

What's inside the delivery package?

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This presentation

- provides an outline about contents of the data package delivered to PI and available via the archive
 - They change from the previous cycles (Cycle 0-4) ones
- provides briefs about QA2 report in SnooPI (new in Cycle 5)
- focuses on Pipeline (PL) processed Interferometer data product



 This presentation makes use of a Cycle 5 delivery package which is by courtesy of Yamashita-san (Ehime Univ.). We sincerely thank his kind understanding.





Package: the archive

ALMA Request Handler

Kouichiro Nakanishi My Requests | Logout

Kouichiro Nakanishi: Request #1169166993 🗹

Request Title: Click to edit

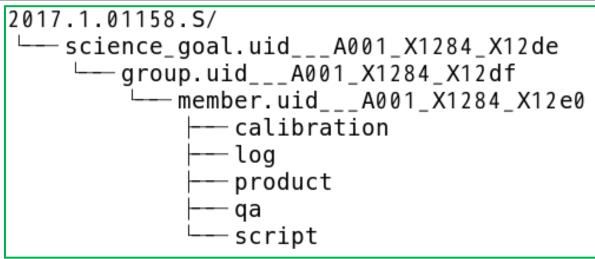
Download Selected

✓ readme ✓ product □ raw □ raw (semipass)

Project / OUSet / Executionblock	File					Size	Accessible
🔻 回 🚞 Request 1169166993							
🔻 回 🚞 Project 2017.1.01158.S							
🗹 📄 readme	2017.1.01158.S.readme						
🔻 📄 🚞 Science Goal OUS uid://A001/X1284/X12de							
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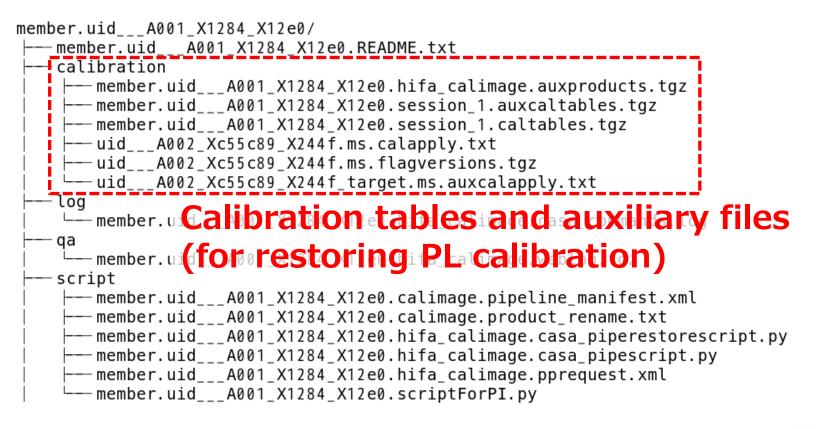
Package: Inside (1)



- calibration: calibration tables
- log: data reduction log files
- product: image products (FITS files)
- qa: PL weblog
- script: script files to restore PL calibration



Package: Inside (2-1)







Package: Inside (2-2)







Package: weblog

لکن ALMA

A Home By Topic By Task

2017.1.01158.S

Observation Overview

Pipeline Summary

Project	uid://A001/X122c/X22f	Pipeline Version	40896 (Pipeline-CASA51-P2-B)
Principal Investigator	takuji		(documentation)
OUS Status Entity id	uid://A001/X1284/X12e0	CASA Version	5.1.1-5 r40000
Observation Start	2017-10-05 05:26:32 UTC	Pipeline Start	2017-11-20 08:07:42 UTC
Observation End	2017-10-05 05:50:02 UTC	Execution Duration	0:45:59

Observation Summary

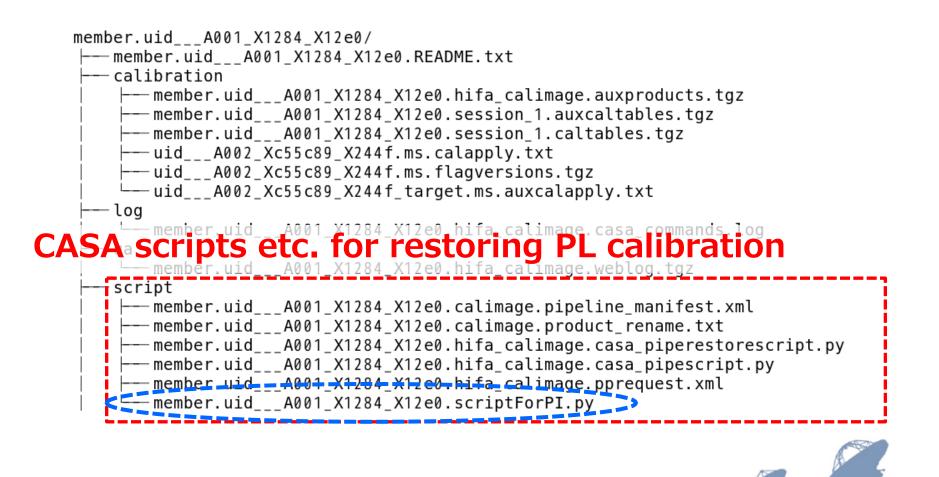
			Time (UTC)			Baseline Length			
Measurement Set	Receivers	Num Antennas	Start	End	On Source	Min	Max	RMS	Size
Observing Unit Set Status: uid://A001/X1284/X12e0 Scheduling Block ID: uid://A001/X1284/X1223									
Session: session_1									
uidA002_Xc55c89_X244f.ms	ALMA Band 6	11	2017-10-05 05:26:32	2017-10-05 05:50:01	0:05:02	8.9 m	48.9 m	27.9 m	732.5 MB
uidA002_Xc55c89_X244f_target.ms	ALMA Band 6	11	2017-10-05 05:44:43	2017-10-05 05:44:43	0:05:02	8.9 m	48.9 m	27.9 m	141.7 MB



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Package: Inside (2-3)





Package: scriptForPI.py

ALMA Data Reduction Script

\$Id: scriptForPI.py,v 1.23 2017/10/17 15:15:15 dpetry Exp \$

Calibration application

```
import os
                       The script for restoring PL calibration
import sys
import glob
from xml.etree import cElementTree as ET
applyonly = True
os.environ["LANG"] = "C"
                                                                                             7
print '*** ALMA scriptForPI ***'
casalog.origin('casa')
casalog.post('*** ALMA scriptForPI ***', 'INFO', 'scriptForPI')
casalog.post('$Id: scriptForPI.py,v 1.23 2017/10/17 15:15:15 dpetry Exp $', 'INFO', 'scriptForPI')
savingslevel=0
if globals().has_key("SPACESAVING"):
   print 'SPACESAVING =', SPACESAVING
   if (type(SPACESAVING)!=int or SPACESAVING<-1):</pre>
       sys.exit('ERROR: SPACESAVING value \"'+str(SPACESAVING)+'\" not permitted, must be int>=-1.\n'
                + 'Valid values: 0 = no saving,\n'
                                1 = delete *.ms.split,\n'
                +
                                2 = delete *.ms and *.ms.split,\n'
                             >=3 = delete *.ms, *.ms.split, and if possible *.ms.split.cal'
                               -1 = do not check disk space')
                +
   savingslevel = SPACESAVING
```



Package: Inside (3-1)

science target

7652901832294517 sci.spw16 18 20 22.cont.I.mask.fits continuum (all the SPWs) 32294517_sci.spw16_18_20_22.cont.I.mask.fits sci.spw16 18 20 22.cont.I.pb.fits.gz e0. 1237652901832294517__sci.spw16.cube.I.mask.fits (Cube) A001_X1284_X12e0._1237652901832294517__sci.spw16.cube.I.pbcor.fits A001_X1284_X12e0._1237652901832294517_sci.spw16.cube.I.pb.fits.gz sci.spw16.mfs.I.mask.fits continuum (per SPV 7 sci.spw16.mfs.I.pbcor.fits sci.spw16.mfs.I.pb.fits.gz sci.spw16.repBW.I.mask.fits line (cube, user request BW sci.spw16.repBW.I.pbcor.fits sci.spw16.repBW.I.pb.fits.az ... (science target cubes and continuum image of other SPWs) member.uid A001 X1284 X12e0.J0006-0623 bp.spw16.mfs.I.mask.fits -member.uid A001 X1284 X12e0.J0006-0623 bp.spw16.mfs.I.pbcor.fits -member.uid A001 X1284 X12e0.J0006-0623 bp.spw16.mfs.I.pb.fits.qz ... (J0006-0623 continuum images of other SPWs) -member.uid___A001_X1284_X12e0.J0141-0928_ph.spw16.mfs.I.mask.fits - member.uid___A001_X1284_X12e0.J0141-0928_ph.spw16.mfs.I.pbcor.fits -member.uid___A001_X1284_X12e0.J0141-0928_ph.spw16.mfs.I.pb.fits.qz ... (J0141-0928 continuum images of other SPWs)



11



Package: Inside (3-2)

science target

product
Continuum A061 X1284 X12e0 327652901832294517_sci.spw16_18_20_22.cont.I.mask.fits member.uid A061 X1284 X12e0 327652901832294517_sci.spw16_18_20_22.cont.I.pbcor.fits member.uid A001 X1284 X12e0. 1237652901832294517_sci.spw16_18_20_22.cont.I.pb.fits.gz
COHUMA @ AM & LA C S P VV S 32294517_sci.spw16_18_20_22.cont.I.pbcor.fits
member.uidA001_X1284_X12e01237652901832294517sci.spw16_18_20_22.cont.I.pb.fits.gz
<pre>momber.vidA001_X1284_X12e01237652901832294517sci.spw16.cube.l.mask.fits</pre>
ine (Cube) A001_X1284_X12e01237652901832294517sci.spw16.cube.I.mask.fits
member.uidA001_X1284_X12e0. 1237652901832294517sci.spw16.cube.I.pb.fits.gz
member.uid A021 X1284 X1264 1177652901832294517sci.spw16.mfs.1.mask.fits
CONTINUUMA (per SPW2)7652901832294517_sci.spw16.mfs.I.pbcor.fits
member.uidA001_X1284_X12e01257652901832294517sci.spw16.mfs.I.pb.fits.gz
ine Cube, Auser request BW) 17_sci.spw16.repBW.I.mask.fits
IIIE CUDE, USEI REQUEST DV 17_sci.spw16.repBw.1.pbcor.fits
member.uidA001_X1284_X12e01237652901832294517sci.spw16.repBW.I.pb.fits.gz
├── (science target cubes and continuum image of other SPWs)
Bandpass call (continuum) p. spw16.mfs. I. mask. fits
member.uidA001_X1284_X12e0.J0006-0623_bp.spw16.mfs.I.pb.fits.gz
├── (J0006-0623 continuum images of other SPWs)
member uid $= 4001 \times 1284 \times 1260 \oplus 10141 - 0028$ pb spw16 mfs T mask fits
Phase Cald (Continuum) -0928_ph.spw16.mfs.I.mask.fits
member.uidA001_X1284_X12e0.J0141-0928_ph.spw16.mfs.I.pb.fits.gz
├── (J0141-0928 continuum images of other SPWs)





Weblog: calibrator images

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By Topic By Task

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16. hifa_timegaincal

17. hif_applycal

18. hifa_imageprecheck

19. hif_makeimlist

20. hif_makeimages

21. hif_checkproductsize

22. hifa_exportdata

23. hif_mstransform

24. hifa_flagtargets

25. hif_makeimlist

26. hif_findcont

27. hif_uvcontfit

28. hif_uvcontsub

29. hif_makeimages

30. hif_makeimlist

31. hif_makeimages

32. hif_makeimlist

33. hif_makeimages

34. hif_makeimlist

35. hif_makeimages

20. Tclean/Makelmages

Make calibrator images

Image Details

Field	Spw	Pol	Image details		Image result
J0006-0623 (BANDPASS)	16	I.	centre frequency of image	221.5480GHz (LSRK)	Type:mage digiterrown field/0006-0623 spw:16 ter:1
			beam	7.69 x 4.22 arcsec	
			beam p.a.	-77.6deg	Conception and the second seco
			final theoretical sensitivity	0.00051 Jy/beam	400 400 400
			cleaning threshold	0.019 Jy/beam Dirty DR: 3.8e+03 DR correction: 19	View other QA images
			clean residual peak / scaled MAD	12.94	
			non-pbcor image RMS	0.0020 Jy/beam	
			pbcor image max / min	1.94 / -0.00934 Jy/beam	

Atacama Large Millimeter/submillimeter Array

33 54

2017.1.01158.S

BACK

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 provides summary of the data (beam size, noise level) and comments by data reducer

– QA2 report replaces Cycle 0-4 README

- at the moment (Dec. 2017), one can access QA2 report solely via SnooPI
 - QA2 report will be available as a part of data package via the archive in the near future

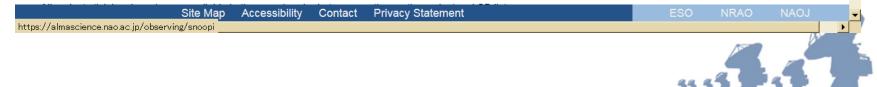




SnooPI: Science Portal

https://a	almascience	.nao.ac.jp/ob	serving/snoopi	Log in
About Science Proposing	Observing Data Process Phase 2 ToO activation	sing Tools Documentation Help	Search Site	٩
The Snooping Project Interface (SnooPI) is projects. SnooPI is available through the lin application and here.	^s ALMA Status Page	-	vational and processing status of their ALMA s ink to the User Manual is available from within	
Bug fixes and new features of the June	2017 release:			
 SB and MOUS history added. Time and visit constraints added at the SI Older updates: 	B details level.			
May 2017:				

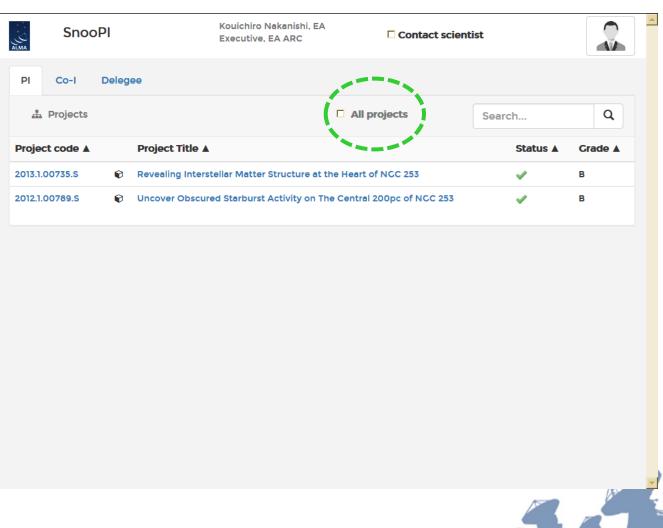
- Loading projects, SBs and detailed EB view is now faster.





SnooPI: PI projects

SnooP	I	÷	
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Home		*	
My Proje	ects	ሑ	
My SBs		Ø	
QUICK LIN	IKS		
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Science	Portal	ß	
Archive	Query		
Helpdes	k	٥	
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SnooPI: project (tree view)

2	♣ 2017.1.01158.S		Executions	<u> </u>	
	ACA Study on the Driving Mechanisms of Starburst and Main-Sequence Star Formation in Local Calaxies				
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	& Member OUS (1237652630714122400)	- a (-	-Status		
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SnooPI: Project overview

	2015 Tel 156.5 ACA Study on the Driving	A	Project	2017.1.01158.S
Med Seq	chanisms of Starburst and Main- uence Star Formation in Local Calaxies		Title	ACA Study on the Formation in Loc
8	Cost UnitSet SC OUS (disk_12376526307141224		Ы	Takuji Yamashita
	 Croup OUS Member OUS 765263071412240 	#	Co-I	Takuma Izumi, Hi Kouichiro Nakani
	12376526_a_06_7M SC OUS (disk_12376517550961459		Туре	Standard
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Ы	Takuji Yamashita	
)-I	Takuma Izumi, Hiroyuki Kaneko, Yusei Koyama, Shuhei Koyama, Tohru Nagao, Kouichiro Nakanishi, Tsutomu Takeuchi, Yoshiki Toba, Junko Ueda, Jing Wang	
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QA2 report: Instructions

Instructions

INSTRUCTIONS OVERVIEW

Introduction Total Power Data Pipeline-calibrated TP data Manually calibrated TP data Primary Beam Correction How to restore the calibrated MeasurementSet for your data

######

INTRODUCTION

This file describes the content of the tar file you have received. The full data structure is inserted below.

At this stage, we are releasing data after completion of one SB (excuted multiple times if required), so you will find only one member_ouss_id directory. This directory contains this README file and the following directories: calibration, script, qa2, log, product.

- 'calibration' contains the files needed for calibration starting from the initial ms to the fully calibrated data.

- 'script' contains the reduction scripts used to process the initial ms to calibrated data, but also to obtain concatenated data (if more than one execution) and imaging products. There are usually several scripts dealing with different parts of the processing.

In case the calibration was done by the automated pipeline, you will also see the Pipeline Processing Request File (PPR).

The most important script for you is the "scriptForPl.py". See the section "How to restore the calibrated MeasurementSet (MS) for your data" further below. (Note: the ALMA archive may prepend the script name with the MOUS UID.)

- 'product' contains the fits files of the selected image products.



Further information

- ALMA Science (CASA) Pipeline
 <u>https://almascience.nao.ac.jp/processing/s</u>
 <u>cience-pipeline</u>
- SnooPI

https://almascience.nao.ac.jp/observing/sn
oopi

QA2 data products (obsolete)
 <u>https://almascience.nao.ac.jp/processing/q</u>
 <u>a2-data-products</u>

