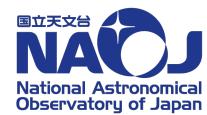
# **ASTE Status Report & Future Prospect**

#### Tetsuhiro MINAMIDANI (ASTE Project Manager - based in NAOJ Mitaka since 2023 April)

National Astronomical Observatory of Japan, NAOJ Observatorio Astronómico Nacional de Japón



## Short Summary of Recent Status ASTE

- ASTE has backed to Science Operation!!
  - Start up: 2023 May -
  - Recovery from Sub-Reflector Driving System malfunction
  - Heterodyne session:
    - CSV: June -
    - Science Observations: July Sep.
  - DESHIMA Session:
    - Oct. Dec.
- Suspended due to Sub-Reflector Driving System Malfunction
  - middle of Nov. -

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- Recovery activities at the ASTE Site will be resumed from middle of Feb. 2024.
- Plan to resume operation as planned.



# <u>Atacama</u> <u>Submillimeter</u> <u>Telescope</u> <u>Experiment</u>

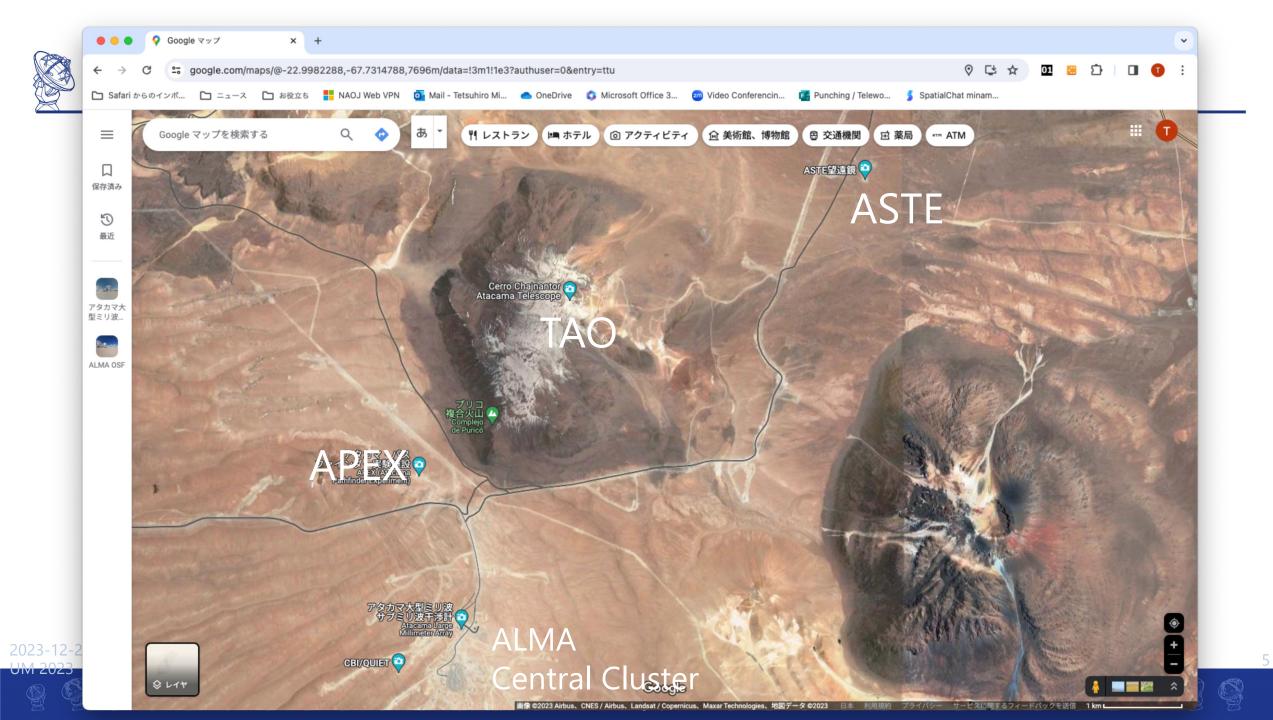
- The prime objectives of ASTE Project
  - **Promote science** with the submm single-dish telescope
  - **Promote development** of instruments and techniques for astronomical observations
  - → A precursor to ALMA and Submm Astronomy in terms of Science and Instrumentation
- 10-m sub-mm telescope at Pampa la Bola 4800m alt. (within the ALMA Site)
  - Since 2002
  - Surface accuracy: 19 $\mu$ m ( $\rightarrow$  ~40 $\mu$ m?)
  - Pointing accuracy: ~2" (rms)
- Site infrastructure

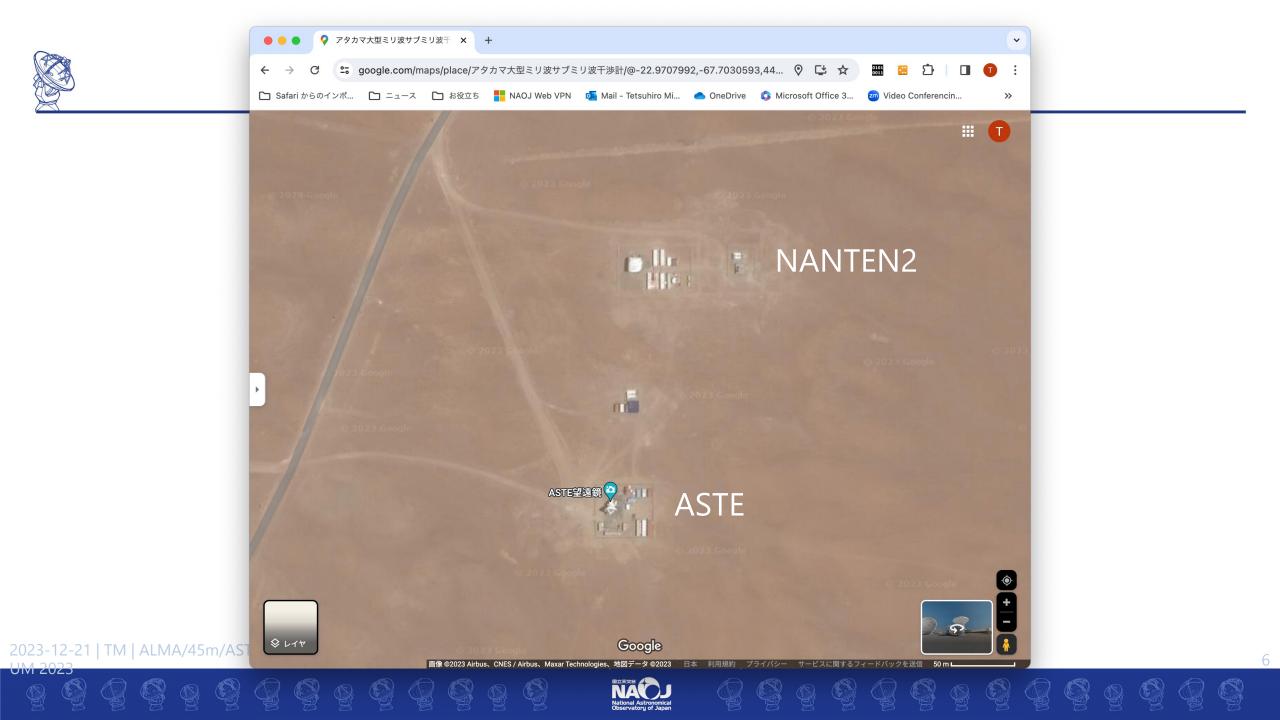
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- Diesel generator (150kW-220V ×2)
- Fuel tank (15kL ×2, consumption 300L/d)
- Satellite network (1Mbps)
- Monitor (weather station, web cameras, etc.)
- <u>https://aste.nao.ac.jp/index\_e.html</u>











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• Struggled with Subreflector Problems

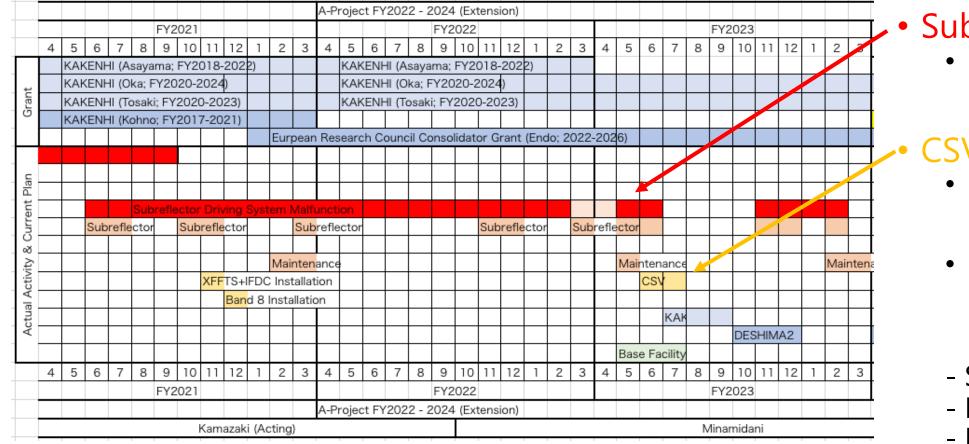
- 2022 Mar.
- 2022 Nov. Dec.
- On-Site Investigation and Recovery Works
  - One of the blocker was solved.
  - Another problem was occurred.
- 2023 Mar.: On-Site Investigation
  - The problem is NOT reproduced.

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#### Sub-Ref Problem

- Recovered by replacing motor for X3 (Chopping axis)
- Wide-IF-bandwidth Band 8 Receiver (CAT8W)
- Digital Spectrometer (XFFTS) and IF Down Converter
- Sub-Ref Parameter
- Pointing
- Linearity
- measurement
- Spurious survey

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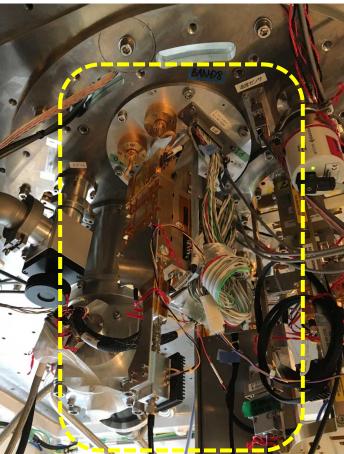
#### Recent Upgrade: Wide-IF-bandwidth Band8 receiver cartridge

- Wide-IF-bandwidth Band8 receiver (CAT8W)
  - Supported by the KAKENHI project (P.I. Oka; Keio U.)
  - The current Band8 receiver cartridge is upgraded with SIS mixers employing high-Jc junctions developed by ATC
  - Same RF range, but IF bandwidth is expanded from 4-8 GHz to 4-18 GHz

e.g.) Simultaneous observations of CO and [CI] in Band 8 become available.

- Trec ~150-250 K,
- IRR ~10-15 dB at Mitaka
- Tsys ~ 600 1000 K (492 GHz)
- Issue
  - one Pol. -> DSB

Receiver	Band8	CAT8W
Beam		1
RF range	387-4	98 GHz
IF range	4-8 GHz	4-18 GHz
Sidebandort & Fu		3, LSB
Polarizations	Х	ζ, Υ



#### Recent Upgrade: Digital Spectrometer and IF Down Converter

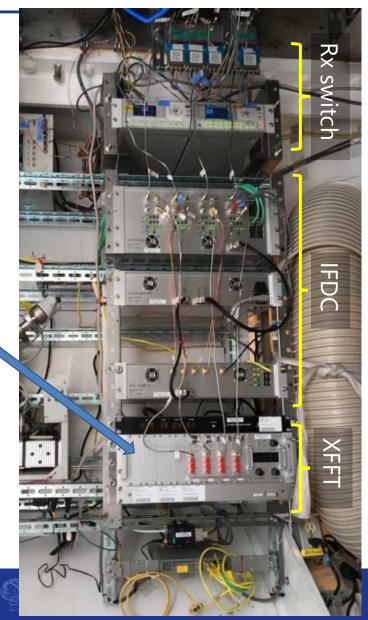
#### RPG eXtended bandwidth <u>FFT Spectrometer</u> (XFFTS)

- Supported by the KAKENHI project (P.I. Tosaki; JUEN)
- 2.5 GHz BW / 32K channels [/Spw]
  - $\Delta v$ =0.047 km/s, velo. width ~1500 km/s@492 GHz
- 10-bit ADC, 5Gsps
  - Good linearity

#### • <u>IF</u> <u>D</u>own <u>C</u>onverter (IFDC)

- 4 spectral windows (2.5 GHz BW) from 4 IF signals of a receiver
- Support of IF BW = 4-8GHz (DASH345, CAT10) and 4-18 GHz (CAT8W)

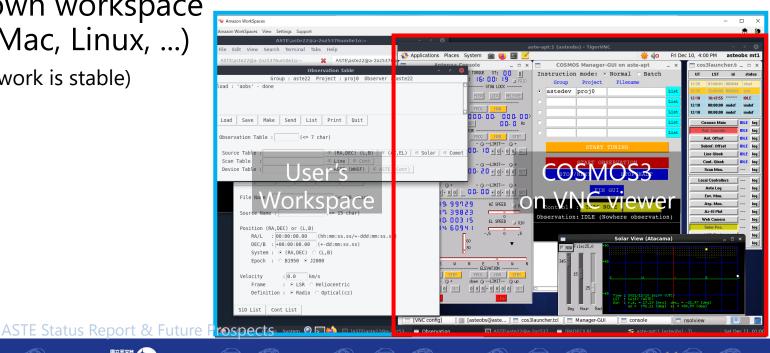
Spectrometer	WHSF -	→ XFFTS
IF bandwidth	2 / 4 GHz	2.5 GHz
# of channels	2,048 *1	32,768
# Of IFStatus Report &	Futur 4 / 2 cts	4
*1 NEWSTAR/NOSTAR	limit	



x4

#### Recent Update: User's Workspace, Remote Control Terminal

- Amazon Workspaces Virtual Desktop Infrastructure
  - A workspace is provided for a user
    - Observation preparation
    - VNC viewer to access the remote-control Virtual PC (on Amazon Web Services)
  - A user can connect its own workspace using AWS client (Win, Mac, Linux, ...) from EA (even from Chile if network is stable)

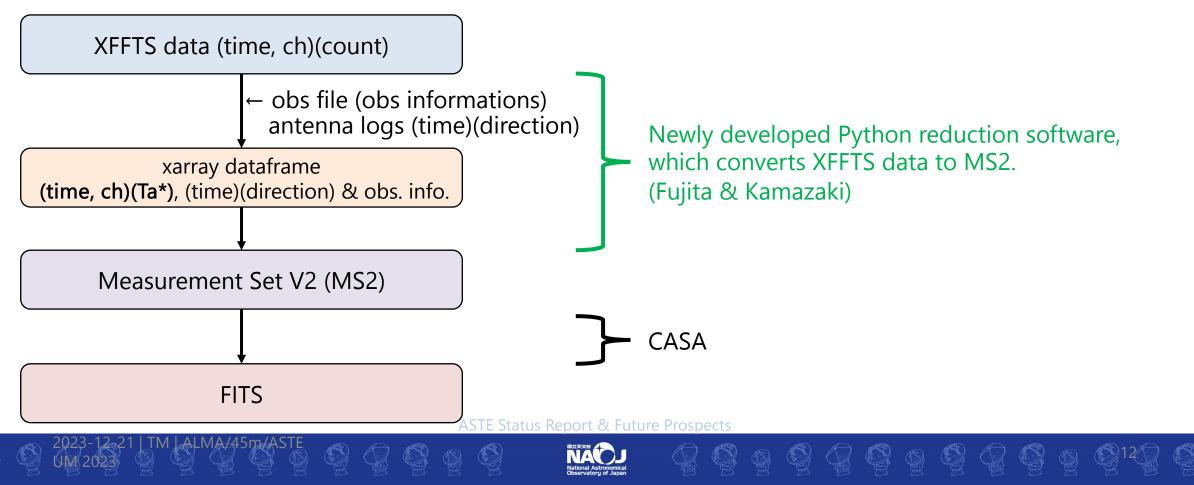




## Recent Update: Data Reduction with CASA

#### Data reduction with CASA

- XFFTS data are reduced with CASA.
  - No plan to update NEWSTAR and NOSTAR to support XFFTS data.



## Recent Update: Data Reduction with CASA

#### Data reduction with CASA

- XFFTS data are reduced with CASA.
  - No plan to update NEWSTAR and NOSTAR to support XFFTS data.
- MSv2 data generator (aka MERGE2) for WHSF and MAC is also under testing.
- These CASA MSv2 format data will be delivered and distributed through the NRO/ASTE Science Data Archive.





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# Science Observations

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### **KAKENHI** Projects

- Tosaki et al.
  - Investigation of physical/chemical properties of ISMs
  - Presentation by Fujita-san.
- Oka et al.
  - Search for Missing Black Holes in the Galaxy based on Submillimeterwave Observations
  - Presentation by Oka-san.

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### Operation/Activities in CY2022 and CY2023

- Relocation of Base Facility in San Pedro de Atacama
  - TAO Base Facility -> Hotel Takha Takha
- NAOJ A-Project Extension for 3 years: FY2022 2024
  - Suspended process due to the COVID-19 and SubRef Malfunction has been resumed and on-going.
  - Completion of process in January is expected.
- Survey of Thesis/Dissertations: PhD, MC, 卒論
  - On-going.

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• Plan to use ryunet / tennet.



# <u>Atacama Submillimeter Telescope Experiment</u>

- 10-m sub-mm telescope at Pampa la Bola
  - Surface accuracy:  $19\mu m (\rightarrow \sim 40\mu m?)$
  - Pointing accuracy: ~2" (rms)
- Heterodyne Receivers
  - Frontends:
    - DASH345 (321-376GHz)
    - Wide-IF-bandwidth Band8 (387-498GHz)
    - Band10 (790-940GHz)
  - Backend: **XFFTS**
- DESHIMA2.0

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<u>https://aste.nao.ac.jp/index\_e.html</u>

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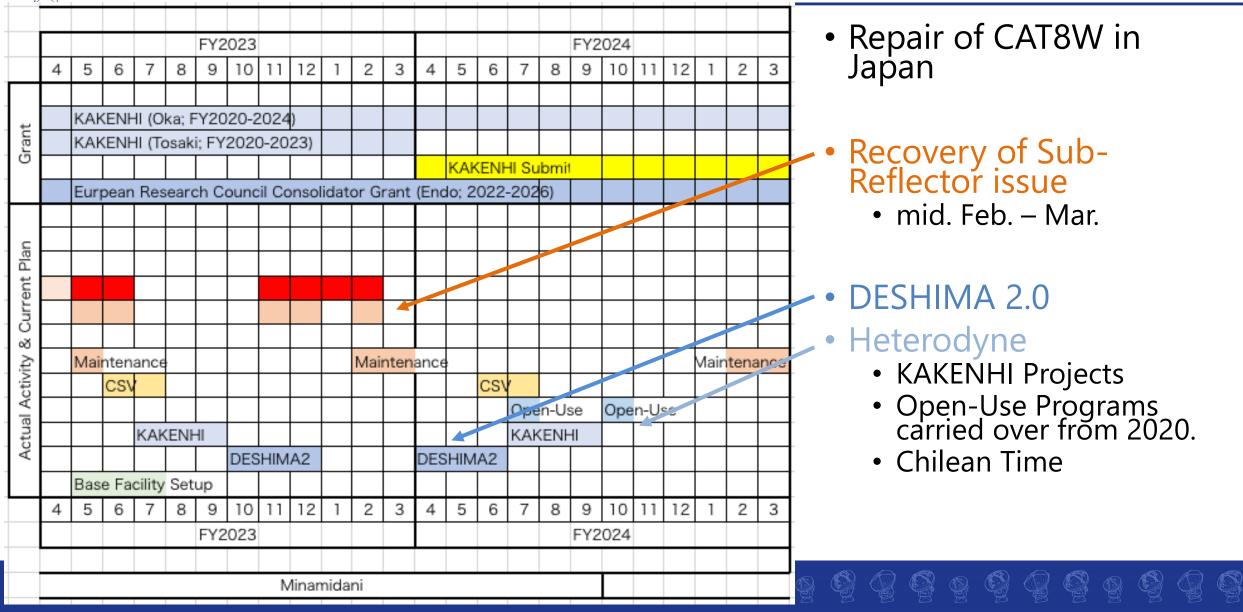




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#### Plan: CY2024 -



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- ASTE Science Workshop
  - planning...

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- International External Review
  - Mar.25 26 @ NAOJ Mitaka
- Improvement of Sub-Reflector Driving System
- A-Project Extension / New A-Project Application for FY2025 27
- NAOJ Science Roadmap, NAOJ Project Implementation Plan for FY2028 -



- The prime objectives of ASTE Project
  - **Promote science** with the submm single-dish telescope
  - **Promote development** of instruments and techniques for astronomical observations

➔ A precursor to ALMA and Submm Astronomy in terms of Science and Instrumentation

- 2022 Jan. 2023 Mar.
  - Struggled with Sub-reflector Problems
  - Updates and Commissioning of New Systems
- 2023 Apr. Dec.
  - 2023 May Sep.: Heterodyne Session (321-376 GHz / 387-498 GHz / 790-940 GHz)
  - 2023 Oct. Dec.: DESHIMA2.0 (220 440GHz)
  - 2023 Nov. Sub-Ref Driving System malfunction
- 2024 Jan. –

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- Recovery of Sub-Ref. Driving System
- 2024 Apr. June: DESHIMA2 (220 440 GHz)
- 2024 June Dec.: Heterodyne Session (321-376 GHz / 387-498 GHz / 790-940 GHz)





- Prediction of Sub-Reflector malfunction?
  - It seems impossible. Need some monitor points, etc. but...
- Surface Accuracy

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- No recent measurements, such as holography.
- Asayama et al. 2022, PASJ, 74, 678
  - ASTE Band 10 (787-950 GHz) heterodyne receiver: System description, commissioning, and science verification
  - https://ui.adsabs.harvard.edu/abs/2022PASJ...74..678A/abstract
  - Aperture efficiency ~ 10 % @ 890 GHz
  - Main beam efficiency ~ 12 % @ 890 GHz: Consistent with Sugimoto et al. 2004 (~ 13 %)
- Approach to the Sub-Reflector Driving System malfunction
  - Short-Term: Replacement of a malfunctioned jack out of 6 jacks to resume observations as soon as possible.
  - Mid-Term (FY2024): Replacement of a whole system.

