       ydatacolumn='data',
       avgtime='1e8', avgscan=True, avgchannel='1',
       iteraxis='spw' )

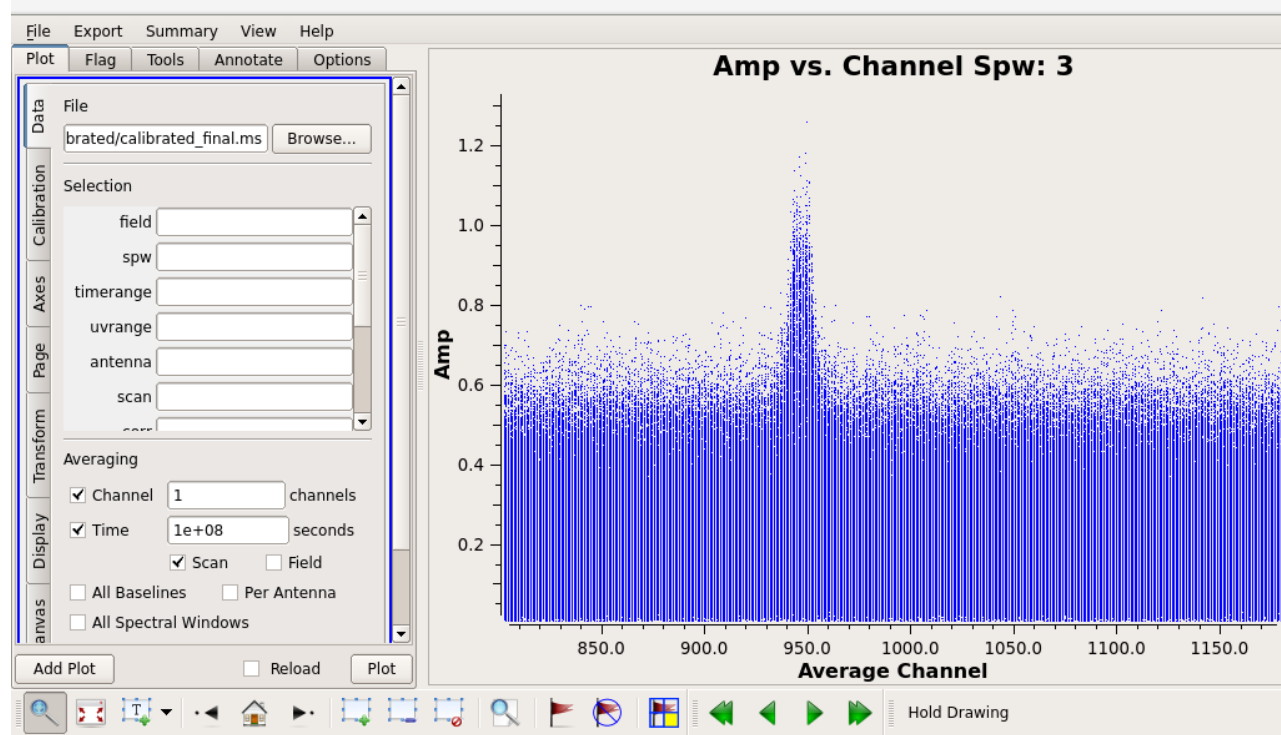
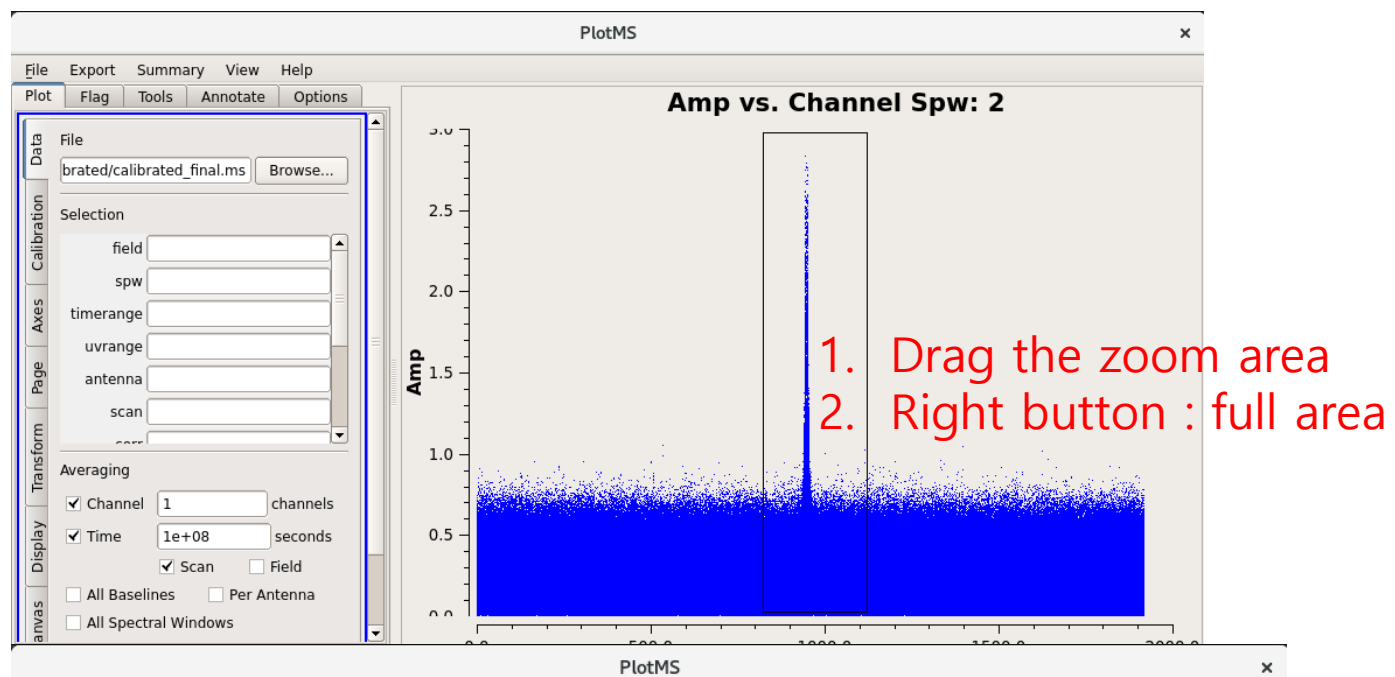
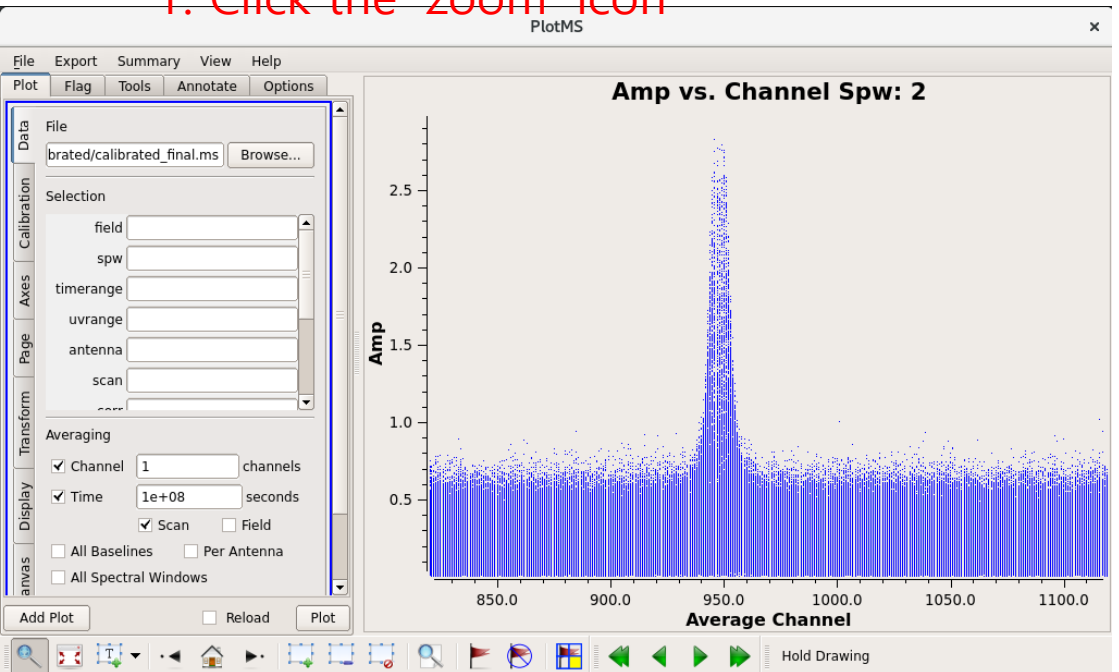
```

SPWs of 19,23,25,27 are renamed with 0,1,2,3





1. Click the 'zoom' icon



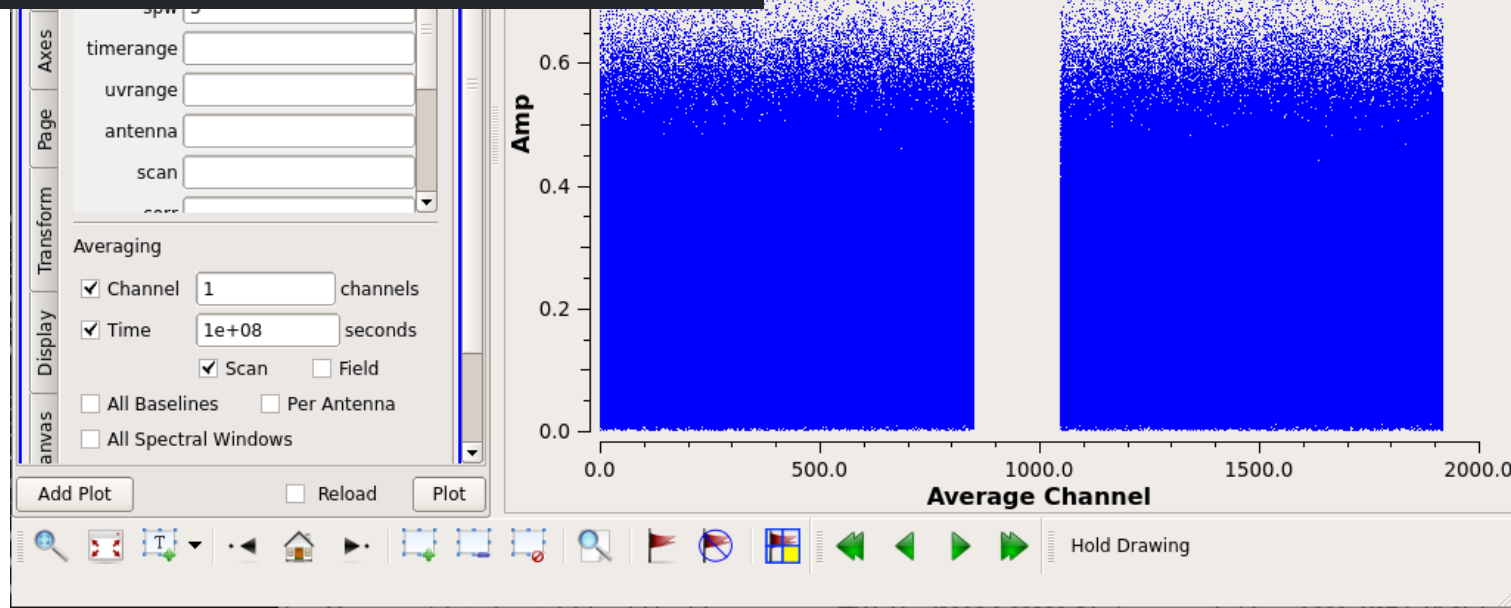
Preparing the continuum ms file (1).

```
# Flag the "line channels"
flagchannels='2:851~1049,3:851~1049' # In this example , spws 2&3 have a line
d 2199 and spectral windows 0 and 1 are line-free.

flagdata(vis=finalvis,mode='manual',
        spw=flagchannels,flagbackup=False)

# check that flags are as expected, NOTE must check reload on plotms
# gui if its still open.
plotms(vis=finalvis,axis='amp',xaxis='channel',
        avgchannel='1',avgtime='1e8',avgscan=True,spw='2')
# avgchannel='1',avgtime='1e8',avgscan=True,iteraxis='spw')
```

1. Flag (delete) line channels.



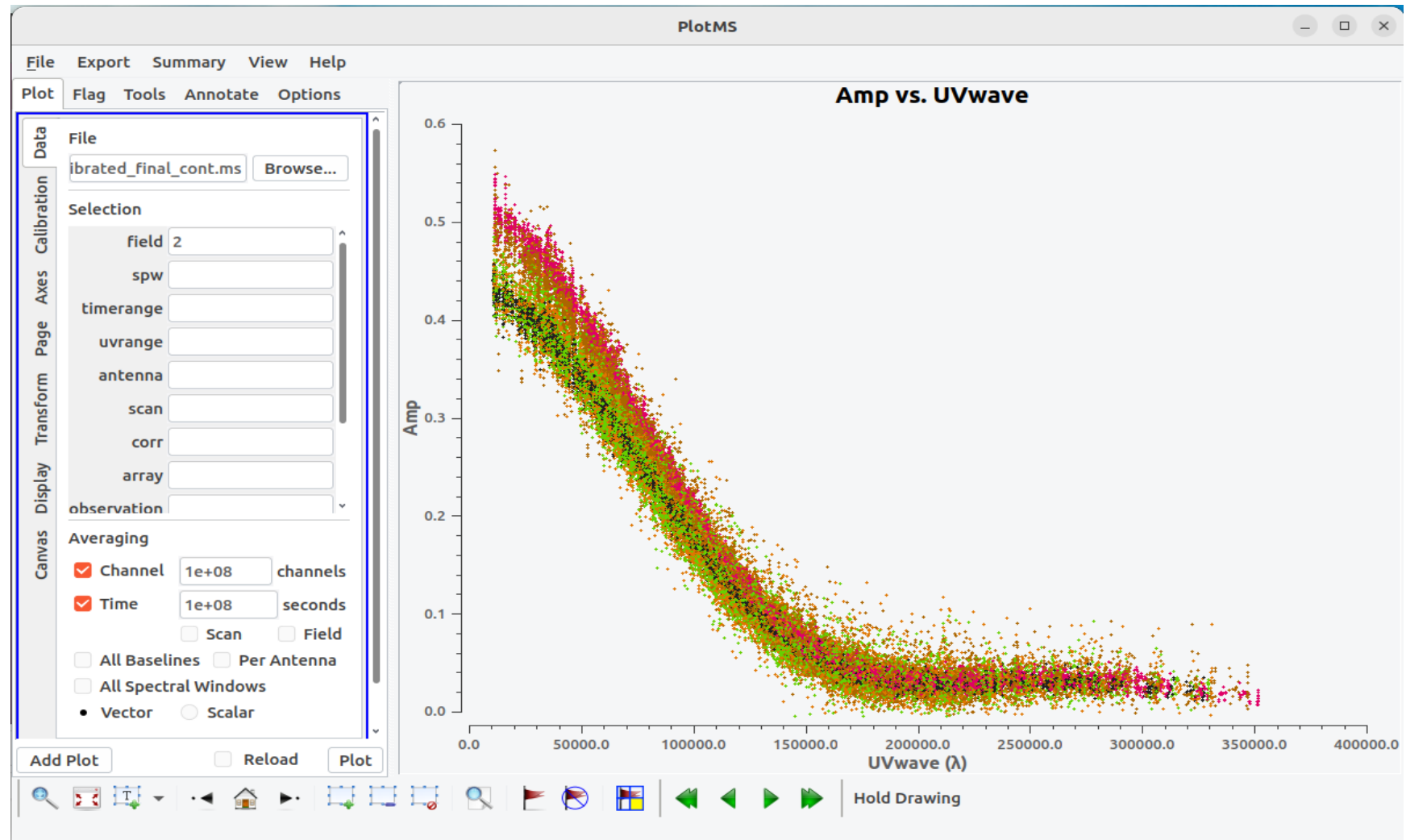
Preparing the continuum ms file.

```
#>>> Note that to mitigate bandwidth smearing, please keep the width
#>>> of averaged channels less than 125MHz in Band 3, 4, and 6, and 250MHz
#>>> in Band 7 for both TDM and FDM modes. For example, for a 2GHz TDM window
#>>> with 15.625 MHz channels, this means that the maximum width parameter
#>>> should be 8 channels for Bands 3, 4, and 6 and 16 channels for Band 7.
#>>> This is especially important for any long baseline data. These limits
#>>> have been designed to have minimize the reduction of the peak flux to
#>>> 95%. See the "for continuum" header for more information on the imaging
#>>> wiki for more infomration.

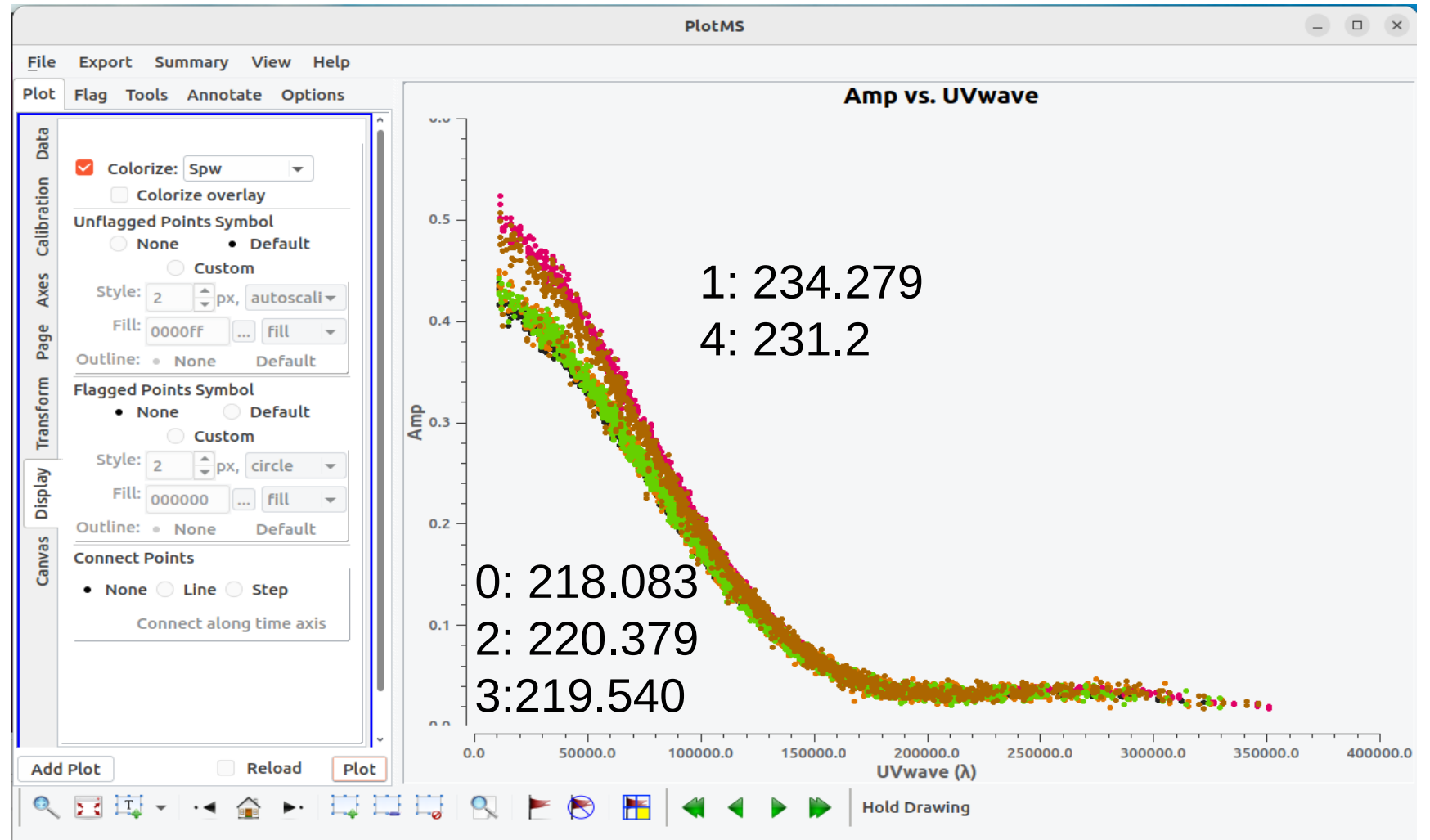
#>>> Note that in CASA 5.1, split2 is now split. Previously split2 was
#>>> needed to deal correctly with channelized weights.
split(vis=finalvis,
      spw=contspws,
      outputvis=contvis,
      width=[16,16,1920,1920,3840], # number of channels to average together. The final
channel width should be less than 125MHz in Bands 3, 4, and 6 and 250MHz in Band 7.
      datacolumn='data')
```

2. Average channels

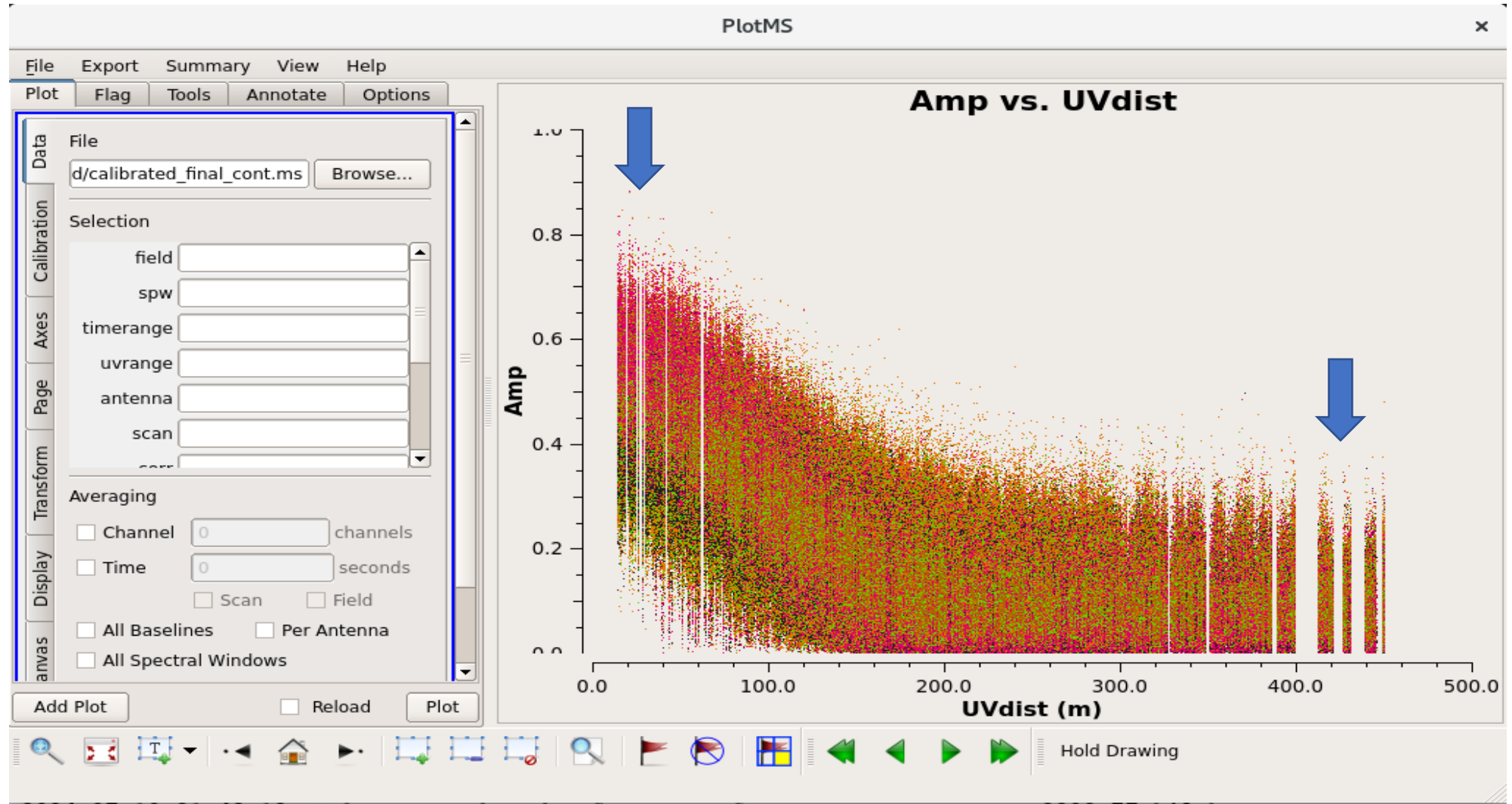
Check UV distance(wave) vs. Amp



Check UV distance(wave) vs. Amp



Discreate UV coverage might produce some features..



Tasks in execution order

- 1. hifa_importdata
- 2. hifa_flagdata
- 3. hifa_fluxcalflag
- 4. hif_rawflagchans
- 5. hif_refant
- 6. hifa_tsyscal
- 7. hifa_tsysflag
- 8. hifa_antpos
- 9. hifa_wvrflag
- 10. hif_lowgainflag
- 11. hif_gainflag
- 12. hif_setij
- 13. hifa_bandpass
- 14. hifa_spwphaseup
- 15. hifa_gfluxscale
- 16. hifa_timegaincal
- 17. hif_applycal
- 18. hif_makeimlist
- 19. hif_makeimages
- 20. hif_checkproductsize
- 21. hif_exportdata
- 22. hif_mstrtransform
- 23. hifa_flagtargets
- 24. hif_makeimlist
- 25. hif_findcont
- 26. hif_uvcontfit
- 27. hif_uvcontsub
- 28. hif_makeimages
- 29. hif_makeimlist
- 30. hif_makeimages
- 31. hif_makeimlist
- 32. hif_makeimages

24. Make image list

Set-up image parameters for target per-spw continuum imaging

Cell and Image sizes used in tclean

BACK

List of Clean Targets

field	Intent	spw	phasecenter	cell	lmsize	imagename	specmode	start	width	nbin	nchan	uvrange
TW_Hya	TARGET	19	ICRS 11:01:51.8126 -034.42.17.276	[0.12arcsec]	[378, 378]	id__A001_X87d_Xb3d.sSTAGENUMBER.TW_Hya_sci.spw19.mfs	mfs			-1	-1	
TW_Hya	TARGET	23	ICRS 11:01:51.8126 -034.42.17.276	[0.12arcsec]	[378, 378]	id__A001_X87d_Xb3d.sSTAGENUMBER.TW_Hya_sci.spw23.mfs	mfs			-1	-1	
TW_Hya	TARGET	25	ICRS 11:01:51.8126 -034.42.17.276	[0.12arcsec]	[378, 378]	id__A001_X87d_Xb3d.sSTAGENUMBER.TW_Hya_sci.spw25.mfs	mfs			-1	-1	
TW_Hya	TARGET	27	ICRS 11:01:51.8126 -034.42.17.276	[0.12arcsec]	[378, 378]	id__A001_X87d_Xb3d.sSTAGENUMBER.TW_Hya_sci.spw27.mfs	mfs			-1	-1	
TW_Hya	TARGET	29	ICRS 11:01:51.8126 -034.42.17.276	[0.12arcsec]	[378, 378]	id__A001_X87d_Xb3d.sSTAGENUMBER.TW_Hya_sci.spw29.mfs	mfs			-1	-1	

Clean Targets Summary

Pipeline QA

Input Parameters

Tasks Execution Statistics

CASA logs for stage 24

- View or download stage24/casapy.log (10.5 KB)



Tasks in execution order

- 1. hifa_restoredata
- 2. hif_mstransform
- 3. hifa_flagtargets
- 4. hifa_imageprecheck
- 5. hif_checkproductsizes
- 6. hif_makeimlist (mfs)
- 7. hif_findcont
- 8. hif_uvcontfit
- 9. hif_uvcontsub
- 10. hif_makeimages (mfs)
- 11. hif_makeimlist (cont)
- 12. hif_makeimages (cont)
- 13. hif_makeimlist (cube)
- 14. hif_makeimages (cube)
- 15. hif_makeimlist (cube_repBW)
- 16. hif_makeimages (cube_repBW)

Representative Target: B335
Representative Frequency: 316.7910 GHz (SPW 25)
Bandwidth for Sensitivity: 1.057 MHz (rounded to nearest integer #channels (9), repBW = 1.099 MHz)
Min / Max Acceptable Resolution: 0.300 arcsec / 0.500 arcsec
Maximum expected beam axial ratio (from OT): 1.5
Goal PI sensitivity: 2.50 mJy
Single Continuum: False

Estimated Synthesized Beam and Sensitivities for the Representative Target/Frequency

Estimates are given for four possible values of the tclean robust weighting parameter: robust = 0.0, +0.5 (default), +1.0, and +2.0. If the "Min / Max Acceptable Resolution" is available ($\geq$ Cycle 5 12-m Array data), the robust value closest to the default (+0.5) that predicts a beam area (defined as simply major x minor) that is in the range of the PI requested beam areas according to the table row for repBW (Bandwidth for Sensitivity) is chosen. If none of these robust values predict a beam area that is in range, robust=+2.0 is chosen if the predicted beam area is too small, and robust=0.0 is chosen if the predicted beam area is too large. The chosen robust value is highlighted in green and used for all science target imaging. In addition to an estimate for the repBW, an estimate for the aggregate continuum bandwidth (aggBW) is also given assuming NO line contamination but accounting for spw frequency overlap. If the Bandwidth for Sensitivity (repBW) is $>$ the bandwidth of the spw containing the representative frequency (repSPW), then the beam is predicted using all spws, otherwise the beam is predicted for the repSPW alone. A message appears on the "By Task" view if a non-default value of robust (i.e., not +0.5) is chosen. Additionally, if the predicted beam is not within the PI requested range using one of the four robust values, Warning messages appear on this page.

These estimates should always be considered as the BEST CASE SCENARIO. These estimates account for Tsys, the observed uv-coverage, and prior flagging. The estimates DO NOT account for (1) subsequent science target flagging; (2) loss of continuum bandwidth due to the hif_findcont process (i.e. removal of lines and other spectral features from the data used to image the continuum); (3) Issues that affect the image quality like (a) poor match of uv-coverage to image complexity; (b) dynamic range effects; (c) calibration deficiencies (poor phase transfer, residual baseline based effects, residual antenna position errors, etc.). *It is also important to note that both the repBW and aggBW beam calculations are intrinsically multi-frequency synthesis continuum calculations, using the relevant spws as described above. The synthesized beam for a single channel in a cube will typically be larger and can be significantly larger depending on the details of uv-coverage and channel width.*

robust	uvtaper	Synthesized Beam	Cell	Beam Ratio	Bandwidth	BW Mode	Effective Sensitivity
0.0	□	0.356 x 0.312 arcsec @ -49.8 deg	0.062 x 0.062 arcsec	1.14	1.099 MHz	repBW	0.0021 Jy/beam
0.0	□	0.346 x 0.312 arcsec @ -58.2 deg	0.062 x 0.062 arcsec	1.14	3237 MHz	aggBW	3.67e-05 Jy/beam
0.5	□	0.372 x 0.343 arcsec @ -51.3 deg	0.069 x 0.069 arcsec	1.08	1.099 MHz	repBW	0.00173 Jy/beam
0.5	□	0.374 x 0.353 arcsec @ -53.7 deg	0.071 x 0.071 arcsec	1.08	3237 MHz	aggBW	2.93e-05 Jy/beam
1.0	□	0.407 x 0.393 arcsec @ -8.71 deg	0.079 x 0.079 arcsec	1.04	1.099 MHz	repBW	0.00159 Jy/beam
1.0	□	0.419 x 0.405 arcsec @ -3.74 deg	0.081 x 0.081 arcsec	1.04	3237 MHz	aggBW	2.68e-05 Jy/beam
2.0	□	0.432 x 0.408 arcsec @ 4.19 deg	0.082 x 0.082 arcsec	1.06	1.099 MHz	repBW	0.00158 Jy/beam
2.0	□	0.447 x 0.425 arcsec @ 12.2 deg	0.085 x 0.085 arcsec	1.06	3237 MHz	aggBW	2.65e-05 Jy/beam

Beam sized and theoretical rms level according to robust parameter.

Tasks in execution order

1. hifa_importdata

2. hifa_flagdata

3. hifa_fluxcalflag

4. hif_rawflagchans

5. hif_refant

6. hifa_tsyscal

7. hifa_tsysflag

8. hifa_antpos

9. hifa_wvrflagcal

10. hif_lowgainflag

11. hif_gainflag

12. hif_setjy

13. hifa_bandpass

14. hifa_spwphaseup

15. hifa_gfluxscale

16. hifa_timegaincal

17. hif_applycal

18. hif_makeimlist

19. hif_makeimages

20. hif_checkproductsizes

21. hif_exportdata

22. hif_mstransform

23. hifa_flagtargets

24. hif_makeimlist

25. hif_findcont

26. hif_uvcontfit

27. hif_uvcontsub

28. hif_makeimages

29. hif_makeimlist

30. hif_makeimages

31. hif_makeimlist

32. hif_makeimages

29. Make image list

Set-up image parameters for target aggregate continuum imaging

BACK

List of Clean Targets

field	intent	spw	phasecenter	cell	imsize	imagename	specmode	start	width	nbin	nchan	uvrange
TW_Hya	TARGET	19,23,25,27,29	ICRS 11:01:51.8126 -034.42.17.276	[0.12arcsec]	[378, 378]	uid___A001_X87d_Xb3d.sSTAGENUMBER.TW_Hya_sci.spw19_23_25_27_29.cont	cont			-1	-1	

Clean Targets Summary

Pipeline QA

Input Parameters

Tasks Execution Statistics

CASA logs for stage 29

- View or download stage29/casapy.log (13.2 KB)

For line imaging, set `threshold` = 3 rms level

Tasks in execution order

1. hifa_importdata

2. hifa_flagdata

3. hifa_fluxcalflag

4. hif_rawflagchans

5. hif_refant

6. hifa_tsyscal

7. hifa_tsysflag

8. hifa_antpos

9. [hifa_wvrqcalflag](#)

10. `hif_lowgainflag`11. `hif_gainflag`

12. hif setiv

13 hifa handpass

14. *bifa* *spwphaseup*15. `bife_afluxscale`16. *life time gain*17. *bif. applied*

1.9. *hif* regulation

4.2. *hif* metabolism

[illegible]

2014-01-01

2.7. [TM_ExportData](#)

22. Thyristor

20. [find_targets](#)24. [Hilf mir](#)29. `fit_indcont`26. `hlf_uvcontfit`27. [hlf_uvcontsub](#)

28. `fit_makeimages`

29. hit_makeimlist

30. `hit_makeimages`

31. hit_makeimlist

32. hif_makeimages

30. Tclean/Makelmages

Make target aggregate continuum images

BACK

Image Details

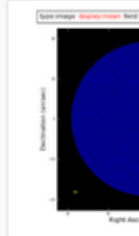
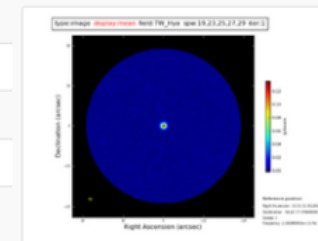
Field	Spw	Pol	Image details		Image result
TW_Hya (TARGET)	19,23,25,27,29	I	center frequency of image	226.1987GHz (LSRK)	
			beam	0.73 x 0.59 arcsec	
			beam p.a.	82.2deg	
			final theoretical sensitivity	1.8e-05 Jy/beam	
			cleaning threshold	0.0033 Jy/beam Dirty DR: 6.8e+03 DR correction: 45	
			non-pbcor image rms	0.00032 Jy/beam	
			pbcor image max / min	0.136 / -0.00310 Jy/beam	
			fractional bandwidth / nterms	7.9% / 1	
			aggregate bandwidth	3.63 GHz (LSRK)	
			score	1.00	
image file			uid__A001_X87d_Xb3d.s30_0_TW_Hya_sci.spw19_23_25_27_29.cont.I.iter1.image		

Image result



[View other QA images...](#)

Real rms level
could be higher
than this value.

Pipeline QA

Input Parameters

Tasks Execution Statistics

CASA logs for stage 30

tclean for continuum image

```
# source parameters
# -----

field='2' # science field(s). For a mosaic, select all mosaic fields. DO NOT LEAVE BLANK
('') OR YOU WILL POTENTIALLY TRIGGER A BUG IN CLEAN THAT WILL PUT THE WRONG COORDINATE
SYSTEM ON YOUR FINAL IMAGE.
gridder='standard' # uncomment if single field
# gridder='mosaic' # uncomment if mosaic or if combining one 7m and one 12m pointing.
cell='0.12arcsec' # cell size for imaging.
imsize = [378,378] # size of image in pixels.

# velocity parameters
# -----

outframe='lsrk' # velocity reference frame.
veltype='radio' # velocity type.

# imaging control
# -----

# The cleaning below is done interactively, so niter and threshold can
# be controlled within clean.

weighting = 'briggs'
robust=0.5
niter=1000
threshold = '0.0mJy'
```

```
contvis = 'calibrated_final_cont.ms'
contimagername = 'cont'

tclean(vis=contvis,
       imagename=contimagername,
       field=field,
       # phasecenter=phasecenter, #
       specmode='mfs',
       deconvolver='hogbom',
       #deconvolver='mtmfs',
       #nterms=2,
       imsize = imsize,
       cell= cell,
       weighting = weighting,
       robust = robust,
       niter = niter,
       threshold = threshold,
       interactive = True,
       gridder = gridder,
       pbcor = True,
       usepointing=False)
```

Deconvolver

hogbom : default

clark

clarkstokes : for polarization

multiscale : Extended sources

scales =[0,5,15] (in unit of pixel)

[point, beam size, 3~5 x beamsize]

mtmfs : Multiterm(multiscale) + multifrequency

too broad frequency is covered in the continuum
imaging and emission varies with frequency.

[Release Information](#)[Index](#)[API \(tasks, tools, GUIs, etc.\)](#)☐ [Task List \(shortcut\)](#)[Input / Output](#)[Information](#)[Flagging](#)[Calibration](#)☒ [Imaging](#)[Home](#) / [API \(tasks, tools, GUIs, etc.\)](#) / [casatasks](#) / [tclean](#)[Edit on GitHub](#)

tclean

```
tclean(vis, selectdata=True, field="", spw="", timerange="", uvrange="", antenna="", scan="",  
observation="", intent="", datacolumn='corrected', imagename="", imsize=[100], cell="1arcsec",  
phasecenter="", stokes='I', projection='SIN', startmodel="", specmode='mfs', reffreq="", nchan=-1,  
start="", width="", outframe='LSRK', veltype='radio', restfreq="", interpolation='linear',  
perchanweightdensity=True, gridder='standard', facets=1, psfphasecenter="", wprojplanes=1,  
vptable="", mosweight=True, aterm=True, psterm=False, wbawp=True, conjbeams=False,  
cfcache="", usepointing=False, computeastep=360.0, rotatepastep=360.0,  
pointingoffsetsigdev="", pblimit=0.2, normtype='flatnoise', deconvolver='hogbom', scales="",  
nterms=2, smallscalebias=0.0, fusedthreshold=0.0, largestscale=-1, restoration=True,  
restoringbeam="", pbcor=False, outlierfile="", weighting='natural', robust=0.5, noise='1.0Jy',  
npixels=0, uvtaper=[""], niter=0, gain=0.1, threshold=0.0, nsigma=0.0, cycleniter=-1,  
cyclefactor=1.0, minpsffraction=0.05, maxpsffraction=0.8, interactive=False, nmajor=-1,  
fullsummary=False, usemask='user', mask="", pbmask=0.0, sidelobethreshold=3.0,  
noisethreshold=5.0, lownoisethreshold=1.5, negativethreshold=0.0, smoothfactor=1.0,  
minbeamfrac=0.3, cutthreshold=0.01, growiterations=75, dogrowprune=True,  
minpercentchange=-1.0, verbose=False, fastnoise=True, restart=True, savemodel='none',  
calcres=True, calcpsf=True, psfcutoff=0.35, parallel=False) \[source\]
```

```
RuntimeError: Error in selectData() : Data selection ended with 0 rows
#####
##### Begin Task: tclean #####
tclean( vis='calibrated_final_cont.ms', selectdata=True, field='2', spw='', timerange='', uvrange='', antenna='', scan='', obs=
Verifying Input Parameters
MS : calibrated_final_cont.ms | Selecting on fields : 2 | [Opened in readonly mode]
  NRows selected : 1640000
Leap second table TAI.UTC seems out-of-date.
Until the table is updated (see the CASA documentation or your system admin),
times and coordinates derived from UTC could be wrong by 1s or more.
imsize with 378 pixels is not an efficient imagesize. Try 384 instead.
Define image coordinates for [cont] :
Impars : start
Shape : [378, 378, 1, 1]Spectral : [2.26199e+11] at [0] with increment [1.81969e+10]
Set Gridding options for [cont] with ftmachine : gridft
Required memory: 0.007725 GB. Available mem.: 189.6 GB (rc, mem. fraction: 80%, memory: -) => Subcubes: 1. Processes on node
Set imaging weights : Briggs weighting: sidelobes will be suppressed over full image
Normal robustness, robust = 0.5
Set Deconvolution Options for [cont] : hogbom
----- Make PSF -----
[cont] Theoretical sensitivity (Jy/bm):1.58883e-05
Time to fit Gaussian to PSF 0.007957
Beam : 0.730786 arcsec, 0.585738 arcsec, 82.91 deg
vi2 : Evaluating Primary Beam model onto image grid(s)
----- Run Major Cycle 1 -----
Absolute Peak residual within mask : 0.123077, over full image : 0.123077
[cont] Initializing new mask to 0.0 for interactive drawing
[cont] Number of pixels in the clean mask : 0 out of a total of 142884 pixels. [ 0 % ]
```

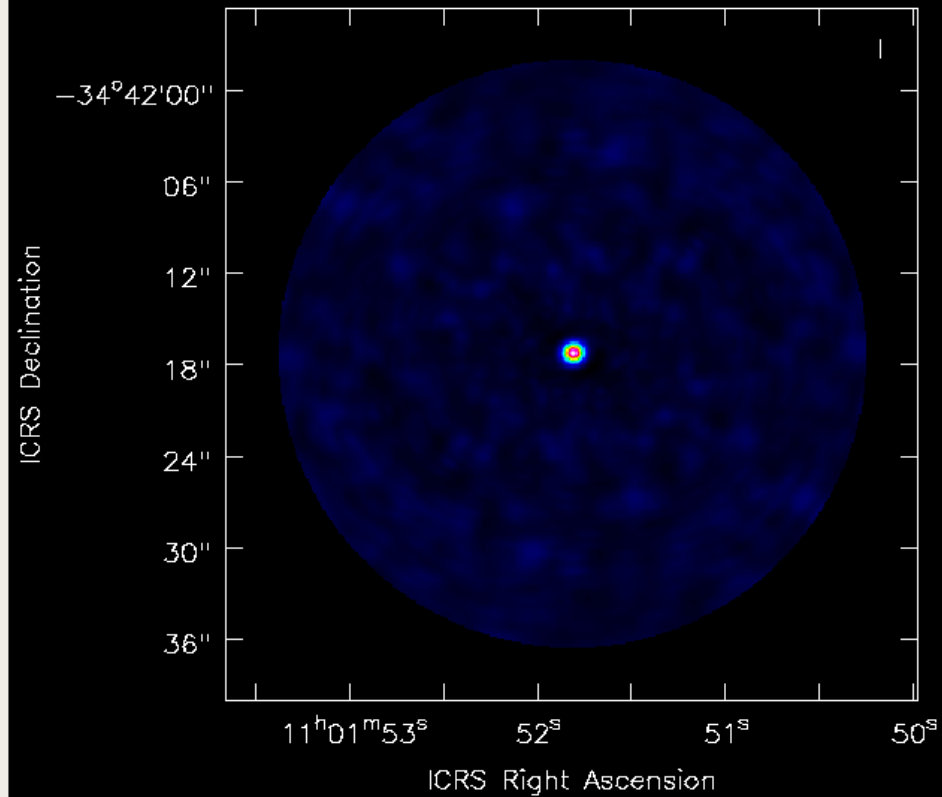
Data Display Panel Tools View Help



☒ Add	☒ This Channel	☒ This Polarization	Next Action:
☐ Erase	☐ All Channels	☐ All Polarizations	
max cycleniter	iterations left	threshold	cyclethreshold
100	1000	0.000000Jy	0.006624Jy

Display

cont.residual-raster



Animators

☐ Stokes☒ Images

Rate: 10

Jump ☐ 0 2

0

1

Cursors

☒ cont.residual-raster☒ cont.mask



Click here with the desired mouse button to assign that button to 'Zooming'

- ☒ Use the assigned mouse button to drag out a rectangle.
- ☐ Use handles to resize.
- ☐ Double click inside rectangle-> zoom in
- ☐ Double click outside rectangle -> zoom out
- ☐ <Esc> to cancel

Next Action:



cyclethreshold

100

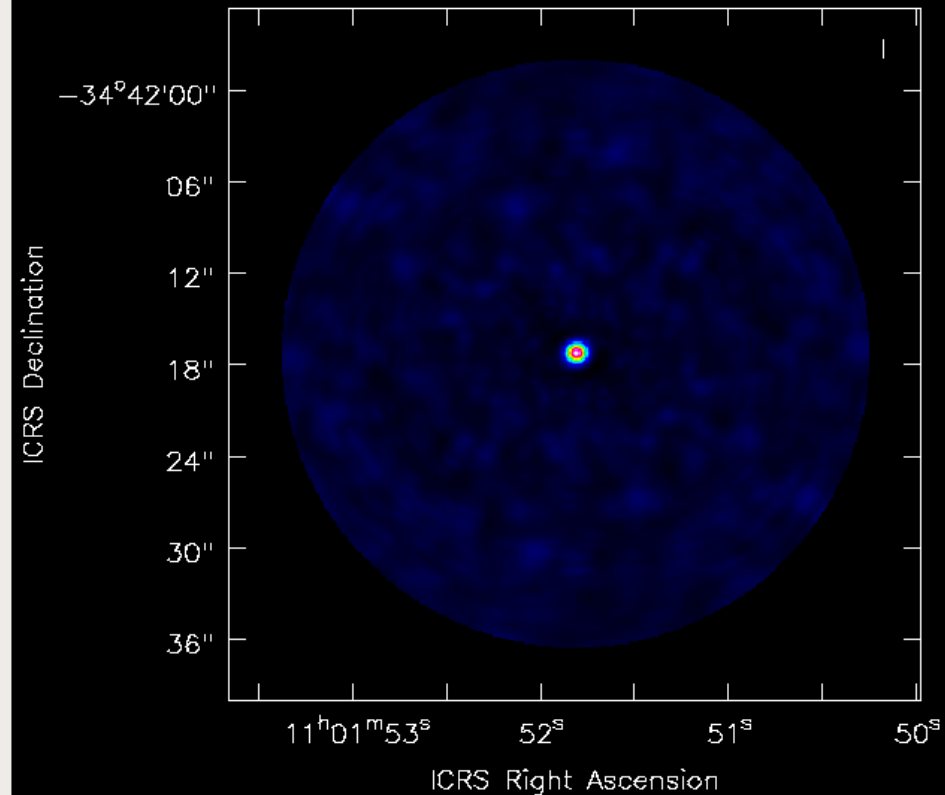
1000

0.000000jy

0.006624jy

Display

cont.residual-raster



Animators

☐ Stokes☒ Images

Rate: 10

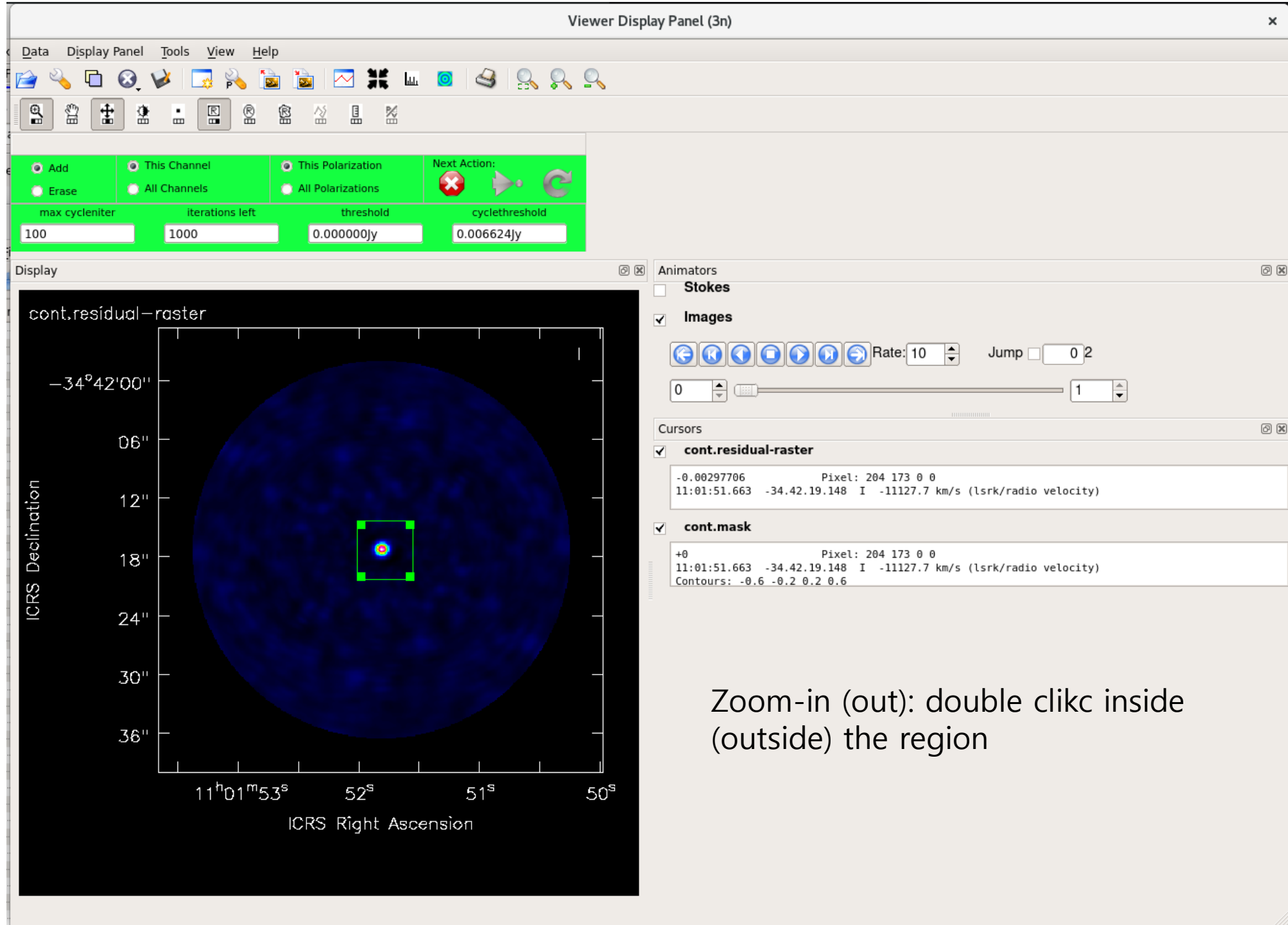
Jump ☐ 0 2

0

1

Cursors

☒ cont.residual-raster☒ cont.mask



Zoom-in (out): double click inside
(outside) the region



Region: make mask where we extract the model signal

Add

Erase

This Channel

All Channels

This Polarization

All Polarizations

Next Action:

max cycleniter

100

iterations left

1000

threshold

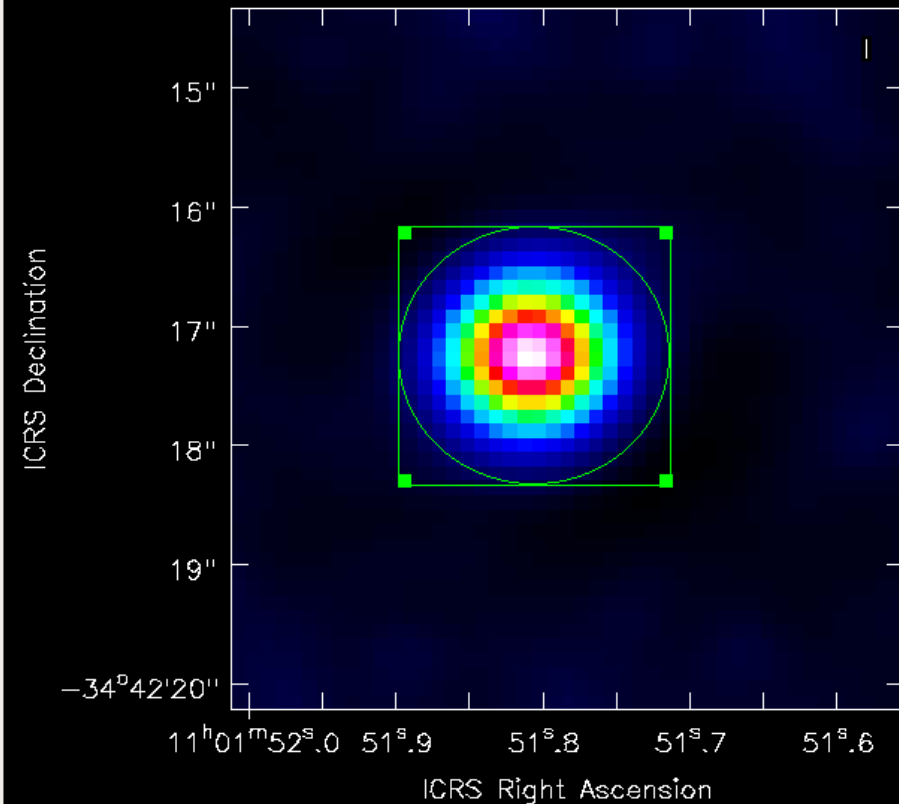
0.000000Jy

cyclethreshold

0.006624Jy

Display

cont.residual-raster



Animators

☐ Stokes☒ Images

Rate: 10

Jump ☐ 0 2

0

1

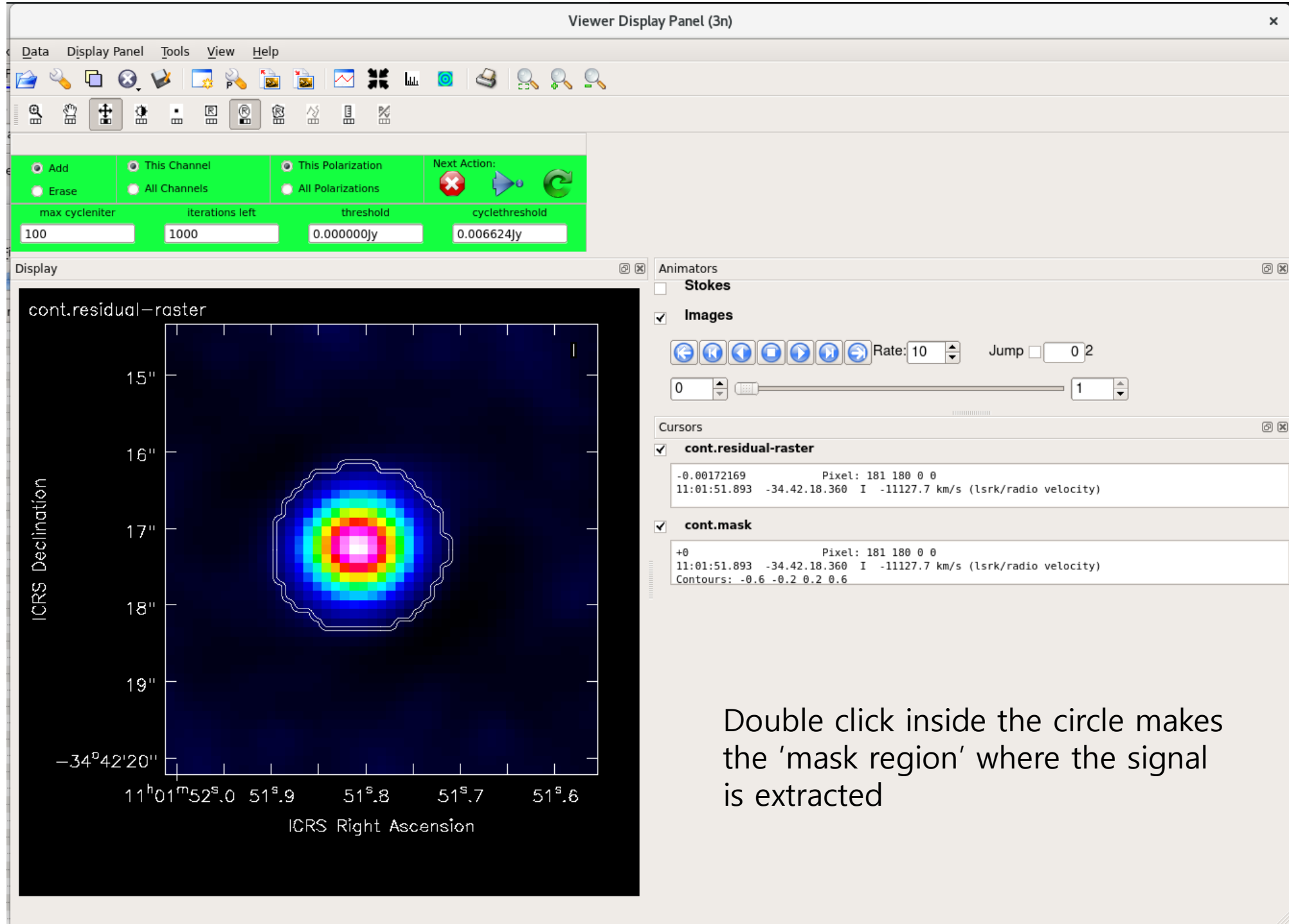
Cursors

☒ cont.residual-raster

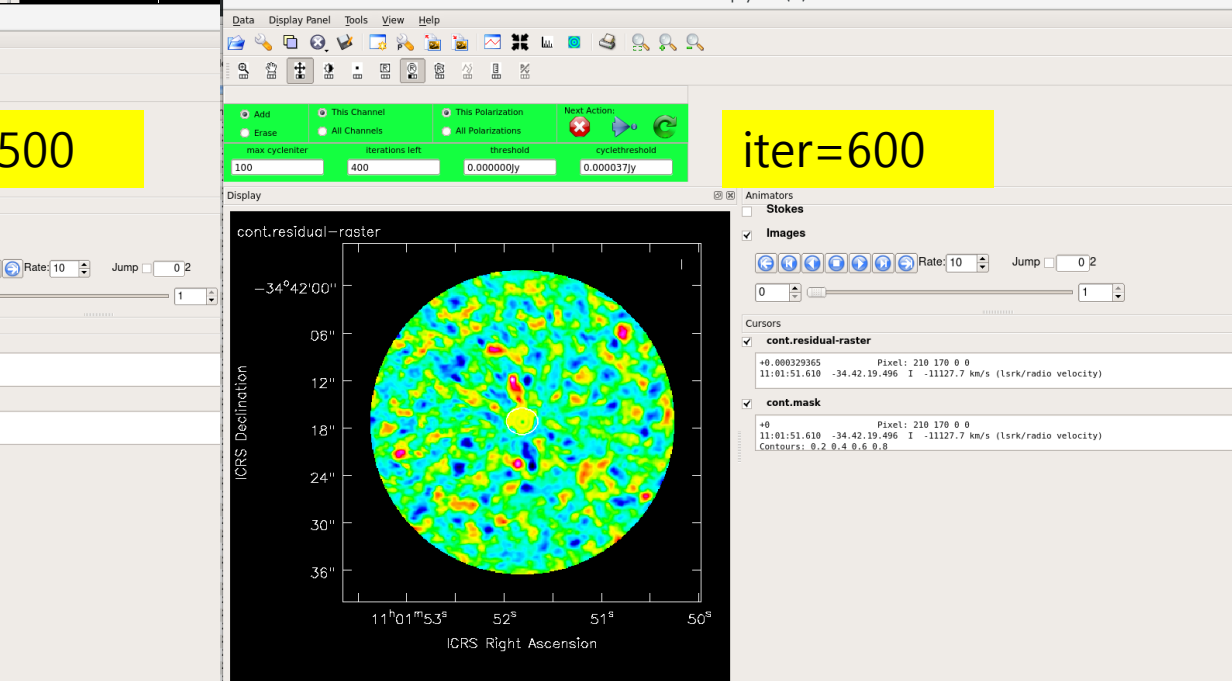
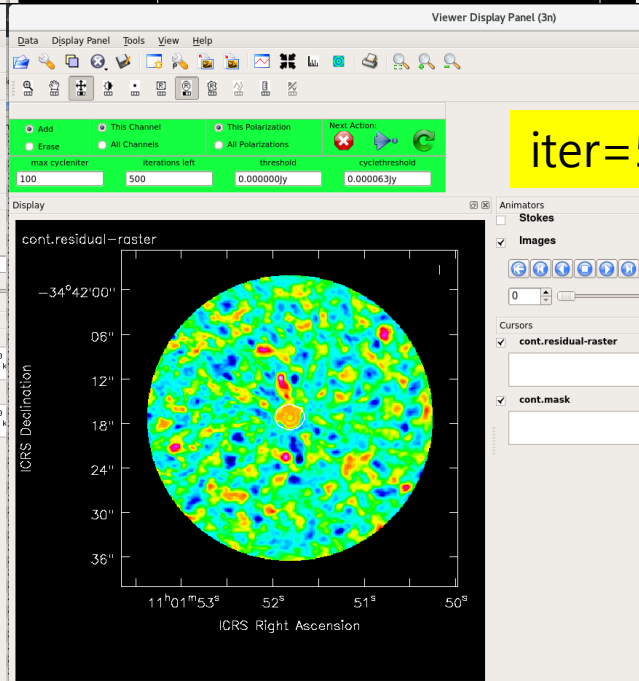
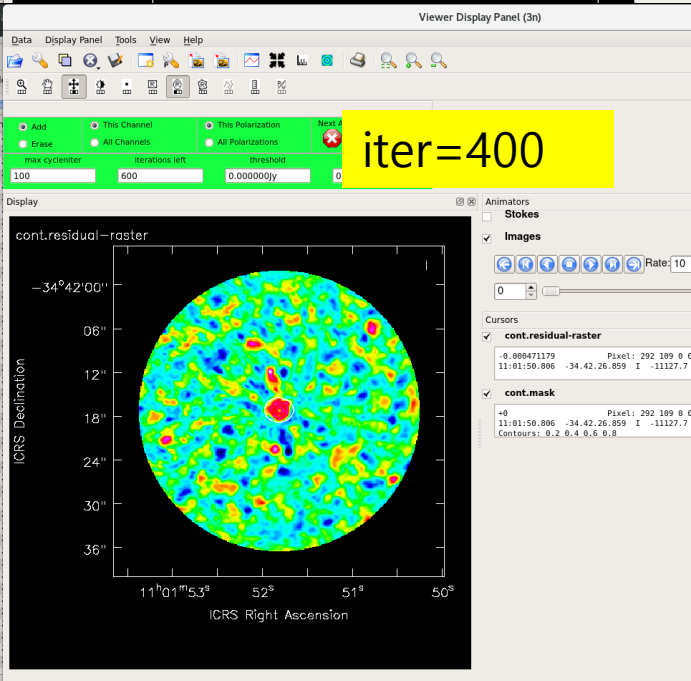
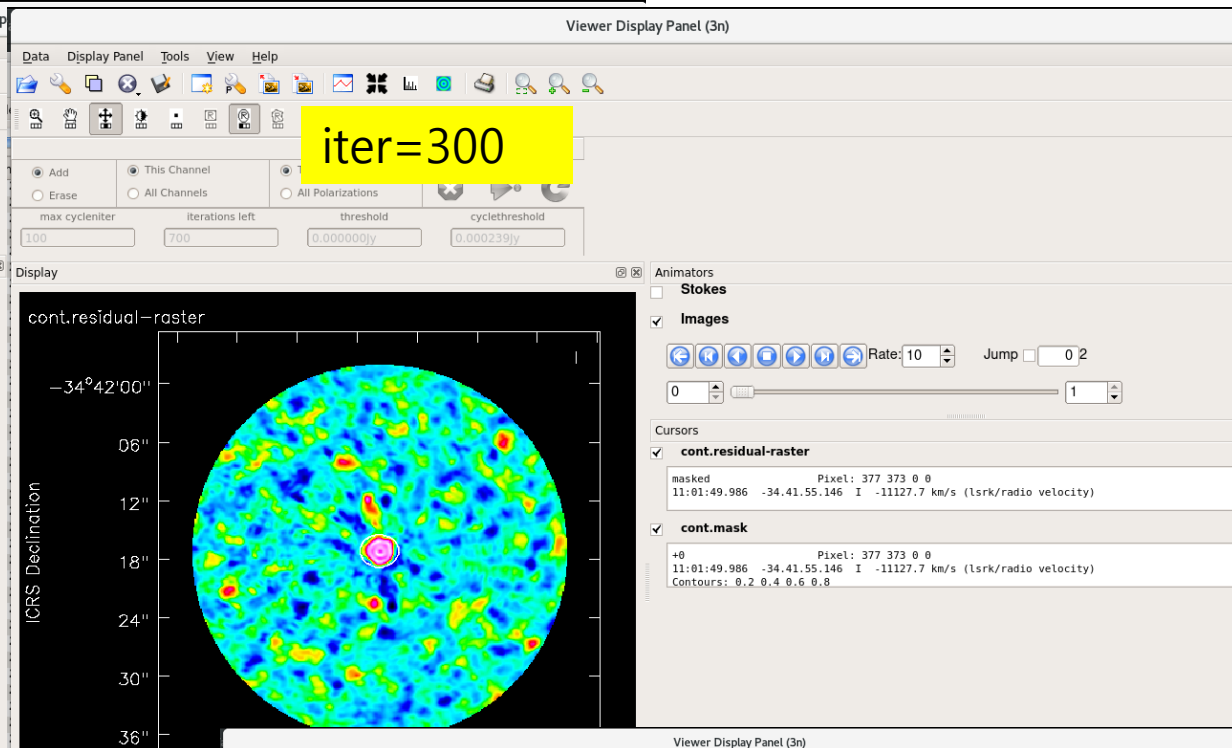
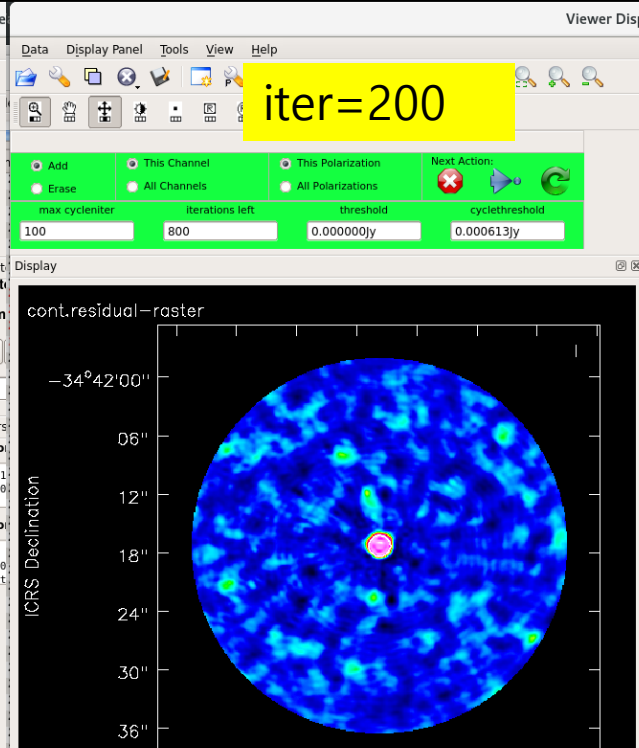
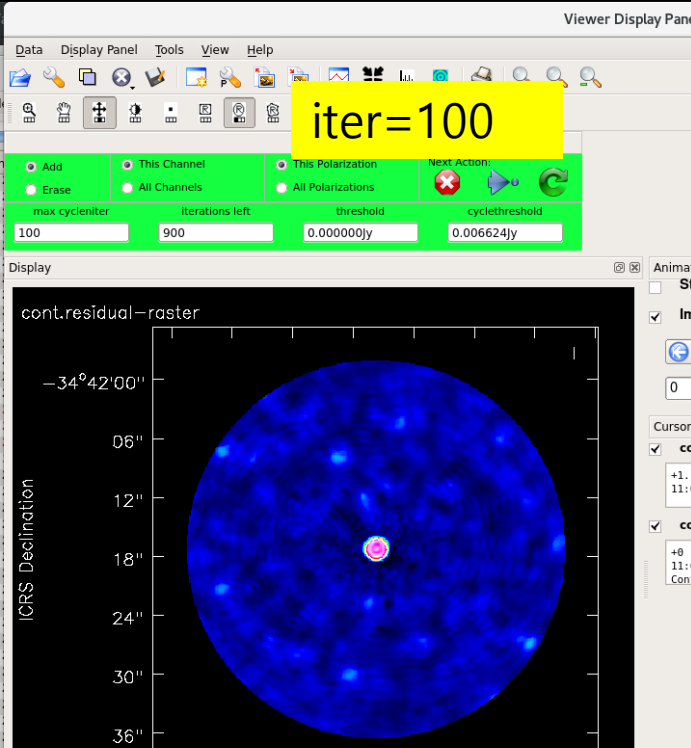
+0.00695825 Pixel: 196 184 0 0
11:01:51.748 -34.42.17.884 I -11127.7 km/s (lsrk/radio velocity)

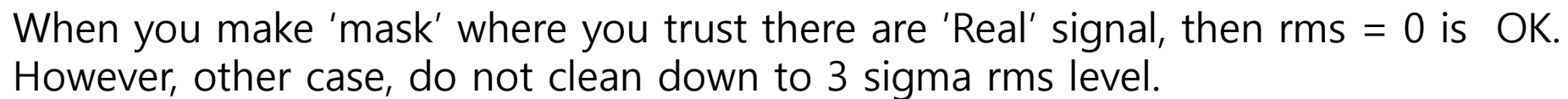
☒ cont.mask

+0 Pixel: 196 184 0 0
11:01:51.748 -34.42.17.884 I -11127.7 km/s (lsrk/radio velocity)
Contours: -0.6 -0.2 0.2 0.6



```
----- Run Major Cycle 1 -----  
Absolute Peak residual within mask : 0.123077, over full image : 0.123077  
[cont] Initializing new mask to 0.0 for interactive drawing  
[cont] Number of pixels in the clean mask : 0 out of a total of 142884 pixels. [ 0 % ]  
[cont.mask] Mask modified from 0 pixels to 267 pixels  
[cont] Mask changed interactively.  
Absolute Peak residual within mask : 0.123077, over full image : 0.123077  
Minor Cycle controls : {'cycleniter': 100, 'cyclethreshold': 0.006623855326324701, 'loopgain': 0.10000000149011612, 'nsigma': 0  
    itsNsigma=0  
[cont] Run Hogbom minor-cycle | CycleThreshold=0.00662386, CycleNiter=100, Gain=0.1  
[cont] iters=0->100 [100], model=0->0.306273, peakres=0.123077->0.0113892, Reached cycleniter.  
Completed 100 iterations.  
----- Run Major Cycle 2 -----
```





Products of 'tclean'

try.psf	Point Spread Function
try.pb	Primary Beam
try.residual	Residual Image (or initial Dirty Image)
try.model	Model Image after deconvolution
try.image	Restored output image
try.image.pbcor	Primary Beam corrected image
try.mask	Deconvolution mask

try.psf.tt0, try.psf.tt1, try.psf.tt2, try.model.tt0, try.model.tt1, try.residual.tt0, try.residual.tt1, try.image.tt0, try.image.tt1, etc...	Multi-term images representing Taylor coefficients (of polynomials that model the sky spectrum)
try.alpha	Spectral index, for multi-term wideband imaging
try.alpha.error	Estimate of error on spectral index
try.beta	Spectral curvature for multi-term wideband images (if nterms > 2)

tclean for line cube image

```
vis_c18o='twhya_c18o_sub.ms'
plotms(vis=vis_c18o,xaxis='channel',yaxis='amplitude',avgtime='1e8',avgscan=True, avgchanne
l='1')

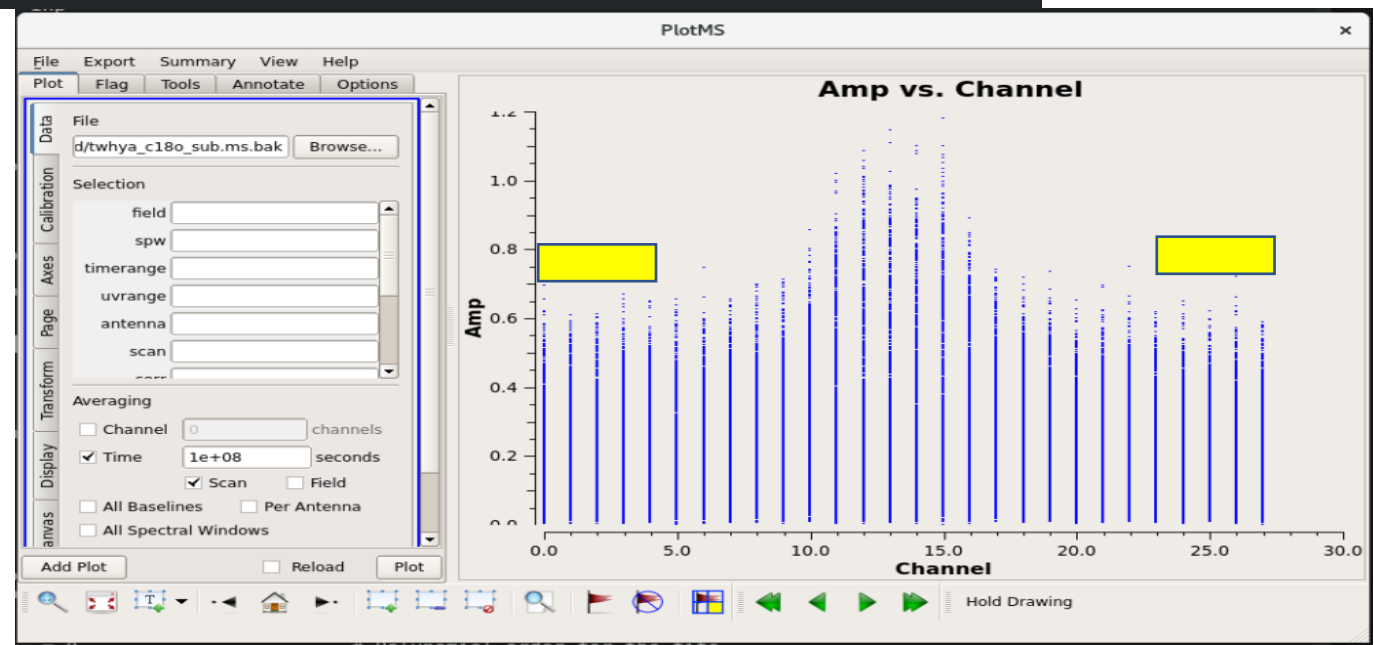
fitspw = '0:0~4;23~27' # *line-free* channels for fitting continuum
linespw = '0' # line spectral windows. You can subtract the continuum from multiple spectra
l line windows at once.

linevis='twhya_c18o_sub_line.ms'

uvcontsub(vis=vis_c18o,
          spw=linespw, # spw to do continuum subtraction on
          fitspw=fitspw, # regions without lines.
          excludechans=False, # fit the regions in fitspw
          solint='int',
          fitorder=0,
          want_cont=False) # This value should not be changed.
```

Input parameters are changed in the recent CASA.

```
uvcontsub(vis=vis_c18o,
          outputvis=linevis,
          fitspec =fitspw,
          fitorder=0)
```



```
gridder='standard' # uncomment if single field
cell='0.12arcsec' # cell size for imaging.
#imsize = [378,378] # size of image in pixels.
imsize = [384,384] # size of image in pixels.
outframe='lsrk' # velocity reference frame.
veltype='radio' # velocity type.
```

```
spw = '0'
field='2'
weighting = 'briggs'
robust=0.5
niter=1000
threshold = '0.0mJy'
```

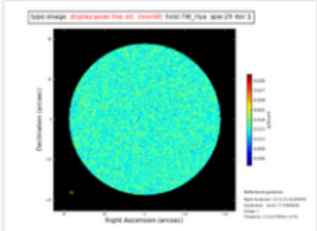
```
start='0' # start velocity. See science goals for appropriate value.
width='1' # velocity width. See science goals.
nchan = 27 # number of channels. See science goals for appropriate value.
restfreq='219.5603541GHz' # Typically the rest frequency of the line: C18O
```

```
linevis='twhya_c18o_sub_line.ms'
lineimagename = 'c18o.ch'
```

Tasks in execution order

1. hifa_importdata
2. hifa_flagdata
3. hifa_fluxcalflag
4. hif_rawflagchans
5. hif_refant
6. hifa_tsyscal
7. hifa_tsysflag
8. hifa_antpos
9. hifa_wvrgcalflag
10. hif_lowgainflag
11. hif_gainflag
12. hif_setjy
13. hifa_bandpass
14. hifa_spwphaseup
15. hifa_gfluxscale
16. hifa_timegaincal
17. hif_applycal
18. hif_makeimlist
19. hif_makeimages
20. hif_checkproductsizes
21. hif_exportdata
22. hif_mstransform
23. hifa_flagtargets
24. hif_makeimlist
25. hif_findcont
26. hif_uvcontfit
27. hif_uvcontsub
28. hif_makeimages
29. hif_makeimlist
30. hif_makeimages
31. hif_makeimlist
32. hif_makeimages



			Dirty DR: 79 DR correction: 2	
			non-pbcor image rms	0.0033 Jy/beam
			pbcor image max / min	0.263 / -0.0530 Jy/beam
			channels	1918 x -0.0611MHz (LSRK)
			score	1.00
			image file	uid__A001_X87d_Xb3d.s32_0.TW_Hya_sci.spw25.cube.l.iter1.image
TW_Hya (TARGET)	27	I	center frequency of cube	219.5575GHz (LSRK)
			beam	0.78 x 0.62 arcsec
			beam p.a.	83.1deg
			final theoretical sensitivity	0.0025 Jy/beam
			cleaning threshold	0.015 Jy/beam Dirty DR: 38 DR correction: 1.5
			non-pbcor image rms	0.0027 Jy/beam
			pbcor image max / min	0.105 / -0.0434 Jy/beam
			channels	1918 x -0.0611MHz (LSRK)
			score	1.00
			image file	uid__A001_X87d_Xb3d.s32_0.TW_Hya_sci.spw27.cube.l.iter1.image
TW_Hya (TARGET)	29	I	center frequency of cube	231.2177GHz (LSRK)
			beam	0.74 x 0.59 arcsec
			beam p.a.	82.3deg
			final theoretical sensitivity	0.0037 Jy/beam
			cleaning threshold	0.015 Jy/beam Dirty DR: 7.6
				


```

tclean(vis=linevis,
       imagename=lineimagename,
       field=field,
       spw=spw,
       # phasecenter=phasecenter,
       # mosweight = True, # unco
       specmode='cube', # comment
       # specmode='cubesource', #
       perchanweightdensity=True,
       start=start,
       width=width,
       nchan=nchan,
       outframe=outframe,
       veltype=veltype,
       restfreq=restfreq,
       niter=niter,
       threshold=threshold,
       interactive=True,
       cell=cell,
       imsize=imsize,
       weighting=weighting,
       robust=robust,
       gridder=gridder,
       pbcor=False,
       restoringbeam='common',
       usepointing=False)

```

specmode = 'cube' ← 'mfs' for continuum

```

start='1km/s'
width='0.2km/s'
nchan = 21

```

```

start='0' #
width='1' #
nchan = 27

```

restfreq = frequency of the targeted line

Niter should be increased by niter (for continuum) x nchan

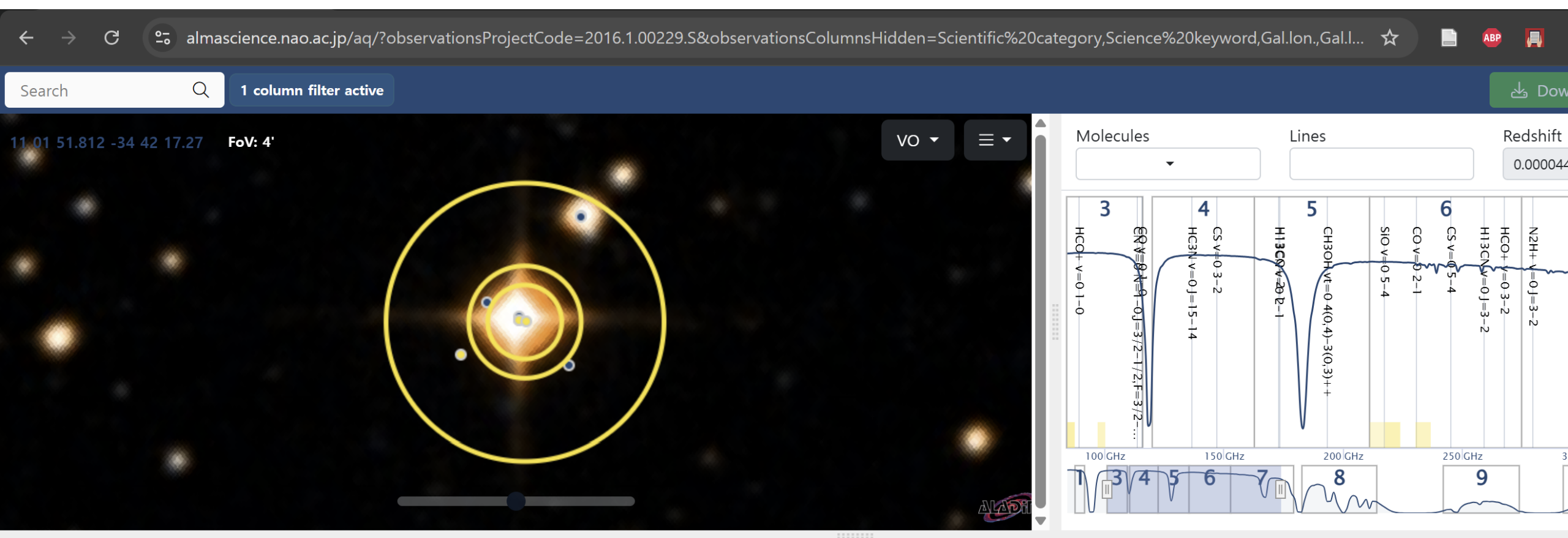
Mask

- You set mask for every channels or the same mask for all channels.
- Automasking
 - usemask='auto-multithresh'
 - Threshold= '## mJy' # 3 * rms level

Automasking

Parameter	7m (continuum/line)	12m (b75 < 300m)	12m (300m < b75 < 400m)	12m (b75 > 400m)	12m + 7m combined <i>TENTATIVE</i>
<i>noisethreshold</i>	5.0	4.25	5.0	5.0	4.25
<i>sidelobethreshold</i>	1.25	2.0	2.0	2.5	2.0
<i>lownoisethreshold</i>	2.0	1.5	1.5	1.5	1.5
<i>minbeamfrac</i>	0.1	0.3	0.3	0.3	0.3
<i>negativethreshold</i>	0.0	0.0 (continuum) 15.0 (line)	0.0 (continuum) 7.0 (line)	0.0 (continuum) 7.0 (line)	0.0
<i>fastnoise</i>	False	False	False	True	False

https://casaguides.nrao.edu/index.php/Automasking_Guide_CASA_6.5.4



Observations (3) Projects (5180) Publications (4472)

Project code: 2016.1.00229.S

	Project code	ALMA source name	FOV	arcsec	Int.Time	s	Min.freq.res.	kHz	L80 BL	m	Obs.date	SB name	Line sens.(10 km/s)	mJy/beam	Group	ous id
☐	2016.1.00229.S															
☐	2016.1.00229.S	tw_hya	16.937		2963.520		244.1314		351.852		2016-11-23	tw_hya_a_07_TM1	0.544			uid://A001/X87d/Xb36
☐	2016.1.00229.S	TW_Hya	25.743		2419.200		122.0798		237.619		2017-04-30	TW_Hya_a_06_TM1	0.539			uid://A001/X87d/Xb3c
☐	2016.1.00229.S	TW_Hya	63.664		2612.736		61.0408		1221.902		2017-08-01	TW_Hya_a_03_TM1	0.480			uid://A001/X87d/Xb30

Tips.

- Save the niter and threshold to run the script automatically.
- Rename the 'imge.mask' to 'image_mask.mask'
- When you run the automatic script, add/change the script
mask = 'image_mask.mask'
Interactive = False

When you rerun the 'tclean', add following paramters,

calcpsf=False

calcres=False

Because we already have '##.psf' and '##.residual' and we do not need to make these again.

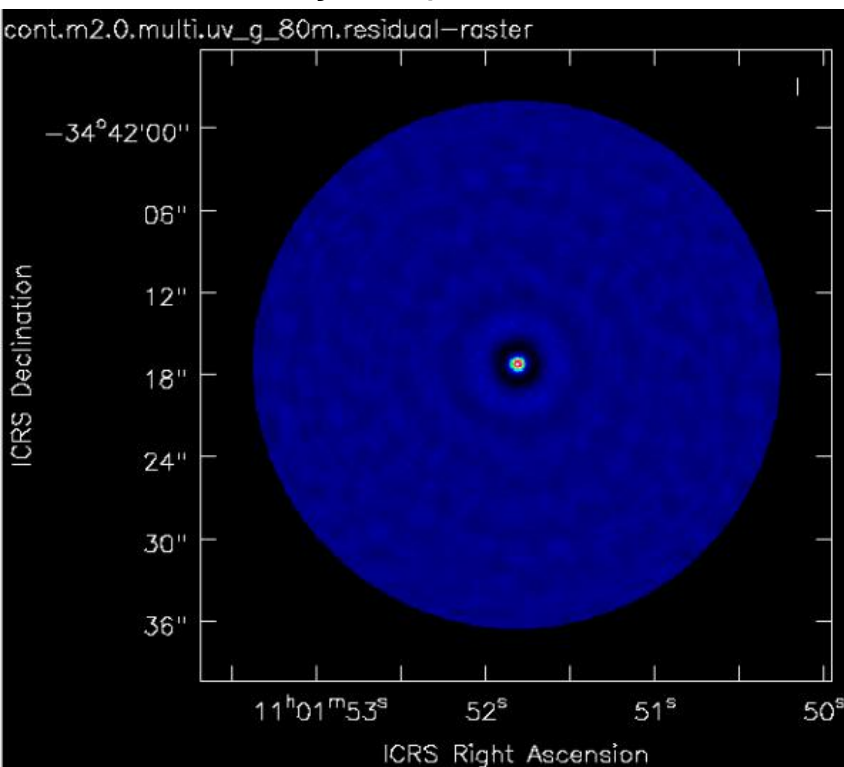
For parallel tclean (for line data)

In the script for tclean, parallel = True

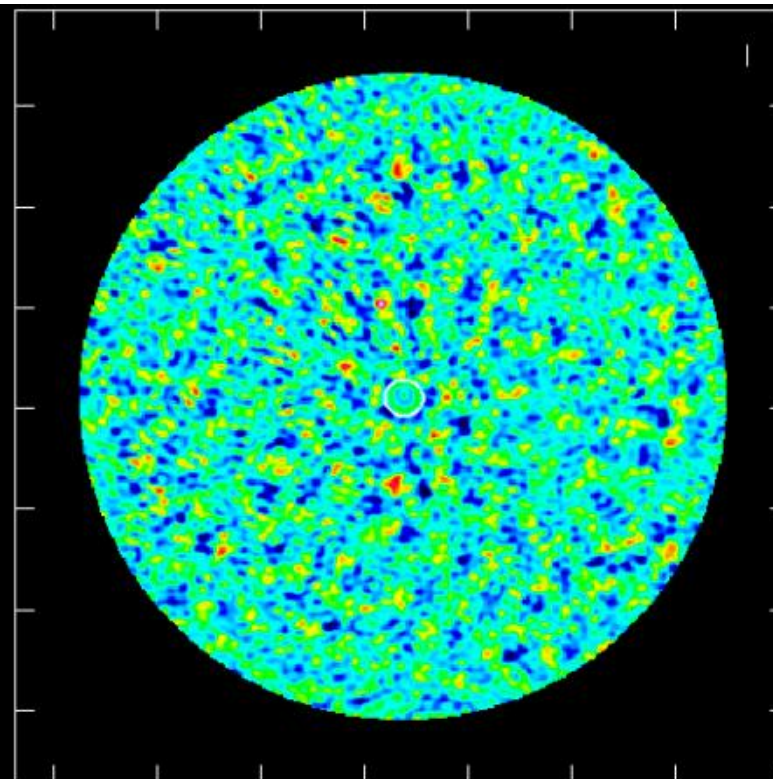
```
/###/bin/mpicasa -n 10 /###/bin/casa -c *.py
```


Exam: Uniform (uvrange > 80m)

Dirty map (niter=0)



Final residual map



Gap and rings in Dirty map is due to sidelobe of PSF. Thus, after several iteration of clean process, the feature disappear and become noise-like.